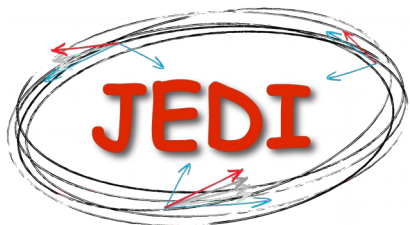




PROGRESS TOWARD A DIRECT MEASUREMENT OF THE DEUTERON ELECTRIC DIPOLE MOMENT AT COSY

HERAEUS SEMINAR 30.03.2021 | VERA SHMAKOVA FOR JEDI COLLABORATION



CHARGED PARTICLE EDM



- No direct measurement for charged hadron EDMs
- Potentially higher sensitivity (compared to neutrons):
 - longer lifetime
 - more stored polarized protons/deutrons
 - can apply larger electric fields in storage rings
- EDM of single particle type not sufficient to identify CPV source

THOMAS - BMT EQUATION



$$\frac{d\vec{S}}{dt} = \vec{\Omega} \times \vec{S} = -\frac{q}{m} \left\{ \overset{\text{MDM}}{G_y \vec{B}} + \left(\frac{1}{\gamma^2 - 1} - G \right) \frac{\vec{\beta} \times \vec{E}}{c} + \overset{\text{EDM}}{d \frac{m}{q \hbar S} (\vec{E} + c \vec{\beta} \times \vec{B})} \right\} \times \vec{S}$$

At storage rings: vertical \mathbf{B} field, radial \mathbf{E} field

MDM causes fast spin precession in horizontal plane

EDM causes slow spin rotation out of horizontal plane, up and down

In an all-electric storage ring, with the frozen spin condition, a radial electric field causes the spin to precess out of the storage plane linearly

THOMAS - BMT EQUATION



$$\frac{d\vec{S}}{dt} = \vec{\Omega} \times \vec{S} = -\frac{q}{m} \left\{ \overset{\text{MDM}}{G_y \vec{B}} + \left(\frac{1}{\gamma^2 - 1} - G \right) \frac{\vec{\beta} \times \vec{E}}{c} + \overset{\text{EDM}}{d \frac{m}{q \hbar S} (\vec{E} + c \vec{\beta} \times \vec{B})} \right\} \times \vec{S}$$

At storage rings: vertical \mathbf{B} field, radial \mathbf{E} field

MDM causes fast spin precession in horizontal plane

EDM causes slow spin rotation out of horizontal plane, up and down

In **pure magnetic ring** motional electric field $(c \vec{\beta} \times \vec{B})$



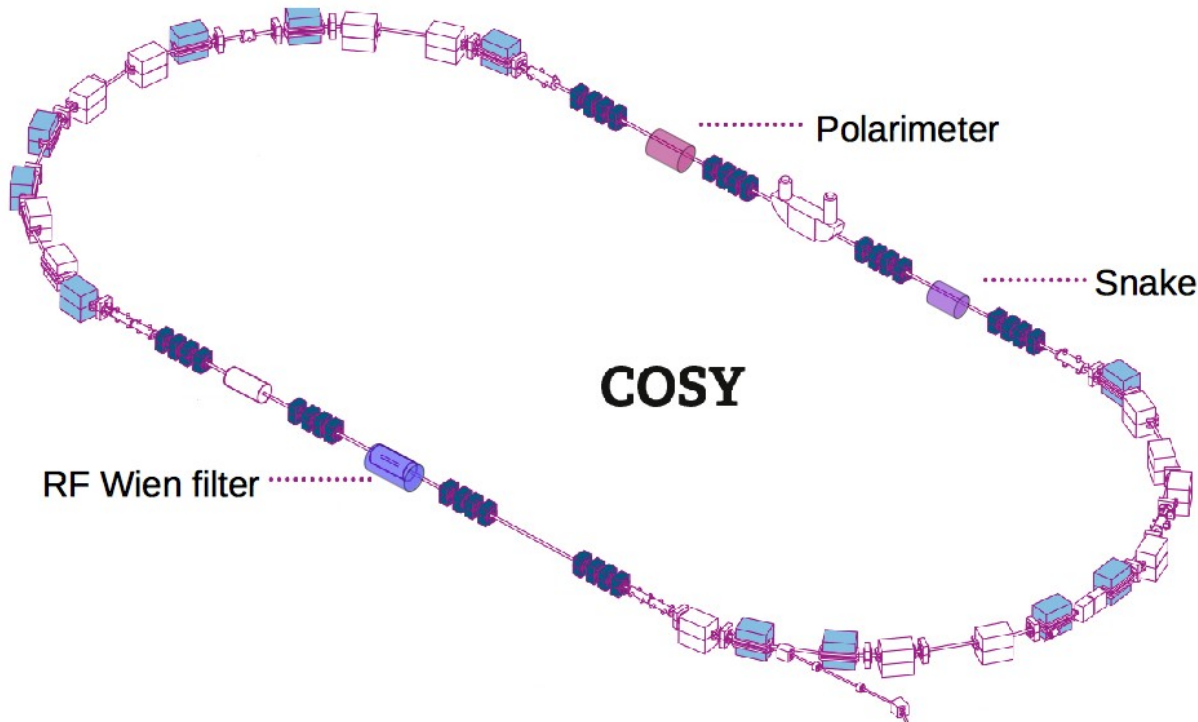
access to EDM

EDM AT COSY



COSY (Jülich, Germany)

- magnetic storage ring
- polarized protons and deuterons
- Momenta $p = 0.3 - 3.7 \text{ GeV}/c$

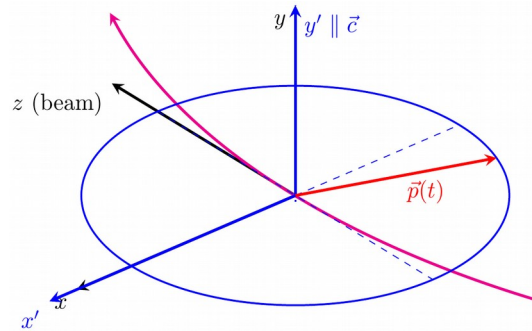


Starting point for
EDM measurement

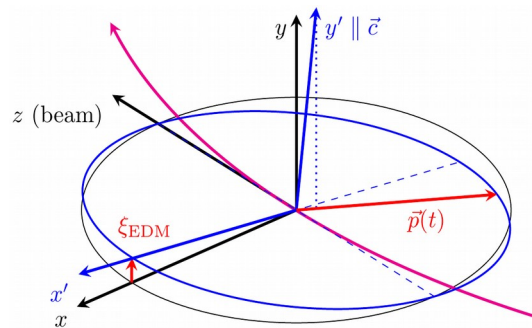
EFFECT ON PRECESSION AXIS



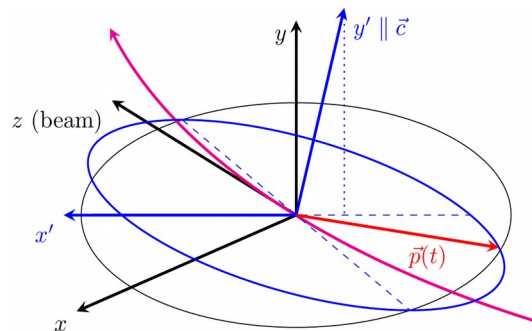
EDM absence case



EDM effect



Magnetic misalignment effect



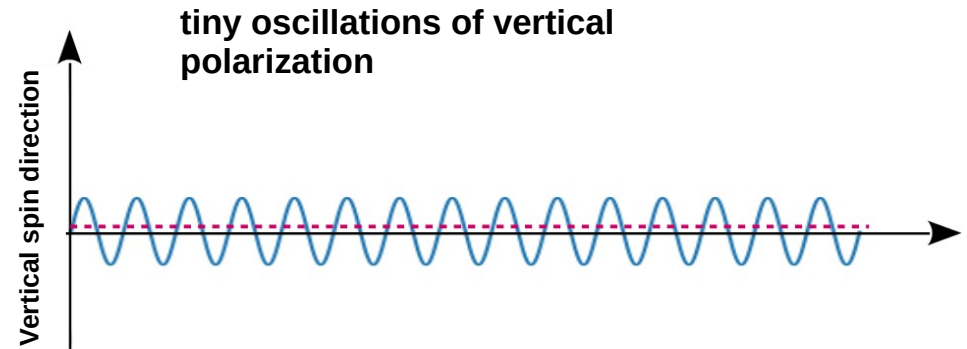
RF WIEN FILTER



In the magnetic ring

momentum $\uparrow\uparrow$ spin \rightarrow spin kicked up
momentum $\uparrow\downarrow$ spin \rightarrow spin kicked down

\downarrow
no accumulation of vertical asymmetry

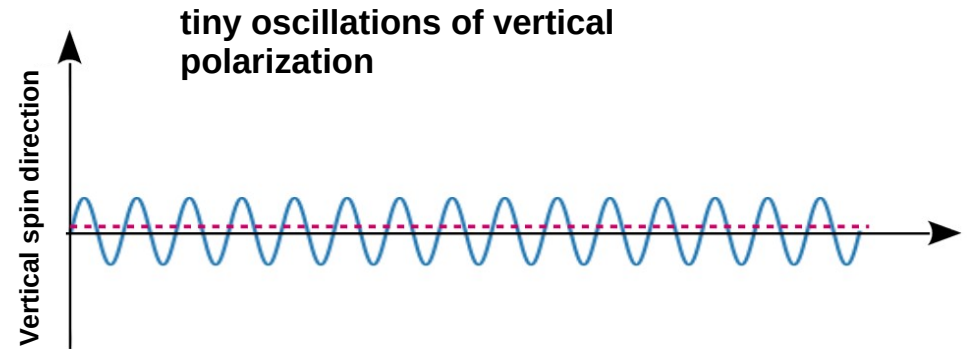


RF WIEN FILTER



In the magnetic ring
 momentum $\uparrow\uparrow$ spin \rightarrow spin kicked up
 momentum $\uparrow\downarrow$ spin \rightarrow spin kicked down

no accumulation of vertical asymmetry

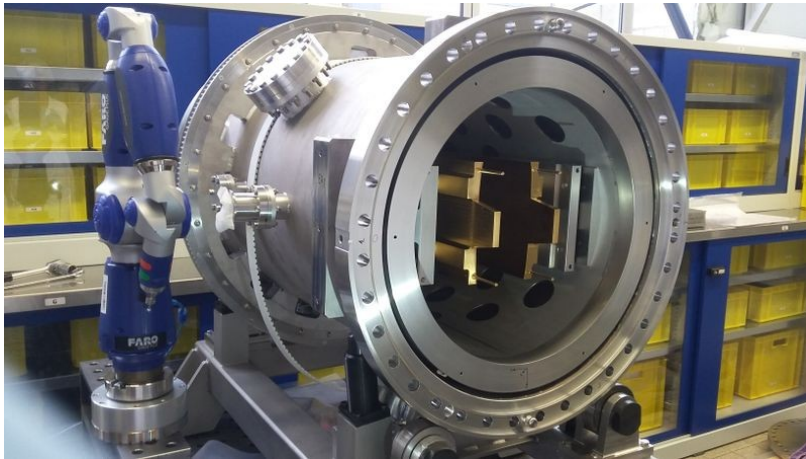


RF Wien filter

Heberling, Hölscher and J. Slim

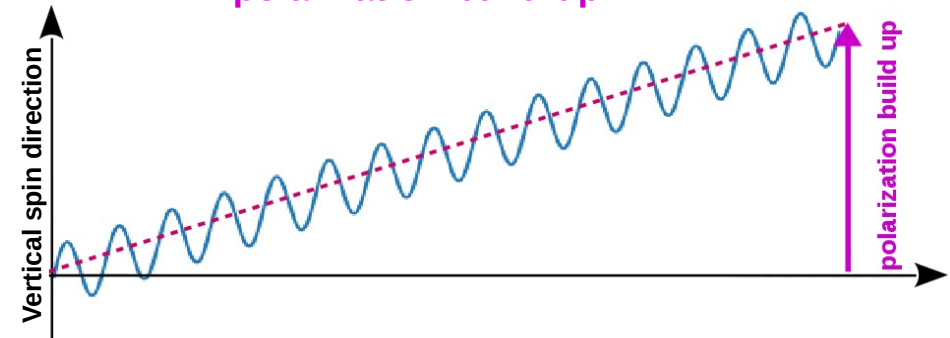
J. Slim et al. Nucl. Instrum. Methods Phys. Res. A 828, 116 (2016)

- Lorentz force $\vec{F}_L = q(\vec{E} + \vec{v} \times \vec{B}) = 0$
- $\vec{B} = (0, B_y, 0)$ and $\vec{E} = (E_x, 0, 0)$

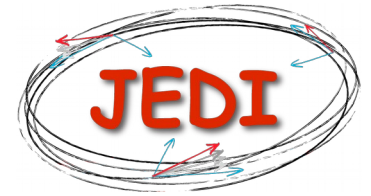


phase lock between spin precession and RF Wien filter

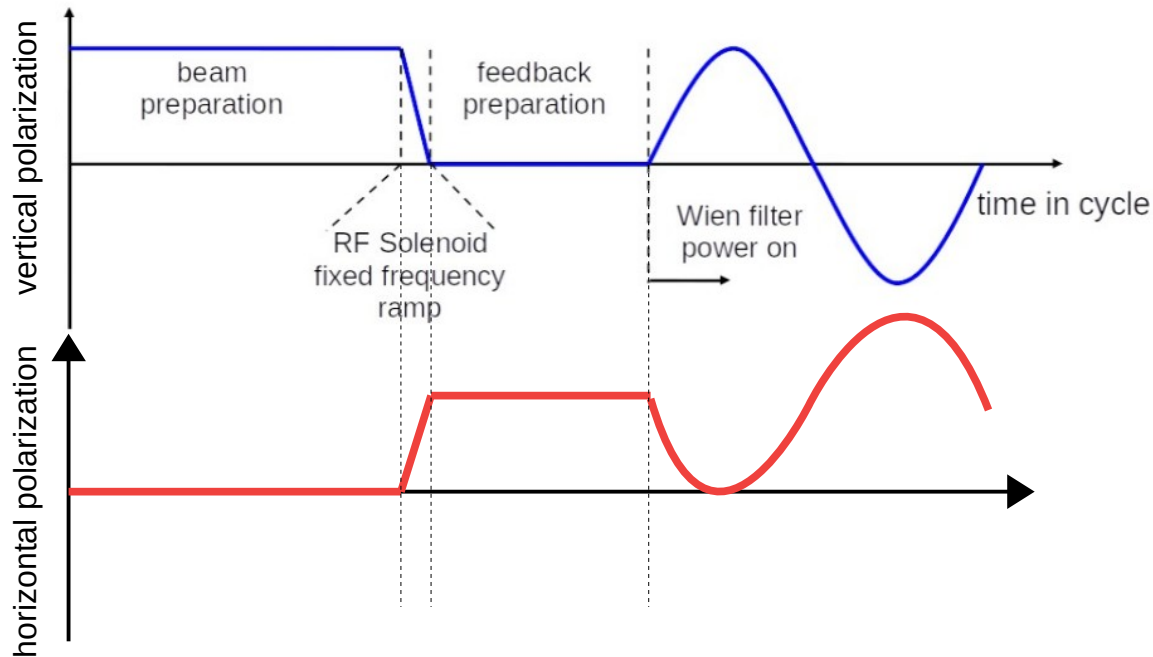
polarization build-up



PRINCIPLE OF MEASUREMENTS



The basic workflow



- Coherent ensembles in ring plane → time of the horizontal polarization decoherence - “spin coherence time” - has to be longer than a measurement

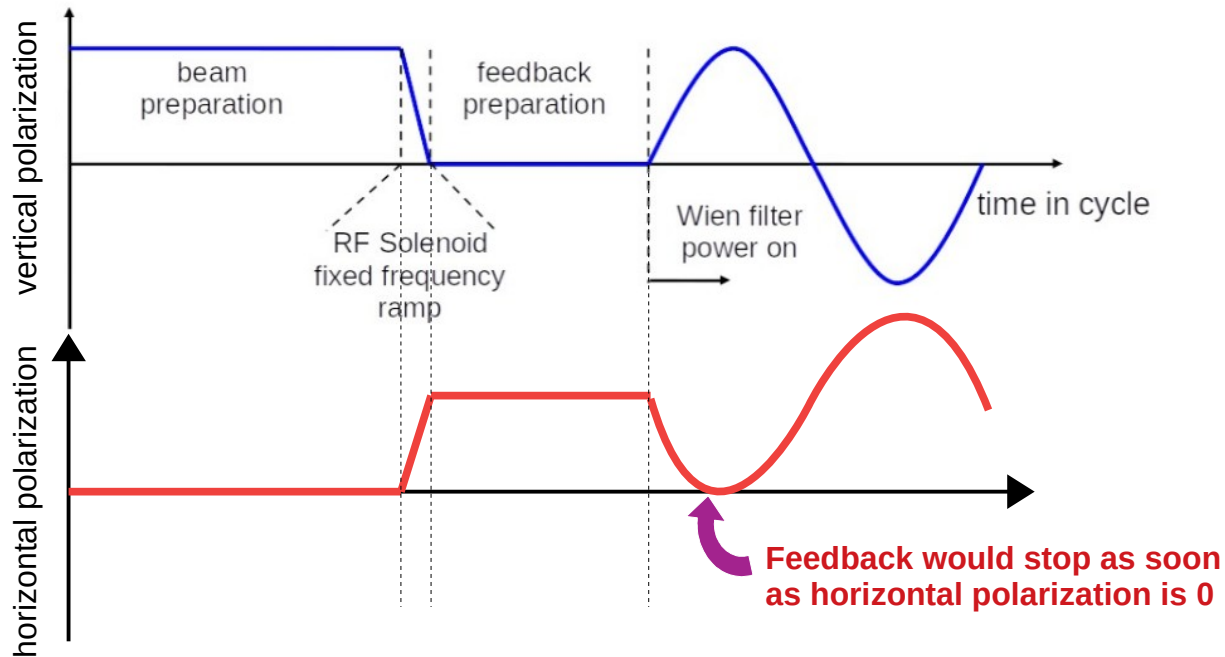
A. Wronska talk

- Spin precesses with 120 kHz.
- Wien filter operates on resonance $f = 871.430$ kHz
- Phase lock between spin precession and Wien filter

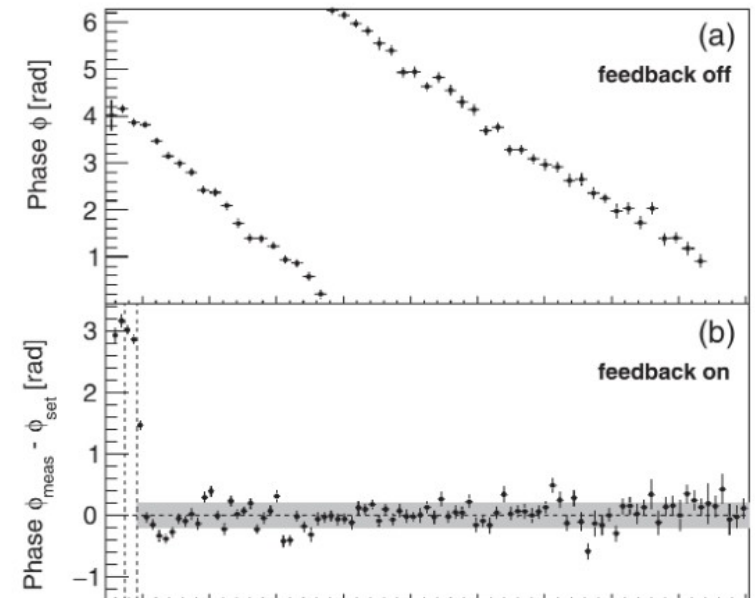
PRINCIPLE OF MEASUREMENTS



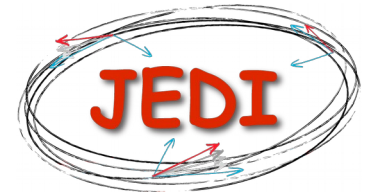
The basic workflow



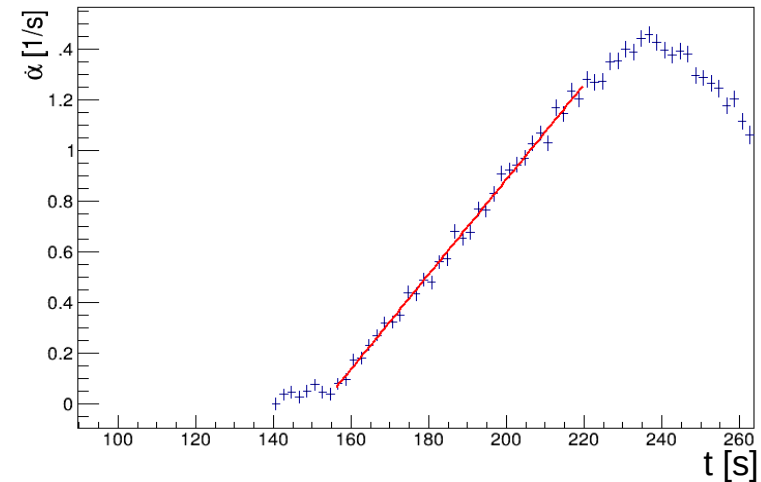
- Feedback monitors spin tune and adjust WF frequency to maintain the relative phase between spin precession and Wien filter
- an error of 0.2 rad



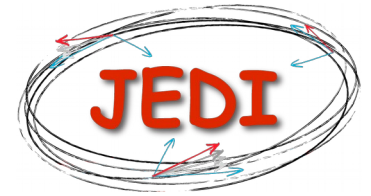
POLARIZATION BUILD-UP



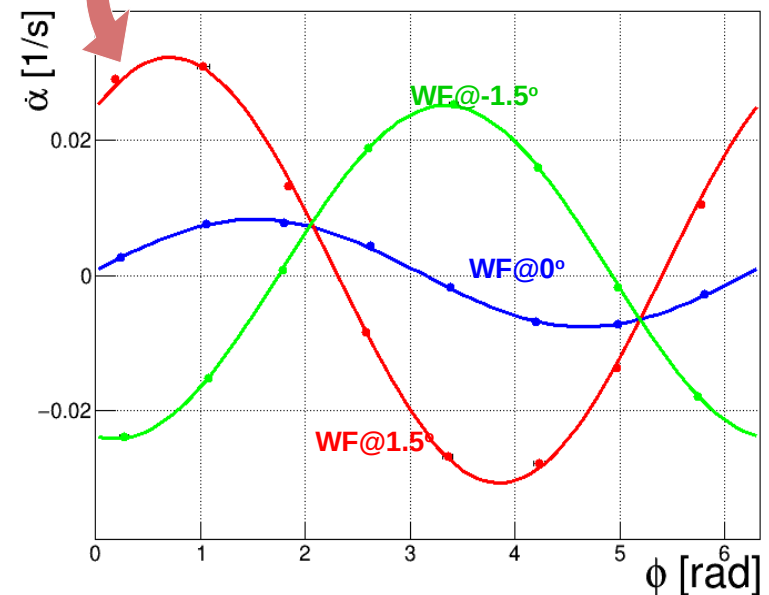
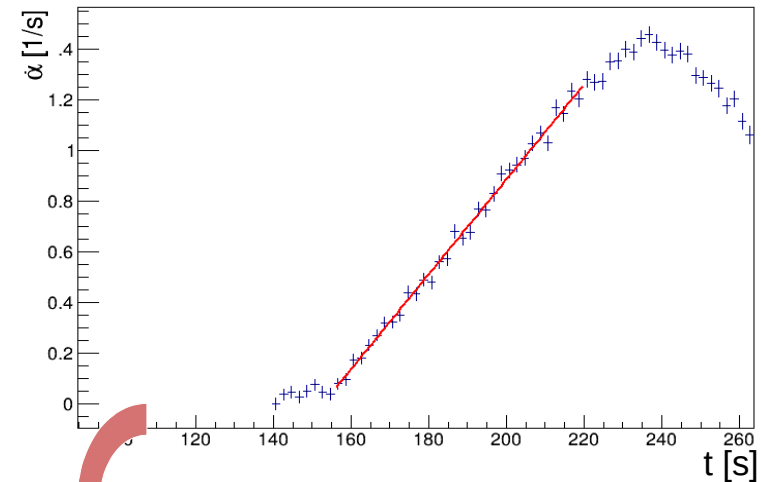
- Wien filter operated with B field normal to the ring plane
- $\alpha(t) = \arctan\left(\frac{P_y}{P_{xz}}\right)$
- Observed initial slopes of polarization build-up varied of Wien filter and solenoid rotations
- Observed slopes would correspond to the EDM of $\sim 10^{-17}$ e·cm



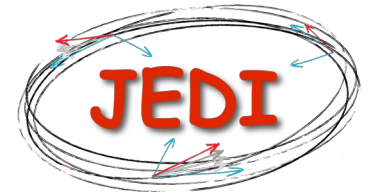
POLARIZATION BUILD-UP



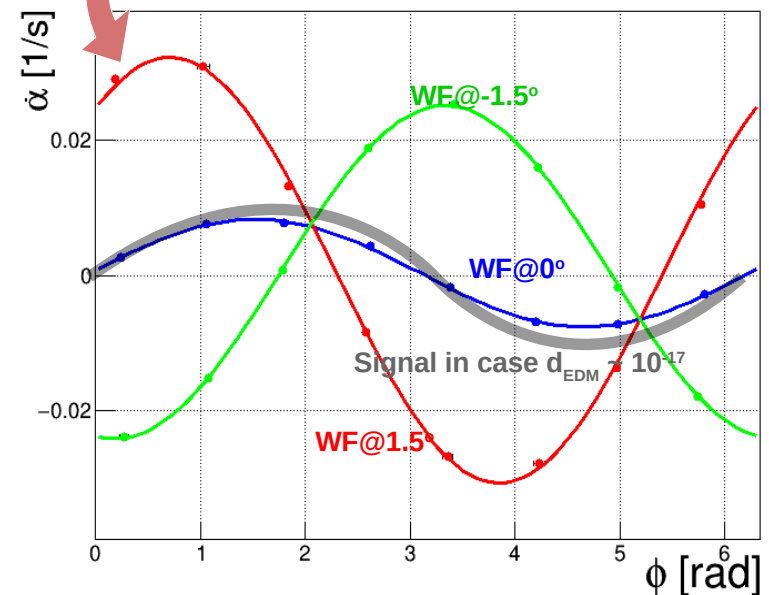
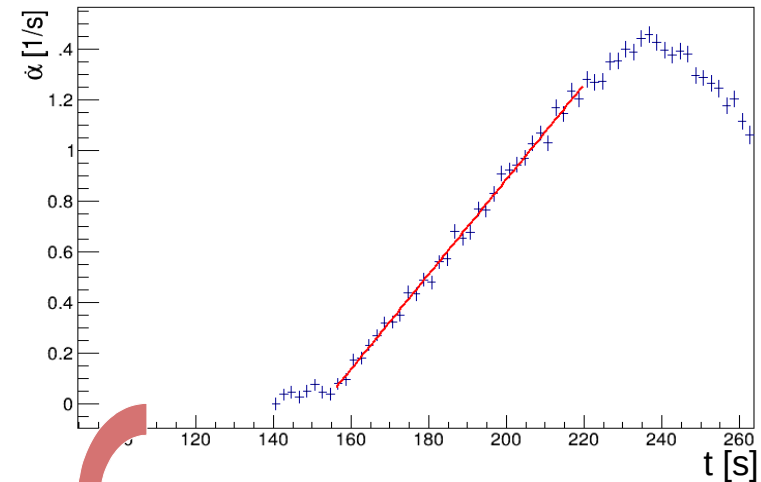
- Wien filter operated with B field normal to the ring plane
- $\alpha(t) = \arctan\left(\frac{P_y}{P_{xz}}\right)$
- Observed initial slopes of polarization build-up varied of Wien filter and solenoid rotations
- Observed slopes would correspond to the EDM of $\sim 10^{-17}$ e·cm



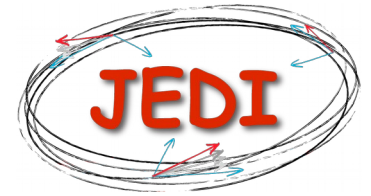
POLARIZATION BUILD-UP



- Wien filter operated with B field normal to the ring plane
- $\alpha(t) = \arctan\left(\frac{P_y}{P_{xz}}\right)$
- Observed initial slopes of polarization build-up varied of Wien filter and solenoid rotations
- Observed slopes would correspond to the EDM of $\sim 10^{-17}$ e·cm



RESONANCE STRENGTH



9+9+14 points on 3 map

Parametric resonance strength based on initial slope $\varepsilon \simeq \frac{\dot{\alpha}}{\omega_{rev}}$

$$\varepsilon = \frac{\psi_{WF}}{4\pi} \sqrt{A_{WF}^2 (\phi^{WF} - \phi_0^{WF})^2 + A_{sol}^2 \left(\frac{\chi_0^{sol} + \chi^{sol}}{2 \sin(\pi \nu_s)} \right)^2} + \varepsilon_0$$

$$\phi_0^{wf} = -3.80 \pm 0.05 \text{ mrad}$$

$$\chi_0^{sol} = -5.51 \pm 0.05 \text{ mrad}$$

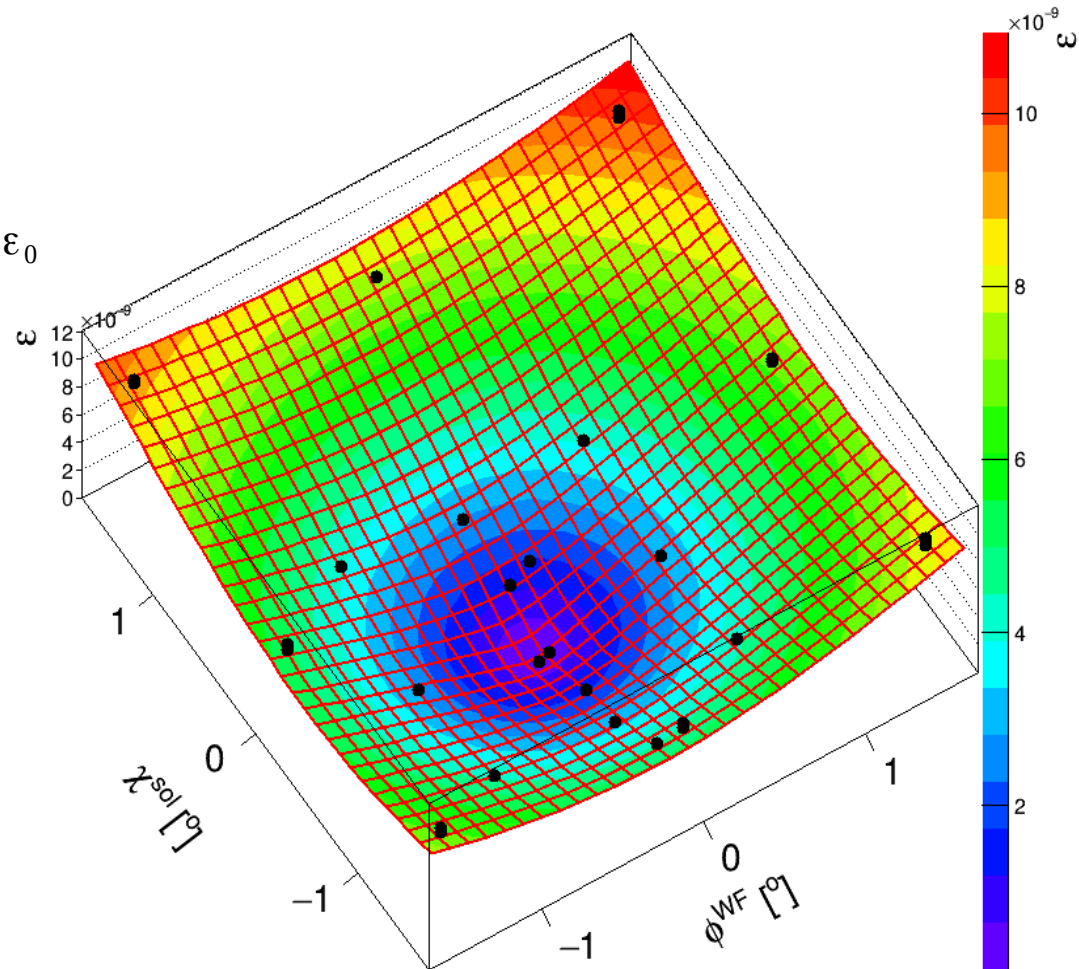
$$A_{WF} = 0.755 \pm 0.004$$

$$A_{sol} = 0.919 \pm 0.004$$

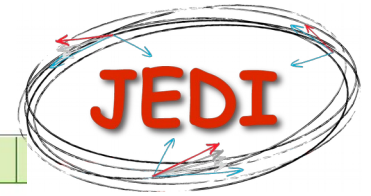
$$\varepsilon_0 = (-1.1 \pm 0.1) \times 10^{-10}$$

Orientation of precession axis at location of RF Wien filter determined from the minimum of the surface

Spin tracking calculations should provide the orientation of precession axis without EDM

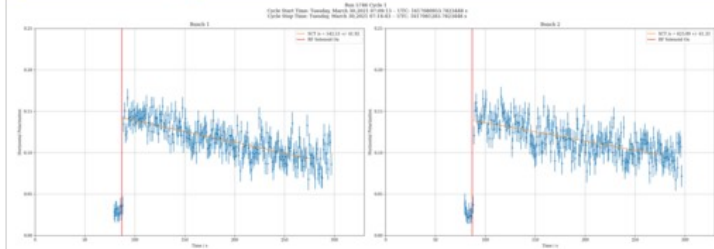


PRECURSOR II IS ONGOING



949	Tuesday, March 30, 2021, 07:09	run	5746	5746	Production	SCT measurement (WF-off Snake-off) for map 3 point 3 WF angle @ -4 deg, Snake 0 A, Phase 5.4 rad	OJ, AK
-----	--------------------------------	-----	------	------	------------	--------------------------------------------------------------------------------------------------	--------

Attachment 1: sct.png



946	Tuesday, March 30, 2021, 07:00	shift summary				all measurements for map3, point1-3 has been done.	AK,OJ
-----	--------------------------------	---------------	--	--	--	----------------------------------------------------	-------

All these measurements has been recorded during day and night shift

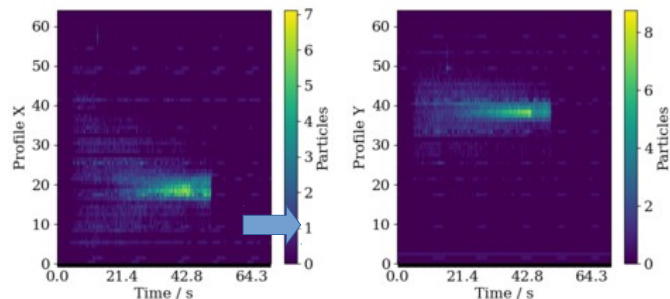
- point 1: WF @-4, Snake at 0 A, phases (0, 0.9, 1.8, 2.7, 3.6, 4.5, 5.4),
- IPM measurement, WF on, Snake on,
- SCT measurement, WF off, Snake on,
- point 2: WF @-4, Snake at 4 A, phases (0, 0.9, 1.8, 2.7, 3.6, 4.5, 5.4),
- IPM measurement, WF on, Snake on,
- SCT measurement, WF off, Snake on,
- SCT measurement, WF off, Snake off,
- point 3: WF @-4, Snake at -4 A, phases (0, 0.9, 1.8, 2.7, 3.6, 4.5, 5.4),
- IPM measurement, WF on, Snake on,
- SCT measurement, WF off, Snake on,
- SCT measurement, WF off, Snake off,

Now Alex can come and rotate Wien filter to 4 degree (+70.36 mrad) and continue measurements untill 8:30 (2 MeV cooler test starts)

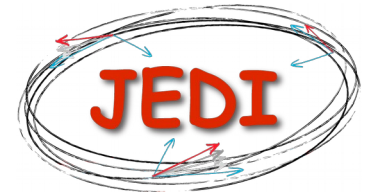
959	Tuesday, March 30, 2021, 06:41	run	5744	5744	Production	IPM measurement for map point 3 WF ON angle @ 0 deg, Snake 0 A,	KL, AN
-----	--------------------------------	-----	------	------	------------	-----------------------------------------------------------------	--------

Attachment 1: beamCooling.png

Beam Cooling Run 5750 Cycle 1
 Cycle Start Time: Tuesday, March 30, 2021 13:58:45 -- UTC: 1617105525.880117
 Cycle Stop Time: Tuesday, March 30, 2021 14:04:15 -- UTC: 1617105855.880117



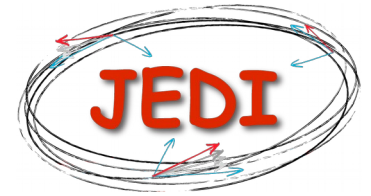
PRECURSOR II IS ONGOING



IMPROVEMENTS AT COSY

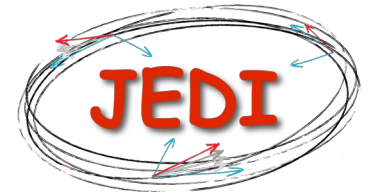
- Alignment campaigns of COSY magnet system
- Beam-based alignment *See T. Warner talk*
PhD thesis T. Wagner T. Wagner et al. JINST 16 T02001 (2021)
- New tool for fast tune and chromaticity measurement
P. Niedermayer and B. Breitkeutz
- Slow control system
I. Bekman and IKP4
- COSY signals and distribution was improved
K. Laihem and V. Hejny
- Rogowski coils at the Wien filter place
F. Abusaif See F. Abusaif talk
- New JEDI polarimeter
I. Keshelashvili and the polarimeter group
- 8 fast switchers to gate the WF power for 1 bunch of 2 of 4
 ➔ pilot bunch technique *See J. Slim talk*
J. Slim, A. Nass, F. Rathmann, G. Tagliente

SUMMARY



- First measurement of deuteron EDM was performed
- Experiment performed is a proof of principle of EDM measurement at storage rings
- Precursor II is ongoing right now
- COSY remains a unique facility for such studies

PILOT BUNCH

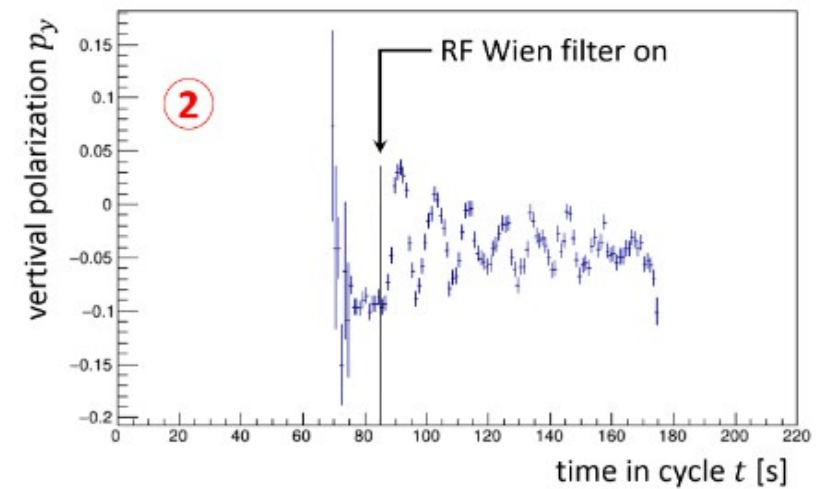
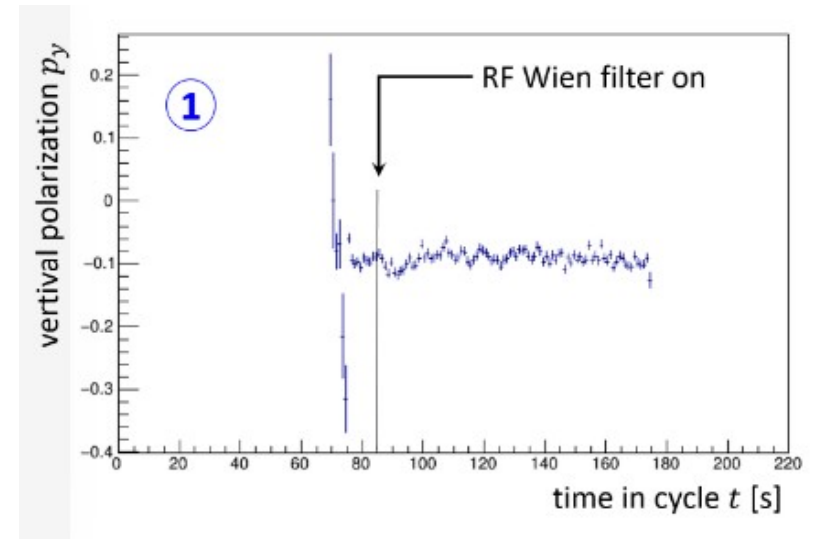
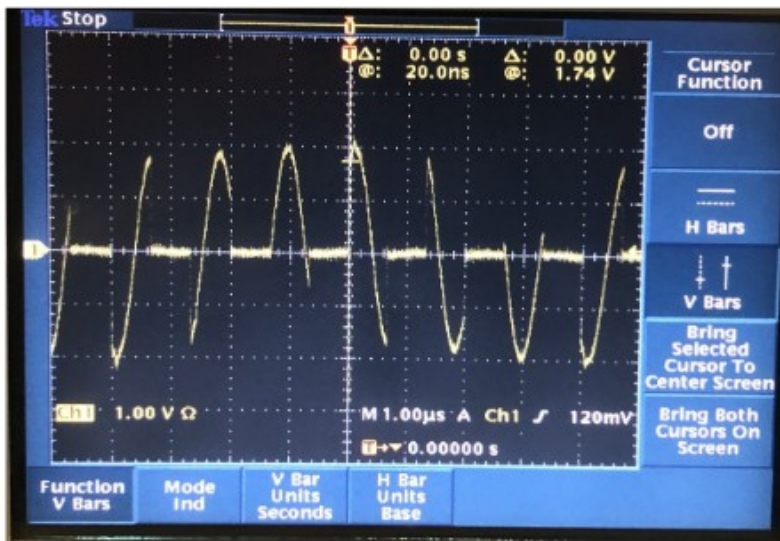


8 fast switchers to gate the WF power for 1 bunch of 2 of 4

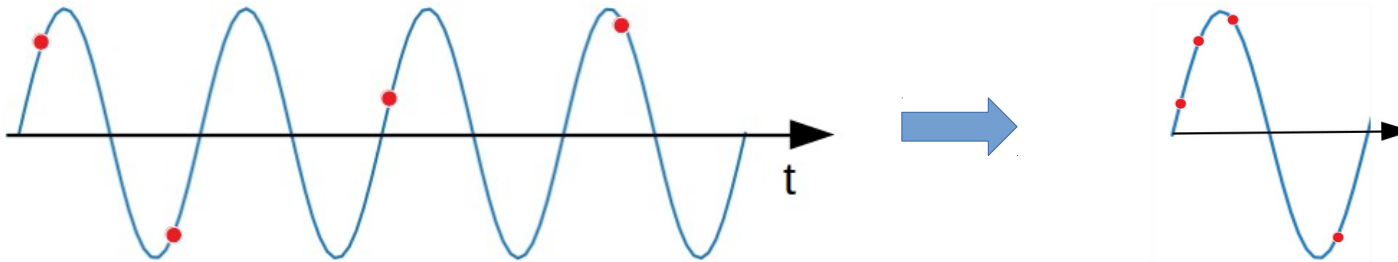
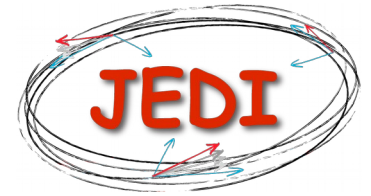
capable of short switch time ~ few ns

One bunch feels the power and oscillate

Another is a pilot bunch



SPINTUNE AND HORIZONTAL POLARIZATION



- Spin precesses with 120 kHz.
- With event rates of $\sim 15000 \text{ s}^{-1}$, there is 1 hit per 10 precessions.
- Not possible to resolve horizontal oscillation directly
- Spintune is determined in each time bin with monitoring phase of measured horizontal asymmetry with fixed spin tune:

$$\nu_s(n) = \nu_s^{fix} + \frac{1}{2\pi} \frac{d\phi}{dn} = \nu_s^{fix} + \Delta \nu_s(n)$$

