

A Search for Axion-like Particles with a Horizontally Polarized Beam in a Storage Ring

Edward J. Stephenson, for JEDI Collaboration
Indiana University

2019 Workshop on Polarized Sources, Targets,
and Polarimetry, September 23-27, 2019



Axions proposed to explain the lack of CP violation. Electric dipole moments (EDMs) did not appear.

Axions are a candidate for interstellar dark matter.

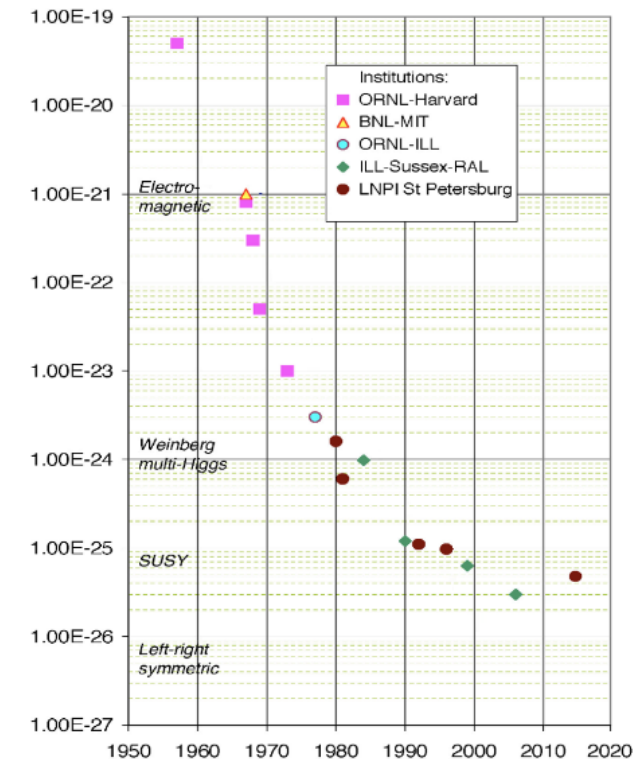
One characteristic is the ability to induce oscillating EDM.

$$d = d_0 + d_1 \cos(\omega t + \varphi)$$

Oscillation frequency
connected to axion mass.

At the time of observation,
the phase is unknown.

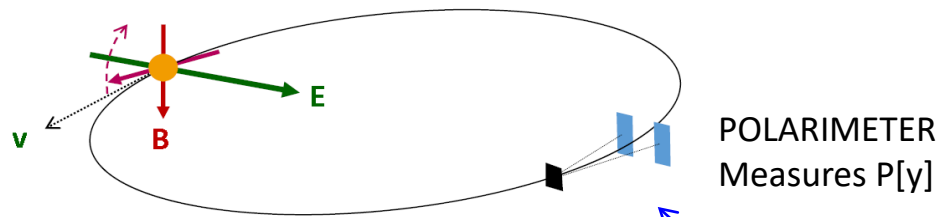
Idea: create an oscillating polarization, look for resonance at ω .



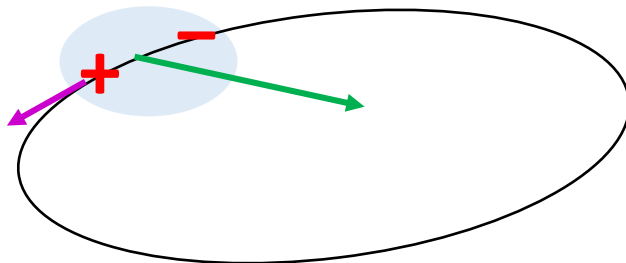
Description of the axion search (demonstration)

We start from a known COSY setup: polarized deuterons at 0.97 GeV/c
Beam is electron (pre-)cooled with long, in-plane polarization lifetime.

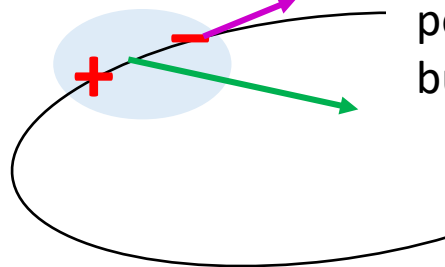
Standard picture of frozen spin EDM experiment:



For deuteron in COSY

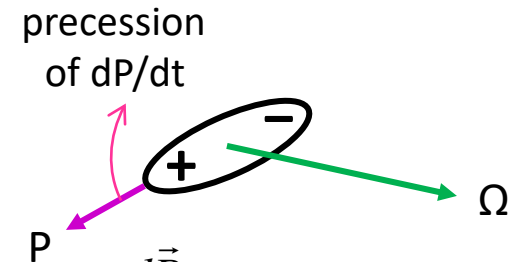


after ~3 revolutions



polarization reversed
but EDM has flipped

combination allows
signal accumulation



$$\frac{d\vec{P}}{dt} = \vec{P} \times \vec{\Omega}$$

$$\vec{\Omega} = \frac{\eta q}{2mc} (\vec{E} + c\vec{\beta} \times \vec{B})$$

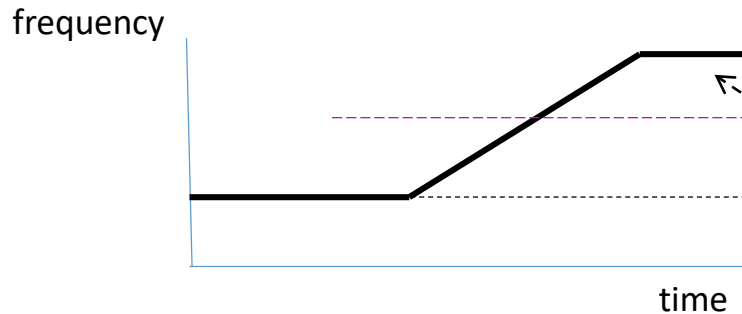
EDM accumulation
depends on $v \parallel p$.

Axion accumulation
is automatic if the
frequency is right.

But phase is unknown.

In the axion search:
(spin is not frozen)

The frequency is unknown.
So scan for it.

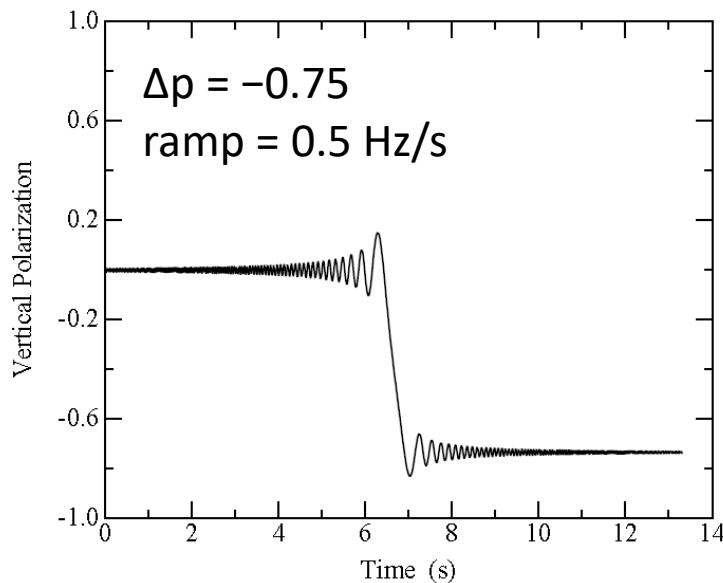


You may be lucky enough to
cross the axion frequency.

Axion frequency

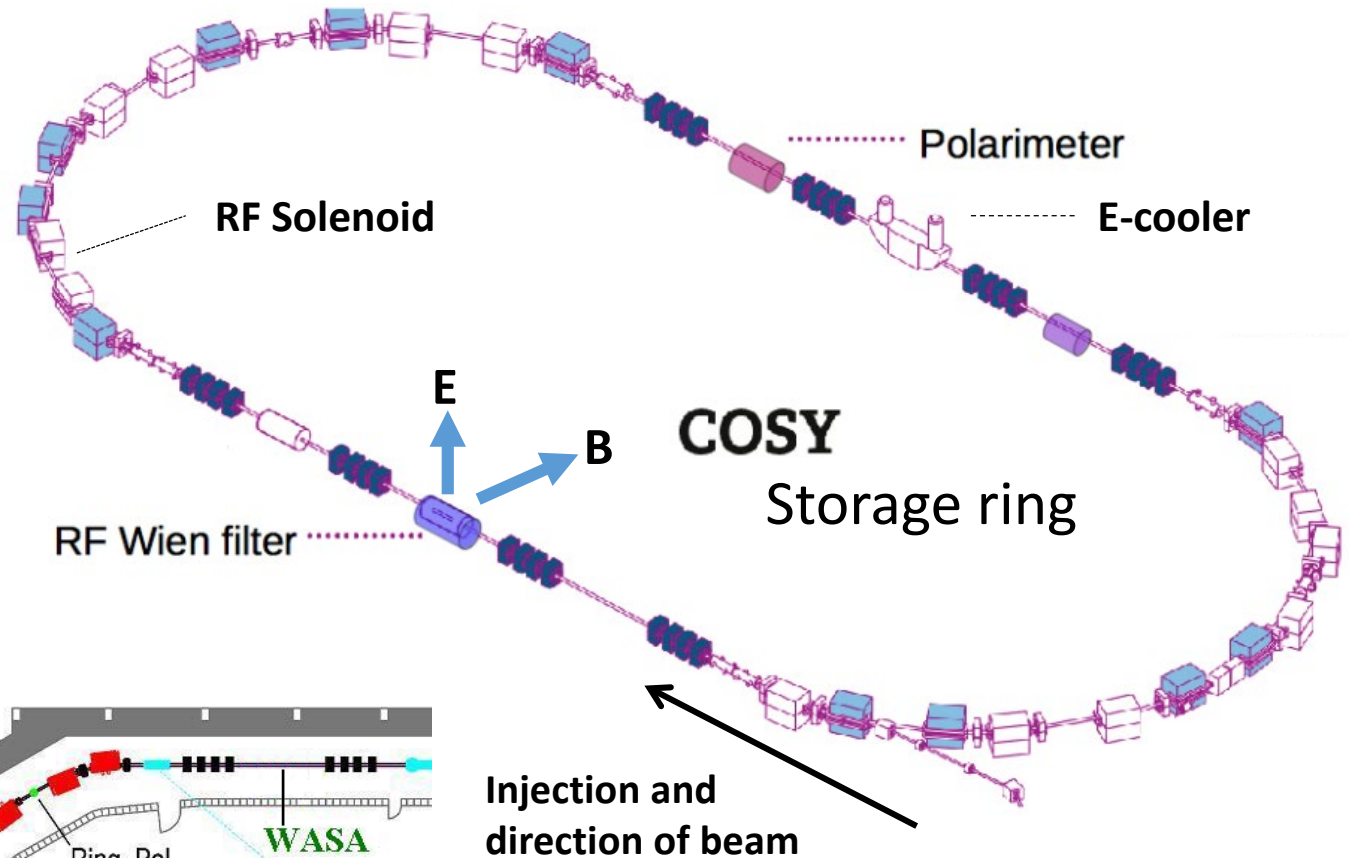
Flat regions allow for a precise
before-after comparison.

Then you will measure:



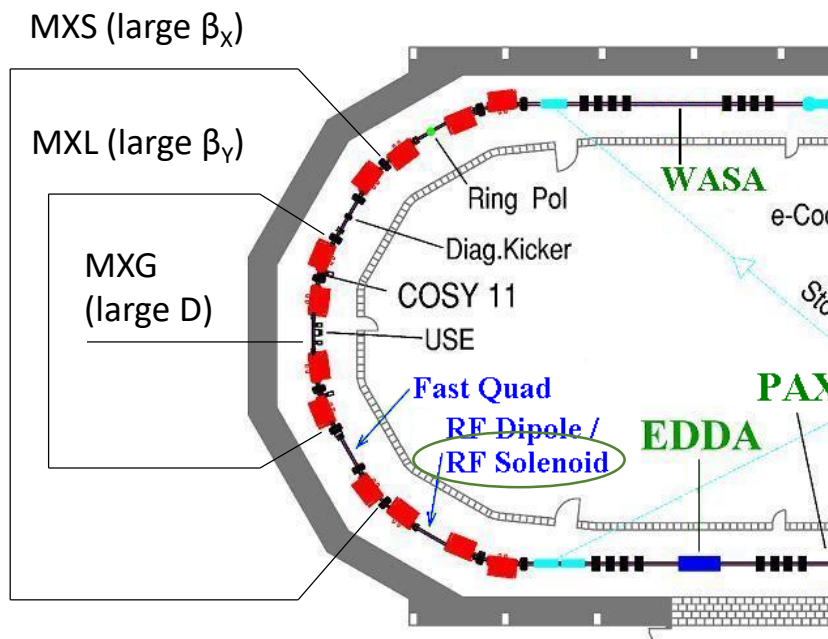
Critical assumptions:

- > The axion is coherent in space.
All particles oscillate together.
The axion is present in all of COSY.
Axions are dense.
- > The axion is coherent in time.
It takes a few seconds to pass COSY.
(only slow axions stay in galaxy,
 $v[\text{esc}] \sim 550 \text{ km/s}$)

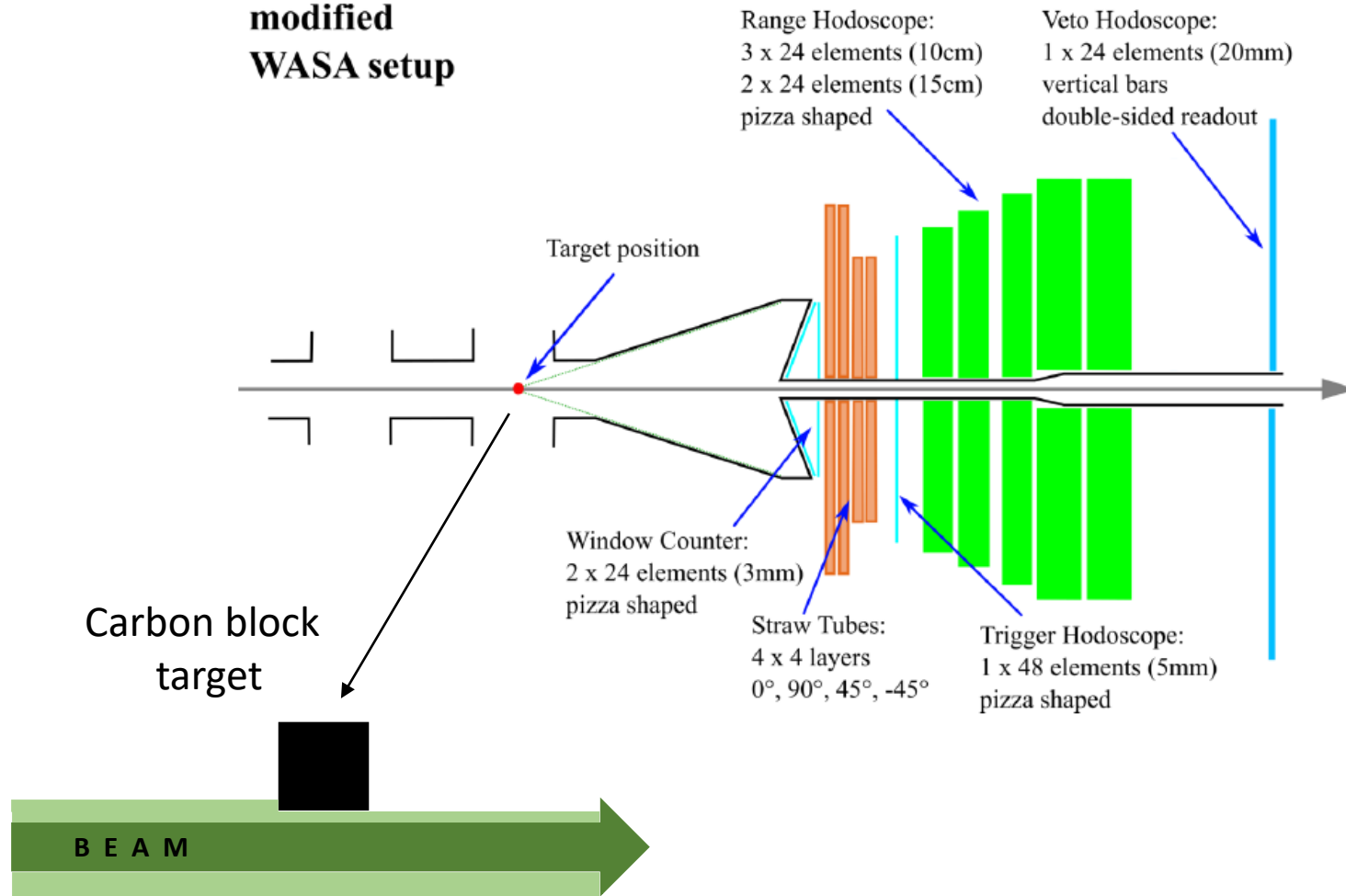


Injection and
direction of beam

Electron cooling and adjustment of
sextupole fields are essential for a
long horizontal polarization lifetime.

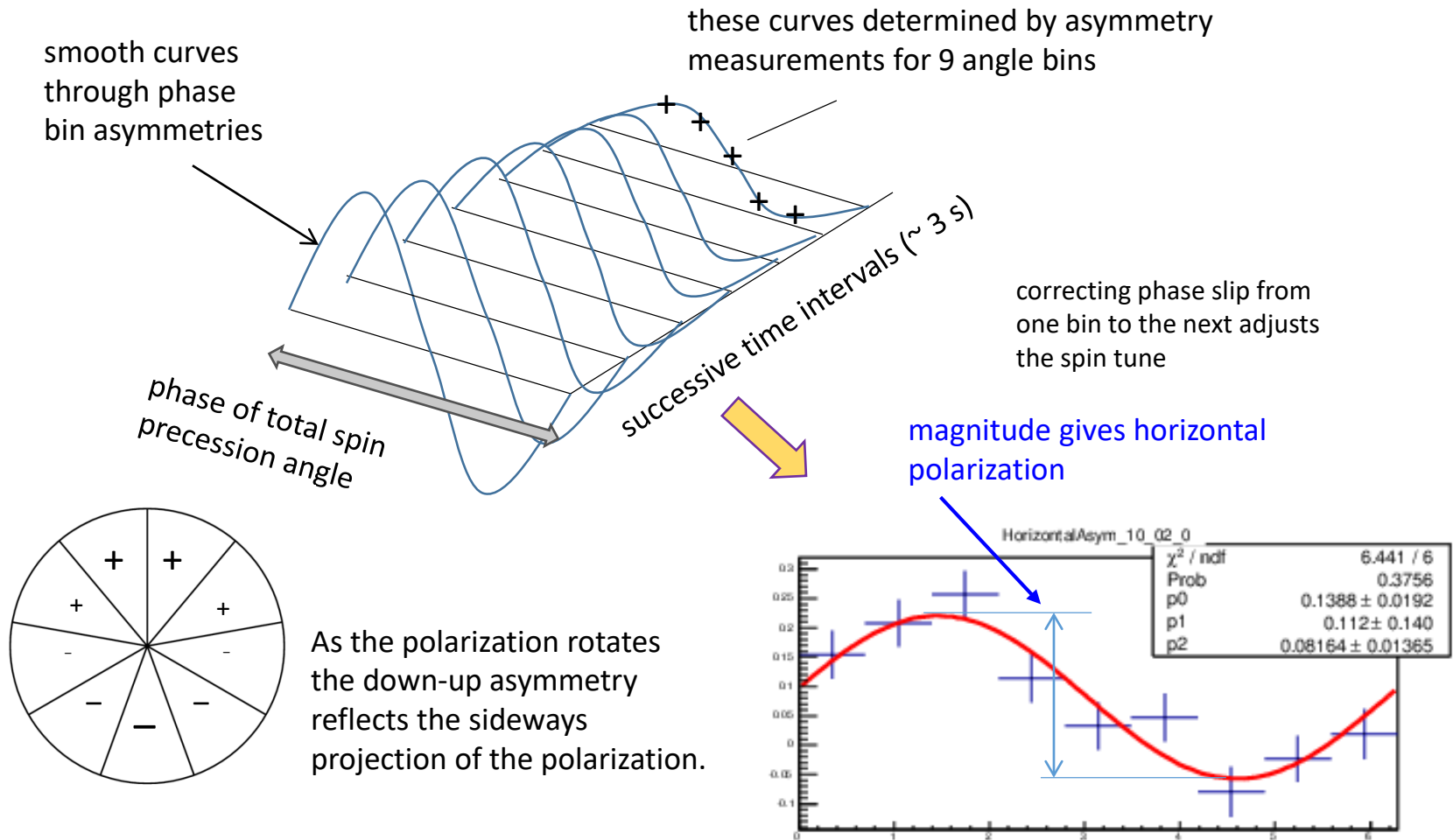


modified WASA setup



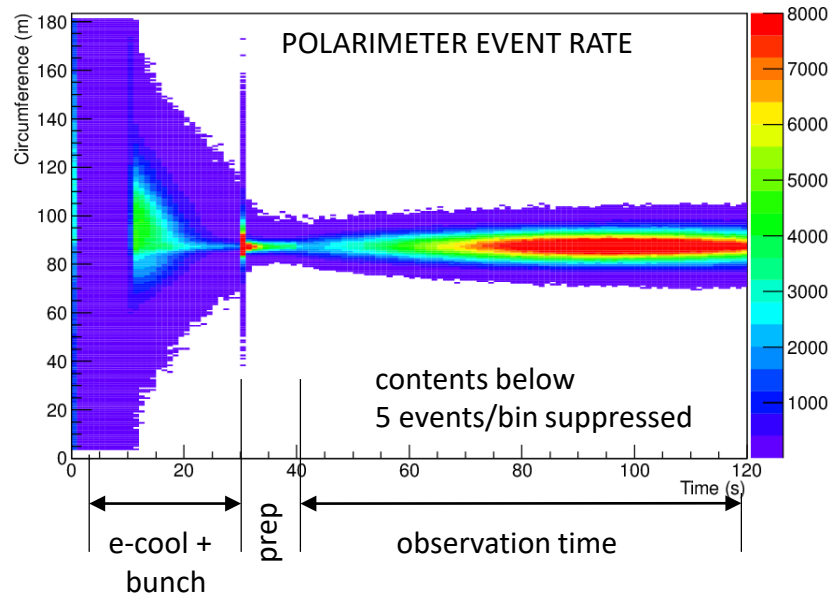
New data acquisition procedure – time stamp every event

- Count turn number (bunched beam) → distribution of turn number
 - Compute total spin precession angle ↘ fraction yields beam distribution
 - Bin by phase around the circle
 - Compute asymmetry in each bin
- based on integral part of turn number

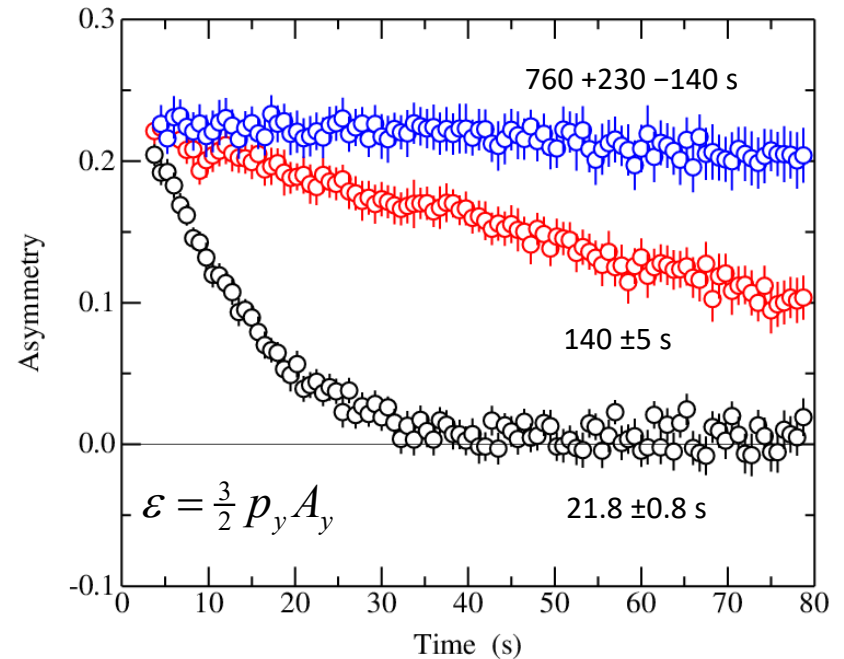


Sample data

Distribution of beam around the ring as a function of time in the store.

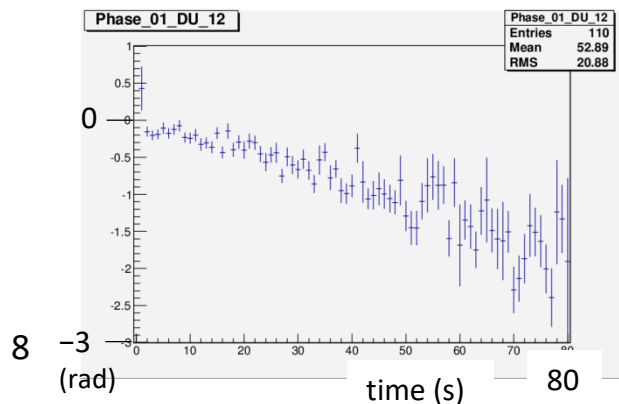


Sample measurements of horizontal polarization loss (corrected for positive bias)



Times are exponential decay rates.

phase in a single store with fixed spin tune



Program searches for highest amplitude in a narrow range. To get maximum asymmetry stationary in one angle bin, spin tune must be accurate to $< 1e-6$. Normal scatter is usually $< 1e-7$.

Best error in phase is $\sim 3^\circ /s$.

Downward slope means spin tune wrong by $3e-8$ ($\delta \sim 10\%$).

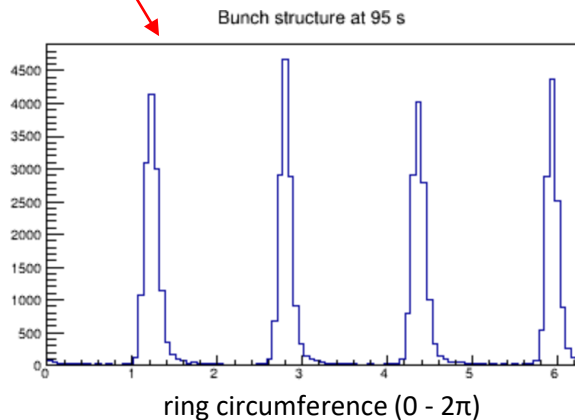
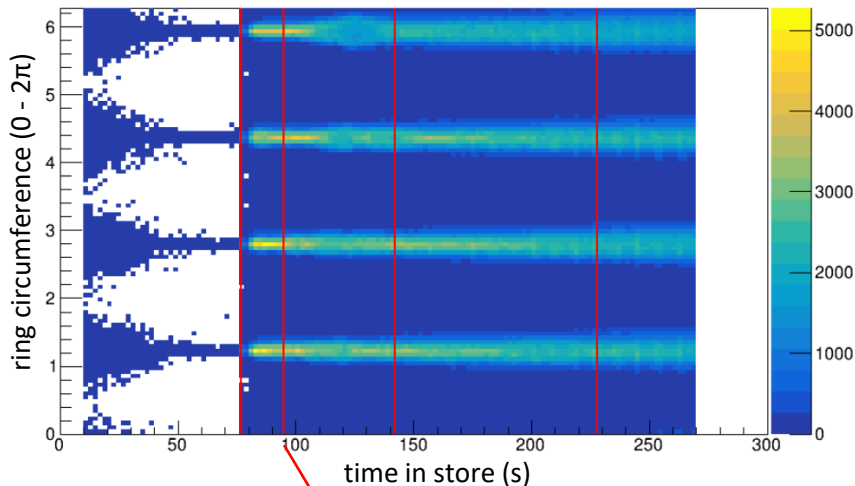
EDM ring requirement is $1e-9$ from feedback.

Problem: Axion phase is unknown.

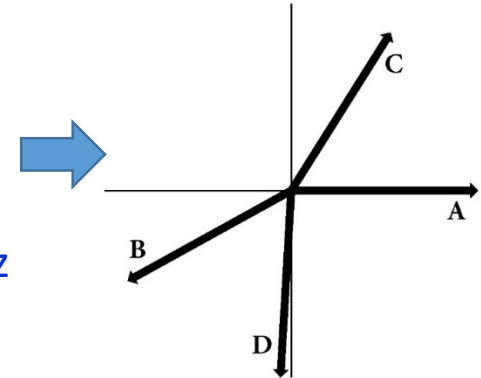
Conduct search simultaneously with beams having perpendicular polarizations.

Combine results for better signal.

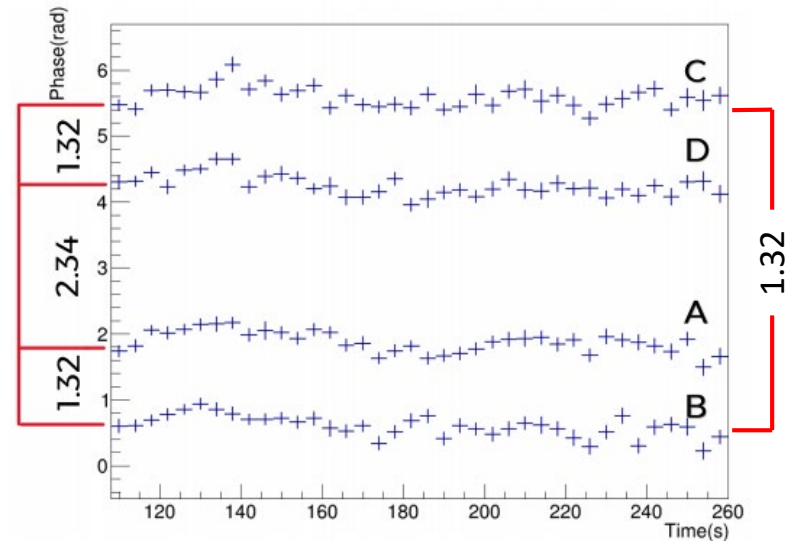
Plan: Use 4 bunches...



Operate RF solenoid at
 $f_{\text{REV}}(1+G\gamma)$
 $= 629771.9 \text{ Hz}$



verification



Calibration based on tracking through rotational history

Calculation shows size of jump
as a function of oscillating
EDM size.

Red – axion in phase (0°)

Magenta – phase = 90°

Black – sum in quadrature

Ramp rate = 10 Hz/s

(most data taken at 0.1 or 0.05 Hz/s)

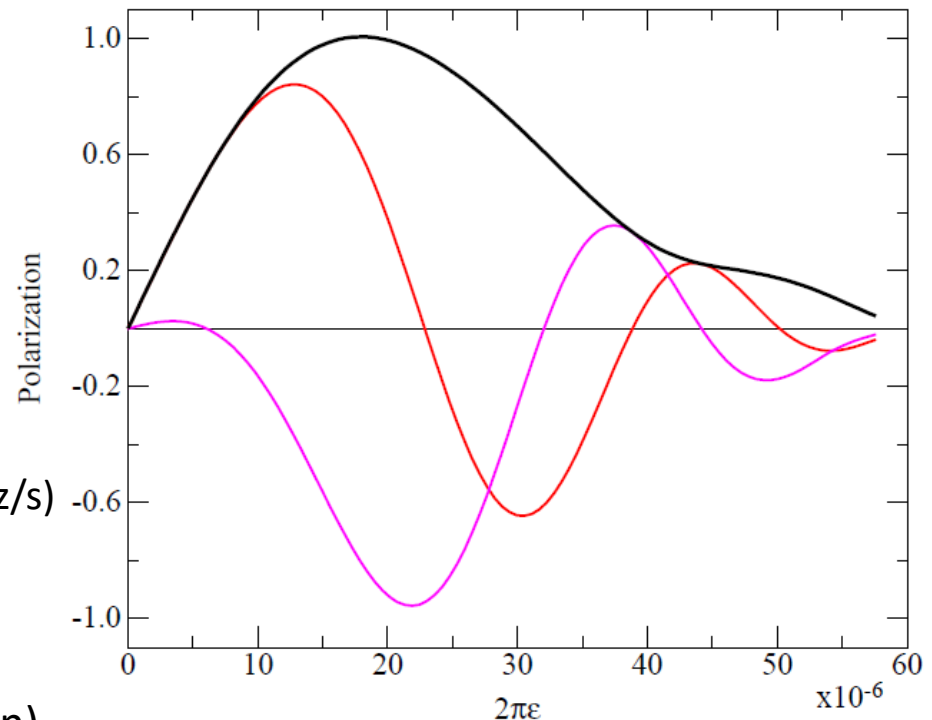
For any other ramp rate,
scale by $1/\sqrt{r}$.

ε is the EDM rotation (rad/turn),

so $\omega_{EDM} = f_{REV}\varepsilon$.

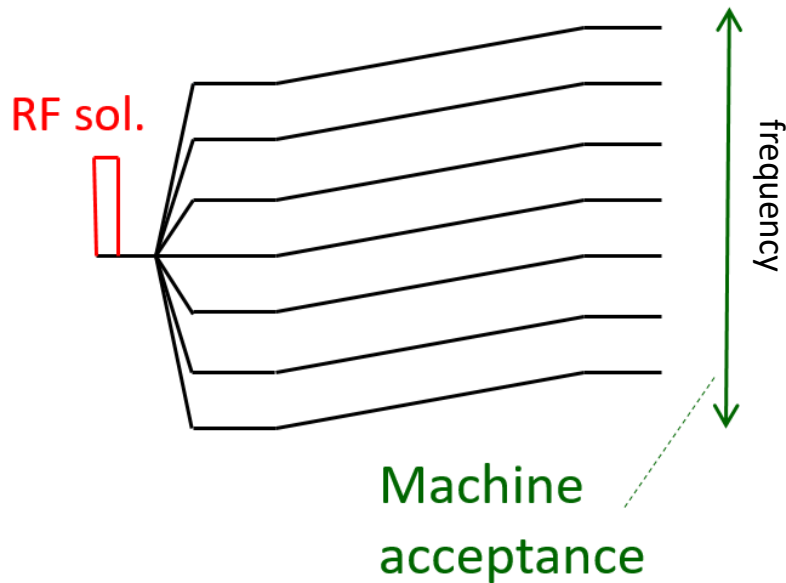
Electric dipole: $d = (6.242 \times 10^{20}) \frac{\hbar\omega_{EDM}(2\pi R)}{v(B\rho)}$

conversion to e·cm

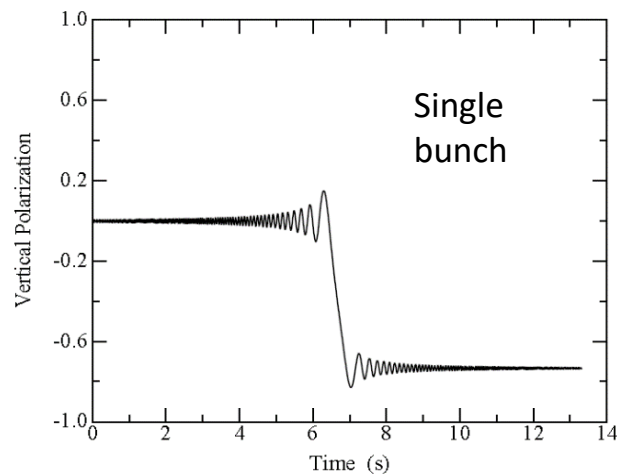


Ring circumference

Run a series of scans in frequency



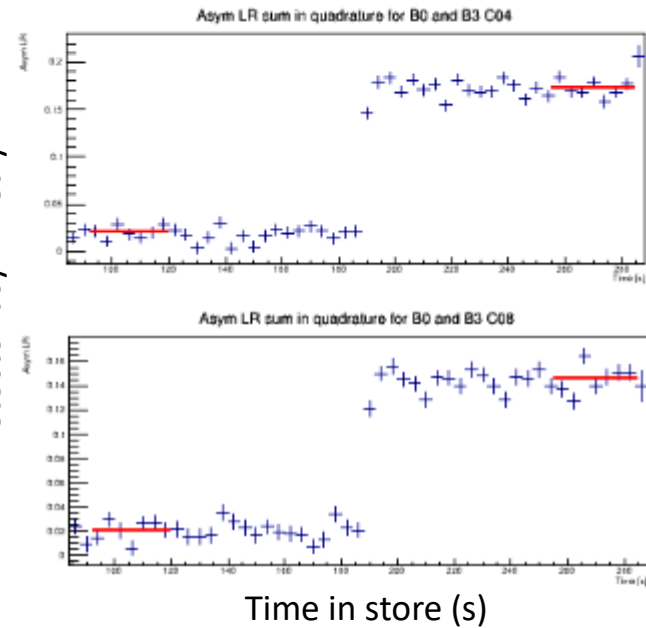
expected signal



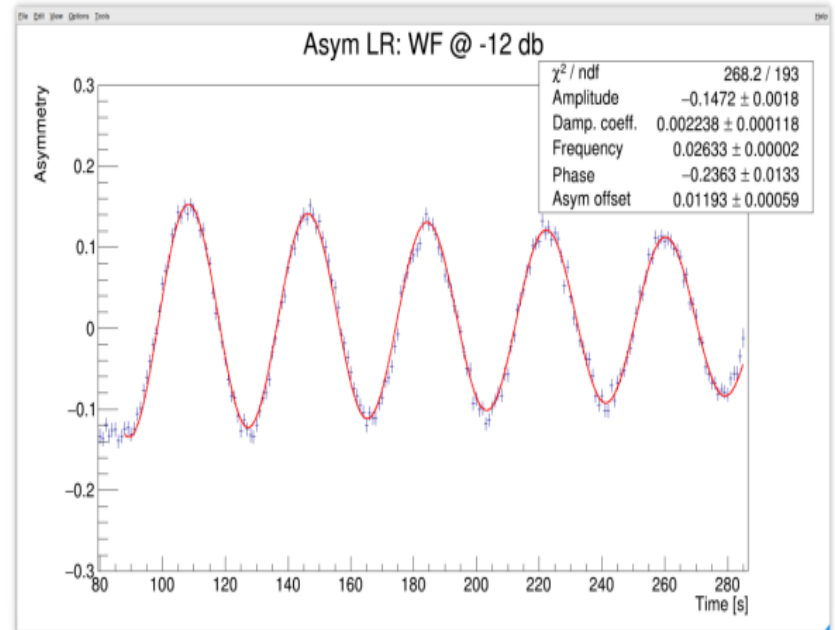
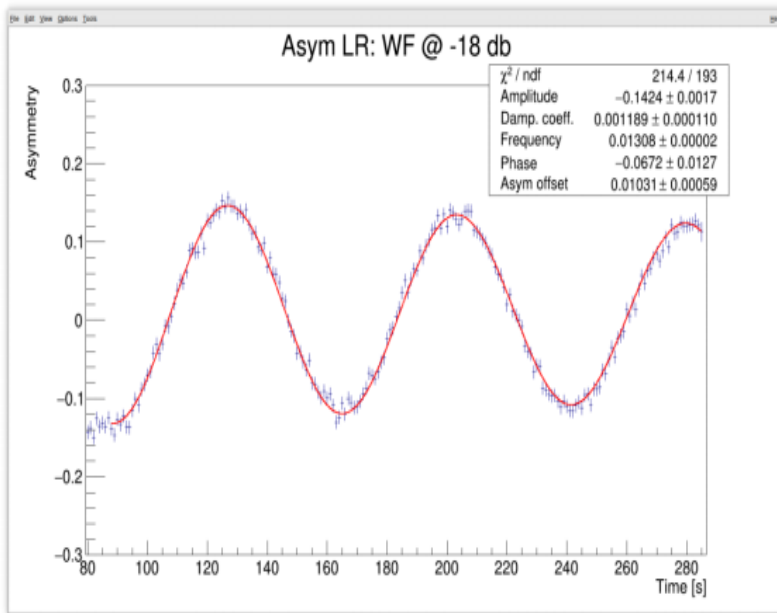
Test calibration

A system test was made using an RF Wien filter to generate signals similar to an axion. The signal response was also used to calibrate the polarization jump.

Detector asymmetry



Check against Wien filter:



Frequency at maximum power level produces average $\varepsilon = 1.42 \times 10^{-7}$

Scale for ramp rate

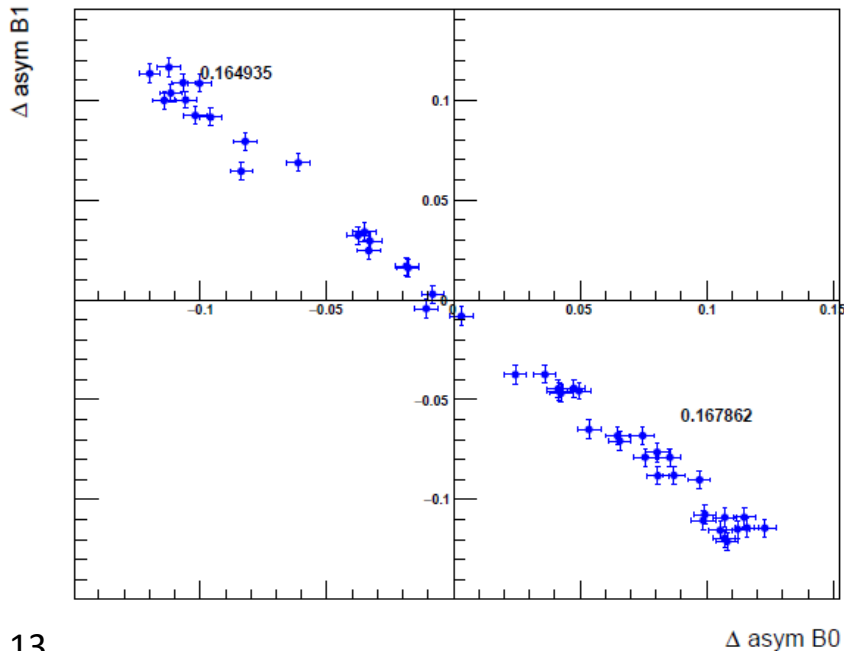
predicted for 0.1 Hz/s $\Delta p = 0.736$

predicted for 0.05 Hz/s $\Delta p = 0.915$

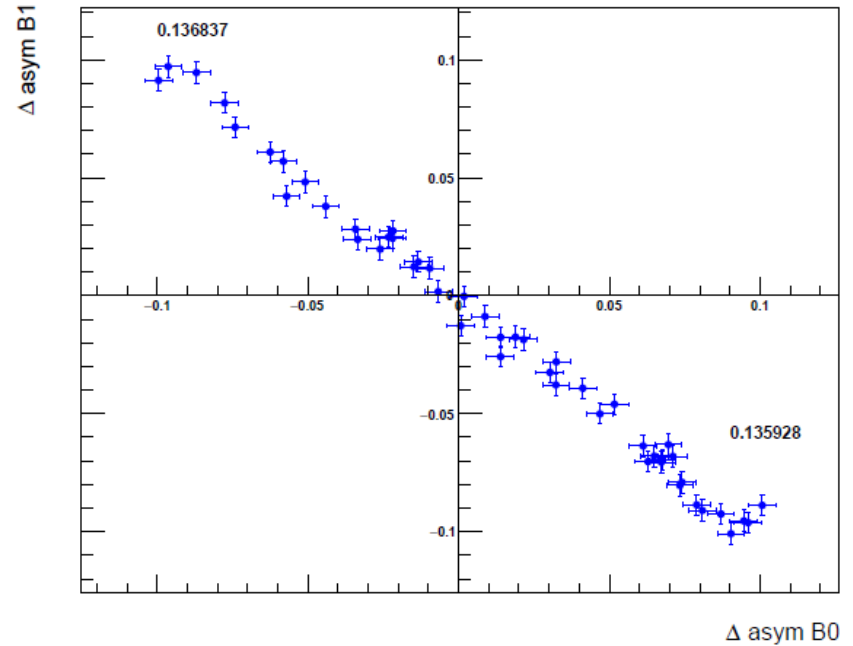
In practice, random phase distributions resulted in a spread of jumps. We want the maximum.

Values in figures are the length of the diagonals.

Δ asym B0 and B1 (ramp speed 0.056)



Δ asym B0 and B1 (ramp speed 0.112)



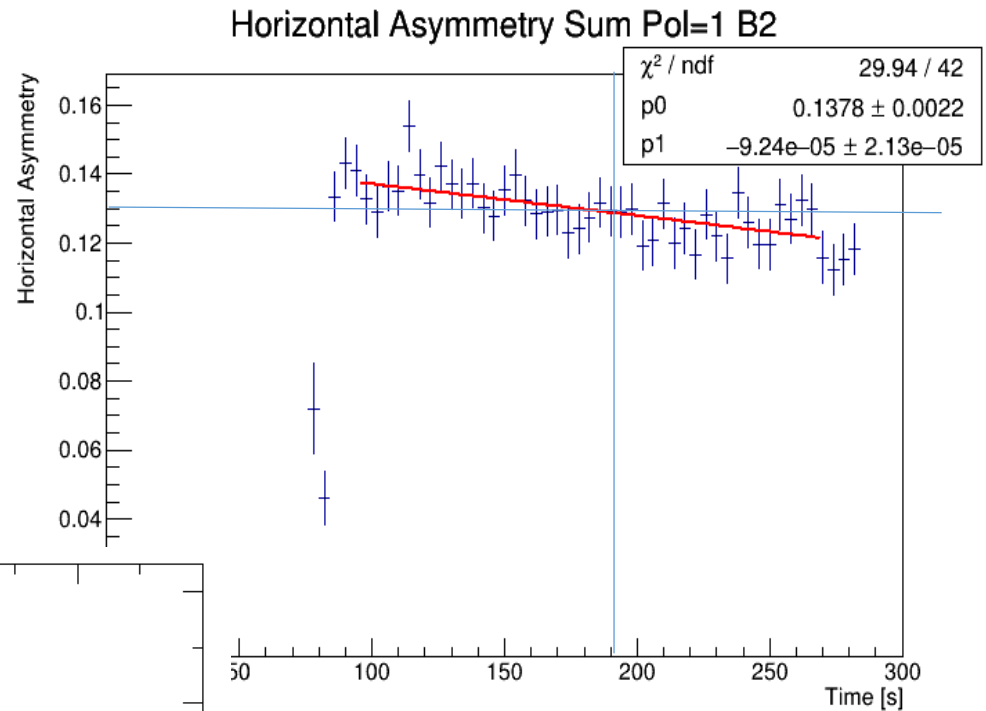
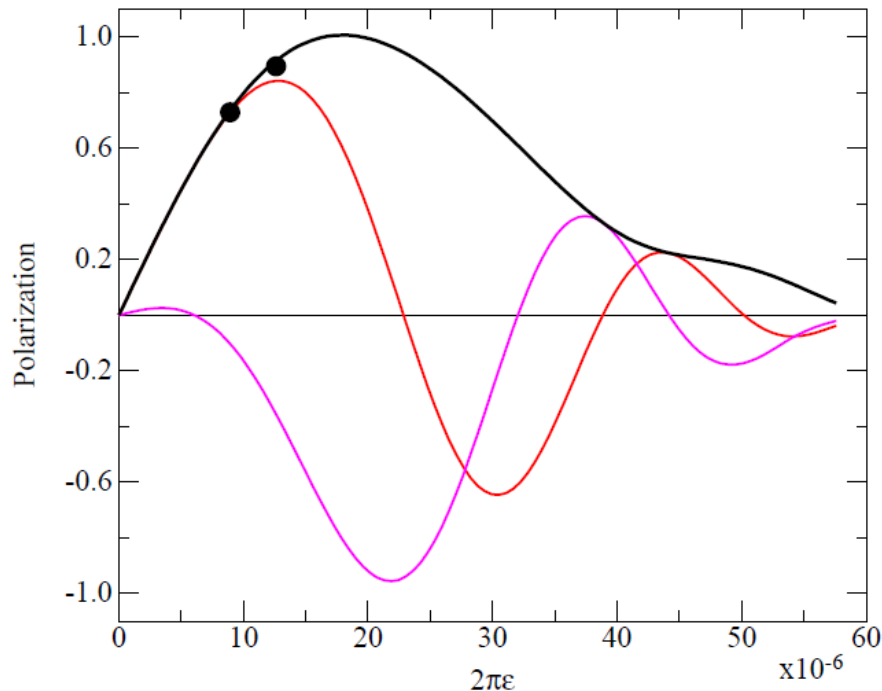
Typical jumps averaged over all cases:

$$0.1 \text{ Hz/s} \rightarrow \Delta(\text{pA}) = 0.0948$$

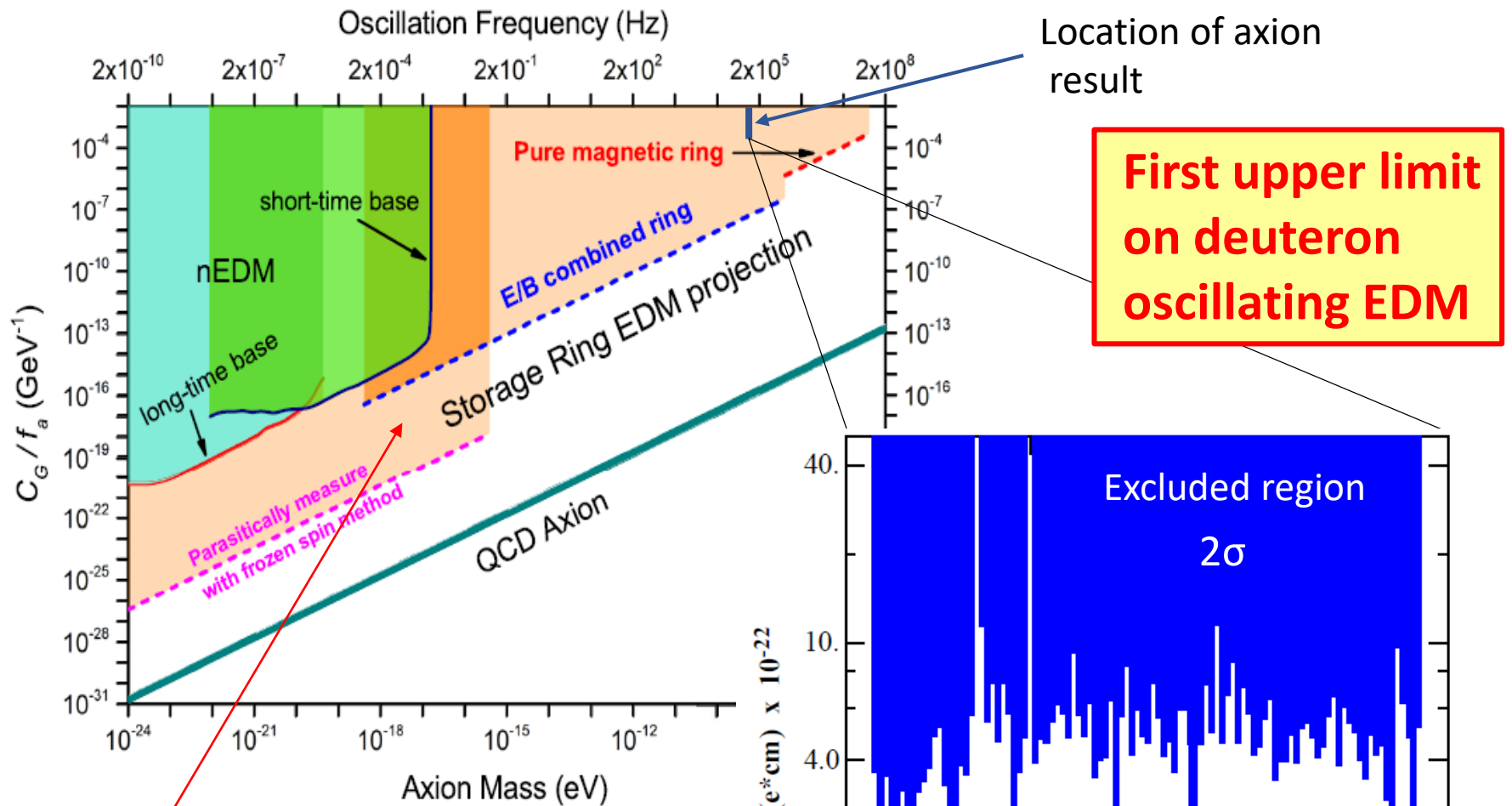
$$0.05 \text{ Hz/s} \rightarrow \Delta(\text{pA}) = 0.1163$$

At 190 s average
horizontal
polarization
is 0.13.

Scales polarizations
to 0.73 and 0.90



The Wien filter test
confirms the
calibration from
the rotation model.



Anticipated limit achievable using:

- Deuteron/proton beams
 - Combined electric/magnetic ring
 - Parasitic analysis using all-electric ring
- see S.P. Chang et al., PRD **99**, 083002 (2019)

