



# NEUTRINO HUNTERS AROUND THE WORLD

Zara Bagdasarian

AUGUST 21 2018 | GEORGIAN-GERMAN SCHOOL AND WORKSHOP IN BASIC SCIENCE

Mitglied der Helmholtz-Gemeinschaft

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[@ZaraBagdasarian](https://twitter.com/ZaraBagdasarian)



# NEUTRINOS IN THE STANDARD MODEL

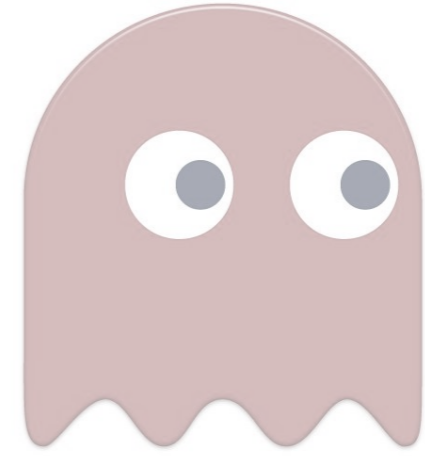


- And beyond



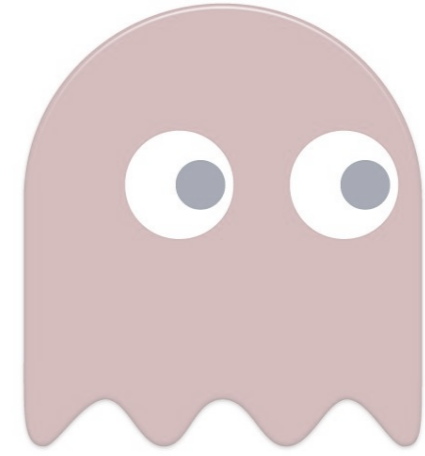
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- No electric charge -> no electro-magnetic interaction
- No color -> no strong interactions
- Weak interaction -> very very small cross sections of interactions with matter



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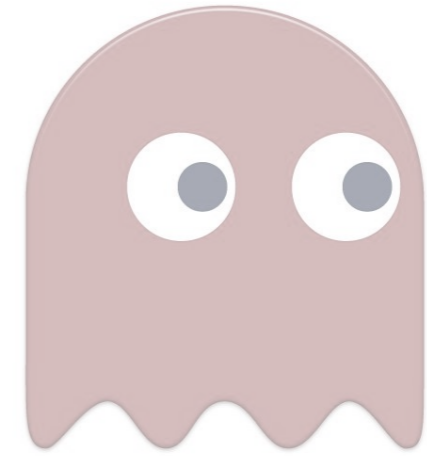


**Often referred to as ghost particles**



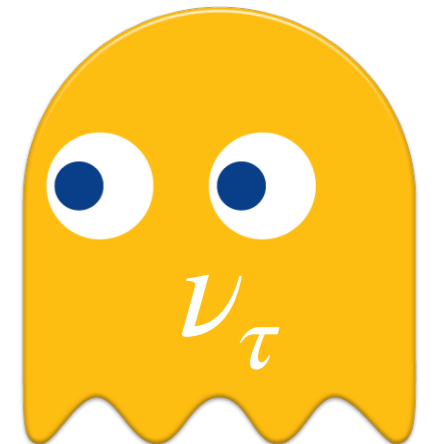
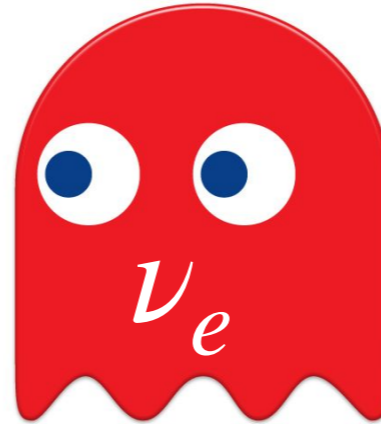
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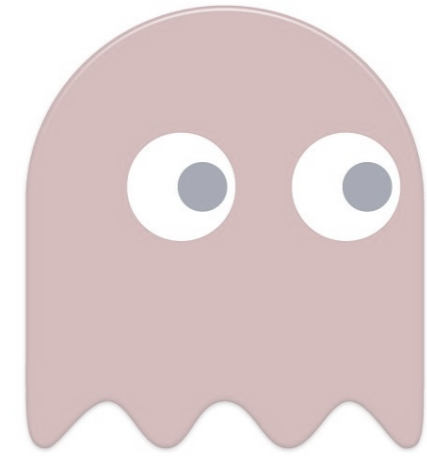
Often referred to as ghost particles

- Tiny masses ( $0.04 \text{ eV} < m < \approx 1 \text{ eV}$ )
- Three flavours
- Changing flavours, aka oscillating



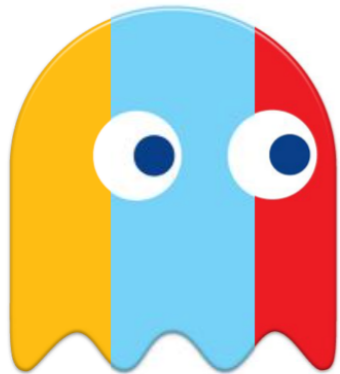
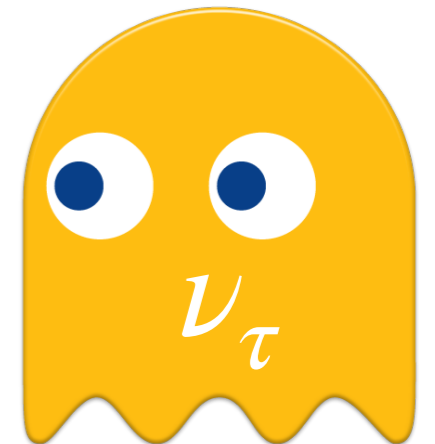
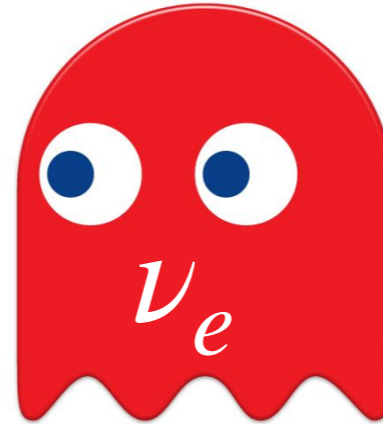
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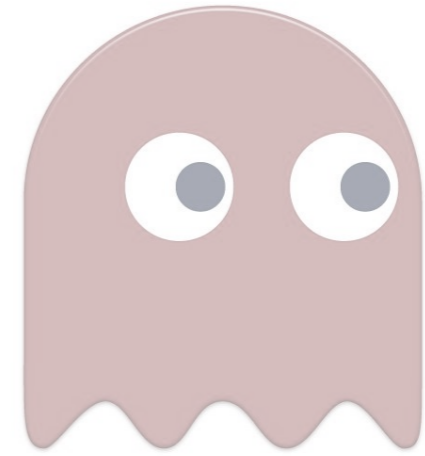
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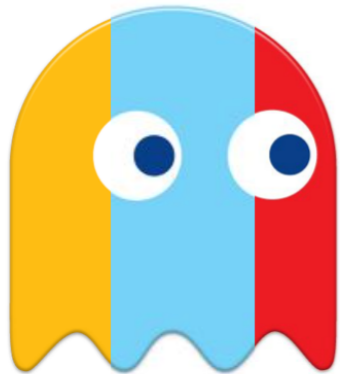
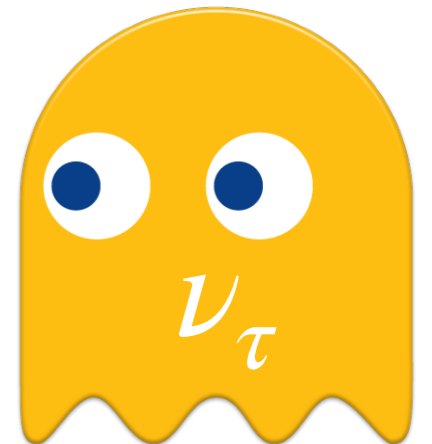
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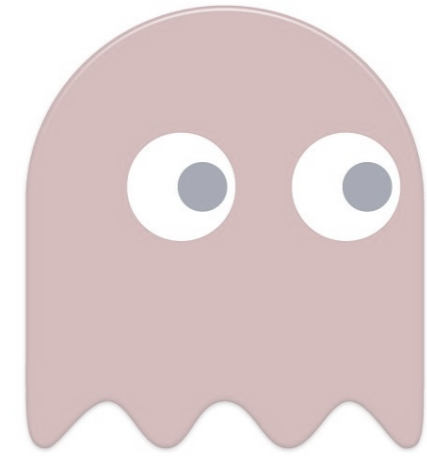
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Ghost particles full of mysteries and surprises

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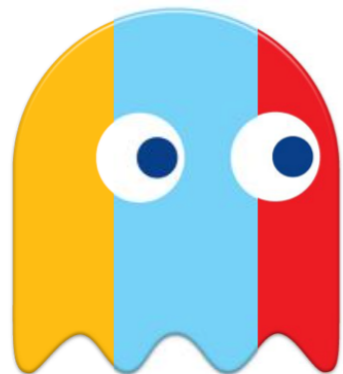
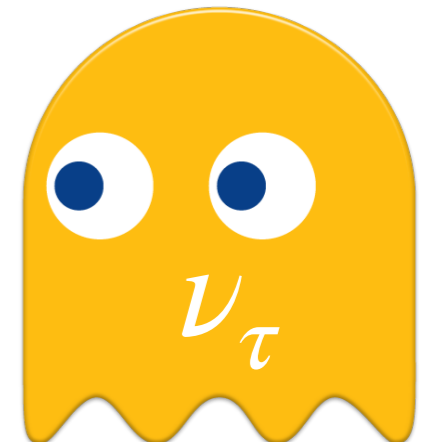
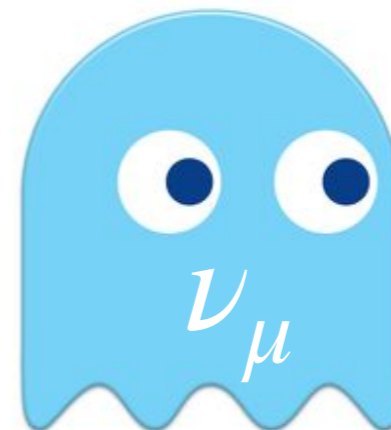
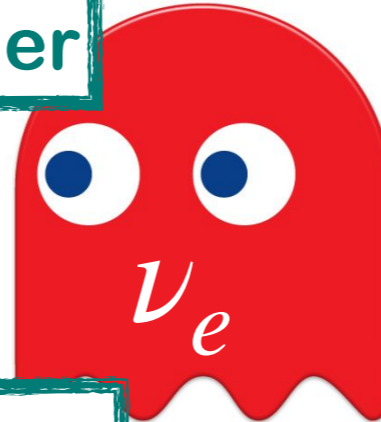
Often referred to as ghost particles

- Tiny masses
- Three flavours
- Changing flavours, aka oscillating

exact values and order

Or more?

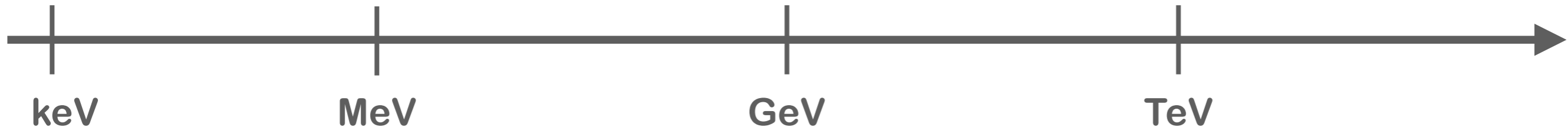
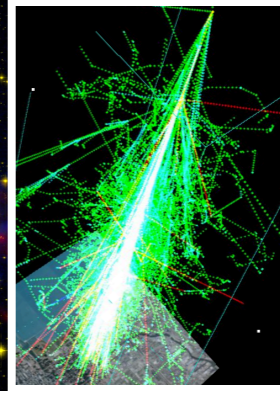
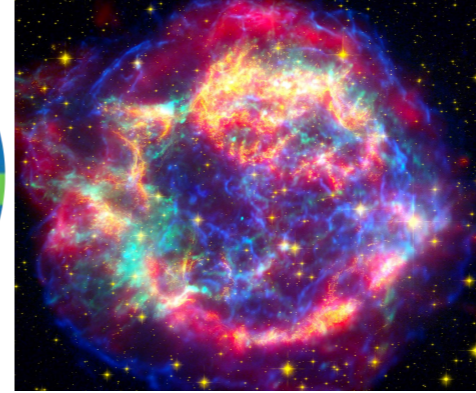
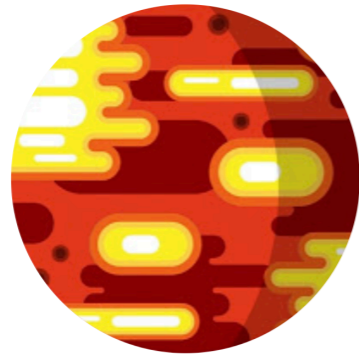
Exact oscillation parameters



Ghost particles full of mysteries and surprises



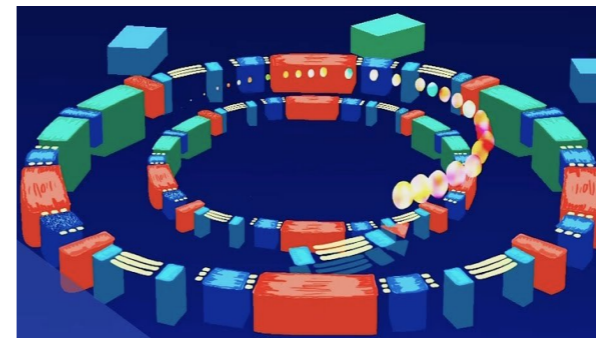
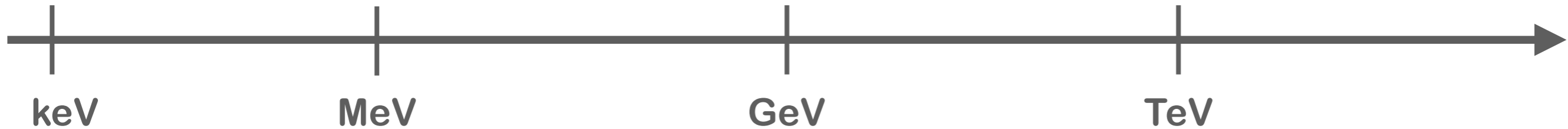
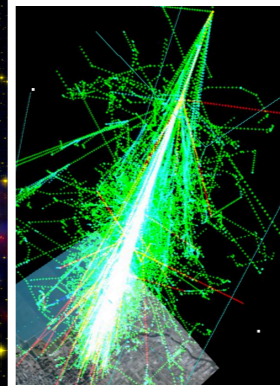
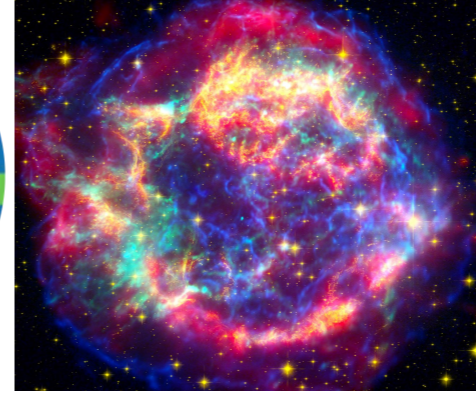
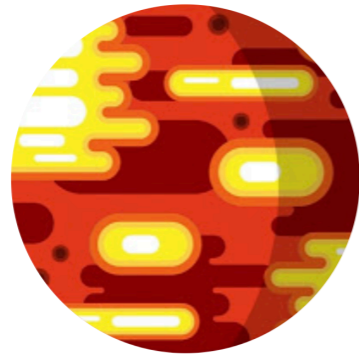
# NEUTRINO SOURCES



## Direct information from the source

Pictures adapted from NASA, Kurzgesagt, the Royal Institution and University of Chicago

# NEUTRINO SOURCES



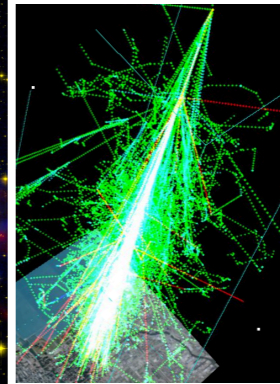
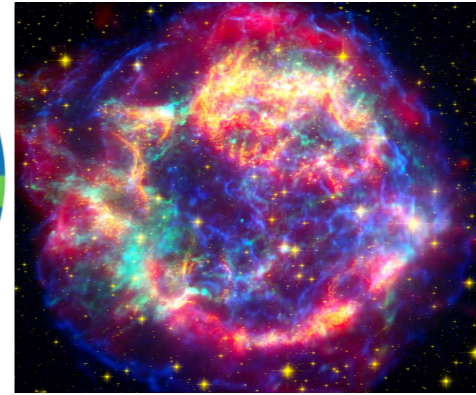
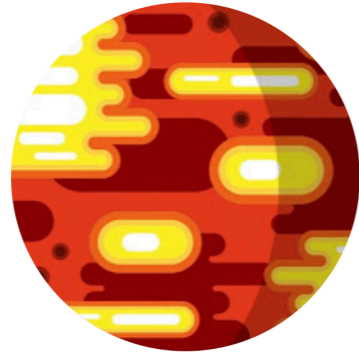
Direct information from the source

Better understanding of neutrino nature

Pictures adapted from NASA, Kurzgesagt, the Royal Institution and University of Chicago



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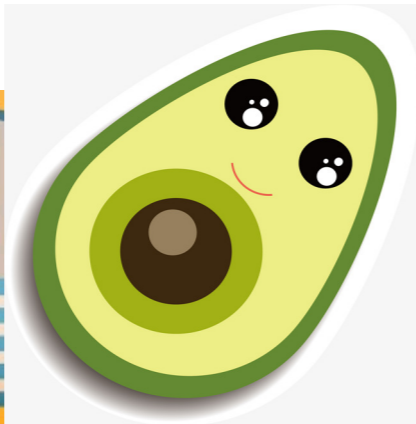


keV

MeV

GeV

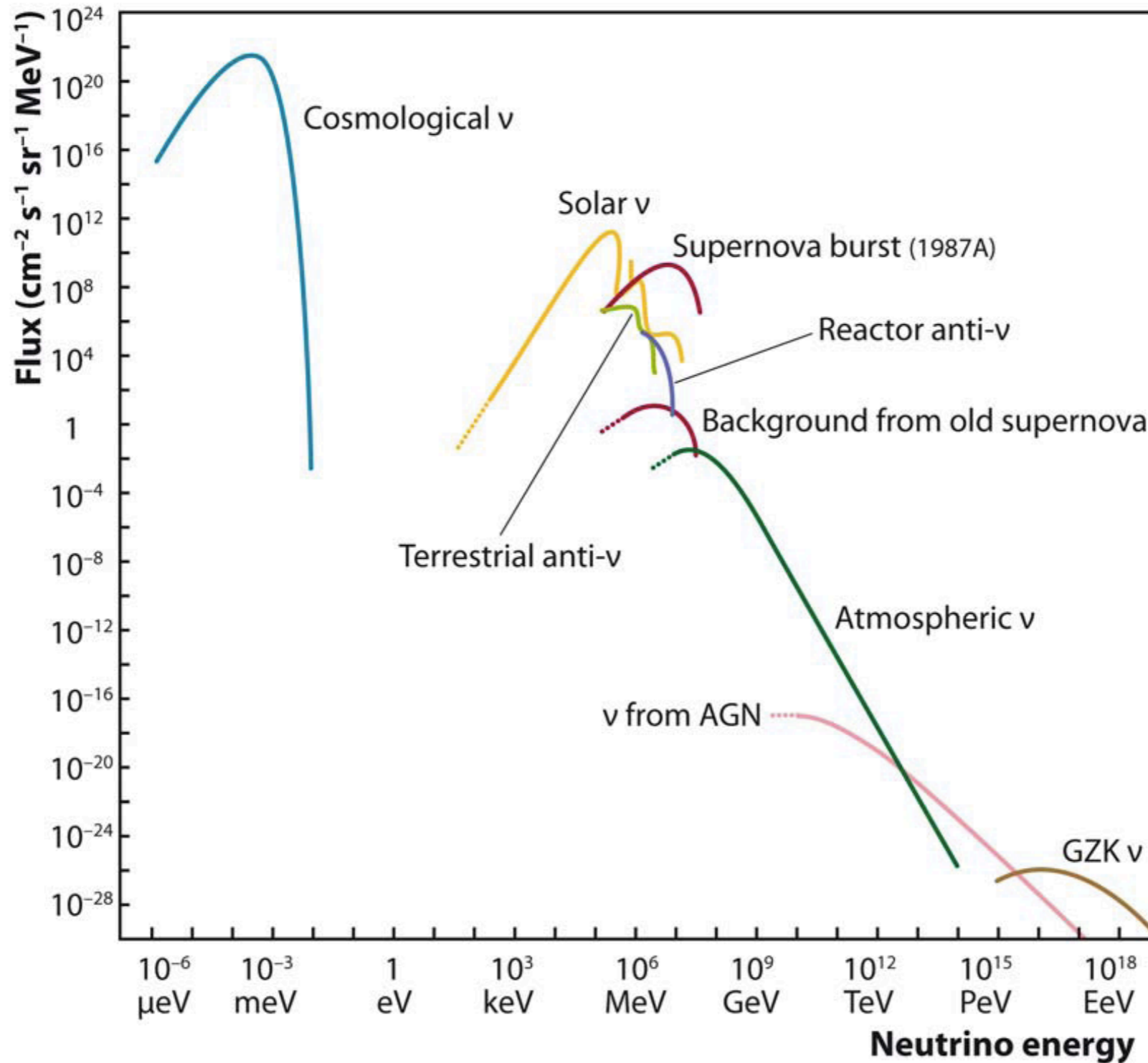
TeV



**Caution! Cute babies radiating in neutrinos**

Pictures adapted from NASA, Kurzgesagt, the Royal Institution and University of Chicago

# NEUTRINO SOURCES





# A BIT OF HISTORY

Year 1930

*Original - Photocopy of PLC 0393*  
Abschrift/15.12.56 **PI**

Offener Brief an die Gruppe der Radioaktiven bei der  
Gauvereins-Tagung zu Tübingen.

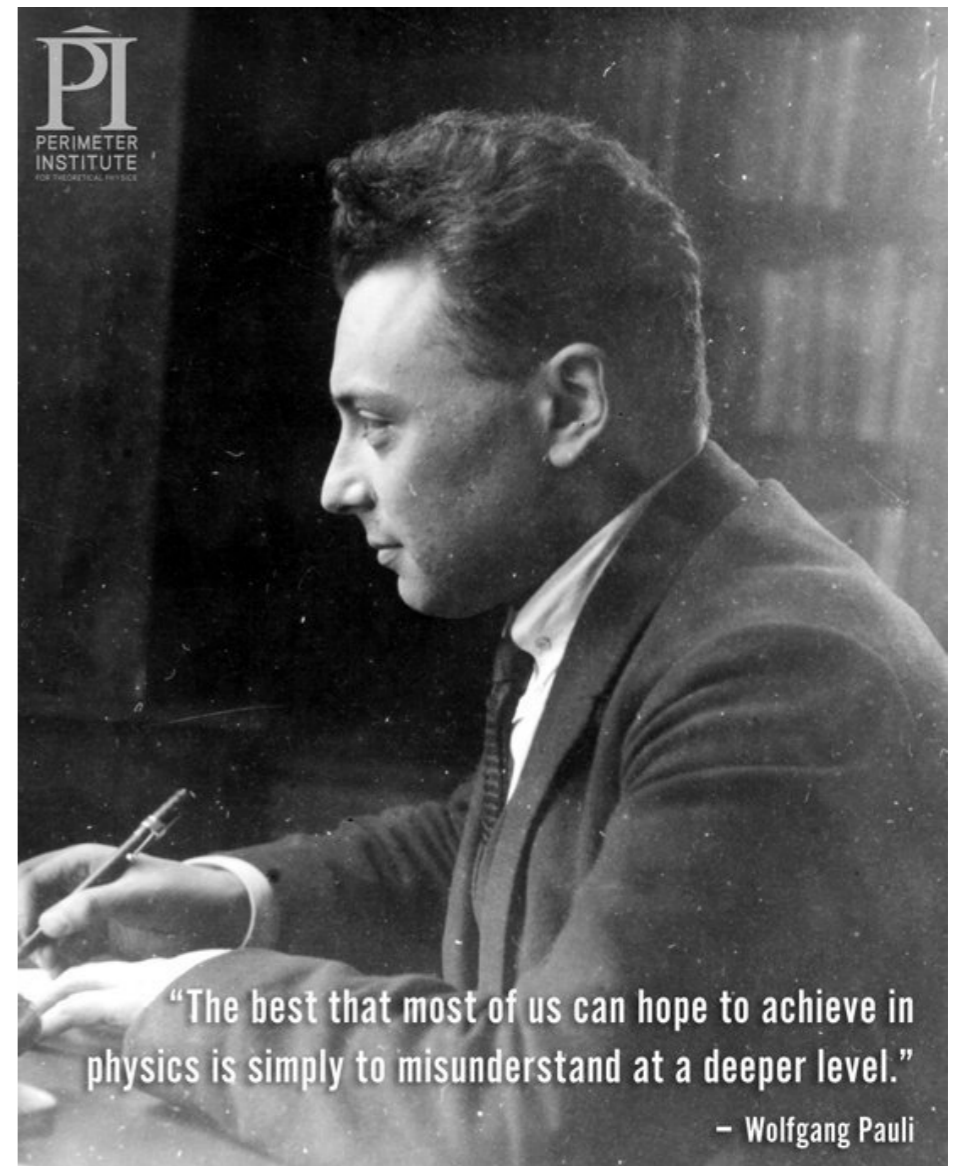
Abschrift

Physikalisches Institut  
der Eidg. Technischen Hochschule  
Zürich

Zürich, 4. Dez. 1930  
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich huldvollst  
anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich  
angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie  
des kontinuierlichen beta-Spektrums auf einen verweifelten Ausweg  
verfallen um den "Wechselsatz" (1) der Statistik und den Energiesatz  
zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale  
Teilchen, die ich Neutronen nennen will, in den Kernen existieren,  
welche den Spin 1/2 haben und das Ausschliessungsprinzip befolgen und  
sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie  
nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen  
müsste von derselben Grossenordnung wie die Elektronenmasse sein und  
jedenfalls nicht grösser als 0,01 Protonenmasse.- Das kontinuierliche  
beta-Spektrum wäre dann verständlich unter der Annahme, dass beim  
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wird, derart, dass die Summe der Energien von Neutron und Elektron  
konstant ist.



Source: Perimeter Institute

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Gle...

Let's call it  
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# A BIT OF HISTORY

Year 1930

Year 1932 James Chadwick



Discovered a neutral particle. Calling dips on "neutron"!

Let's call it "neutron"



"The best that most of us can hope to achieve in physics is simply to misunderstand at a deeper level."

– Wolfgang Pauli



# A BIT OF HISTORY

Year 1930

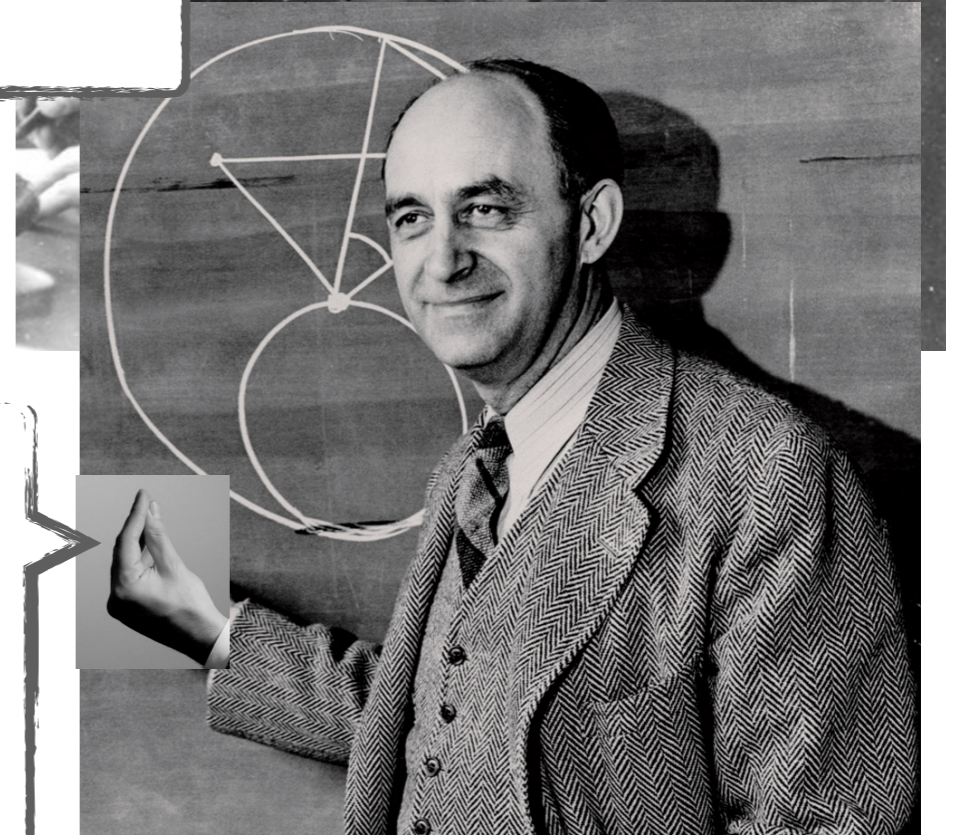
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Discovered a neutral particle. Calling dips on "neutron"!

Let's call it "neutron"

Let's call this one "neutrino"



Year 1934



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Year 1930

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Abschrift/15.12.56 **PH**

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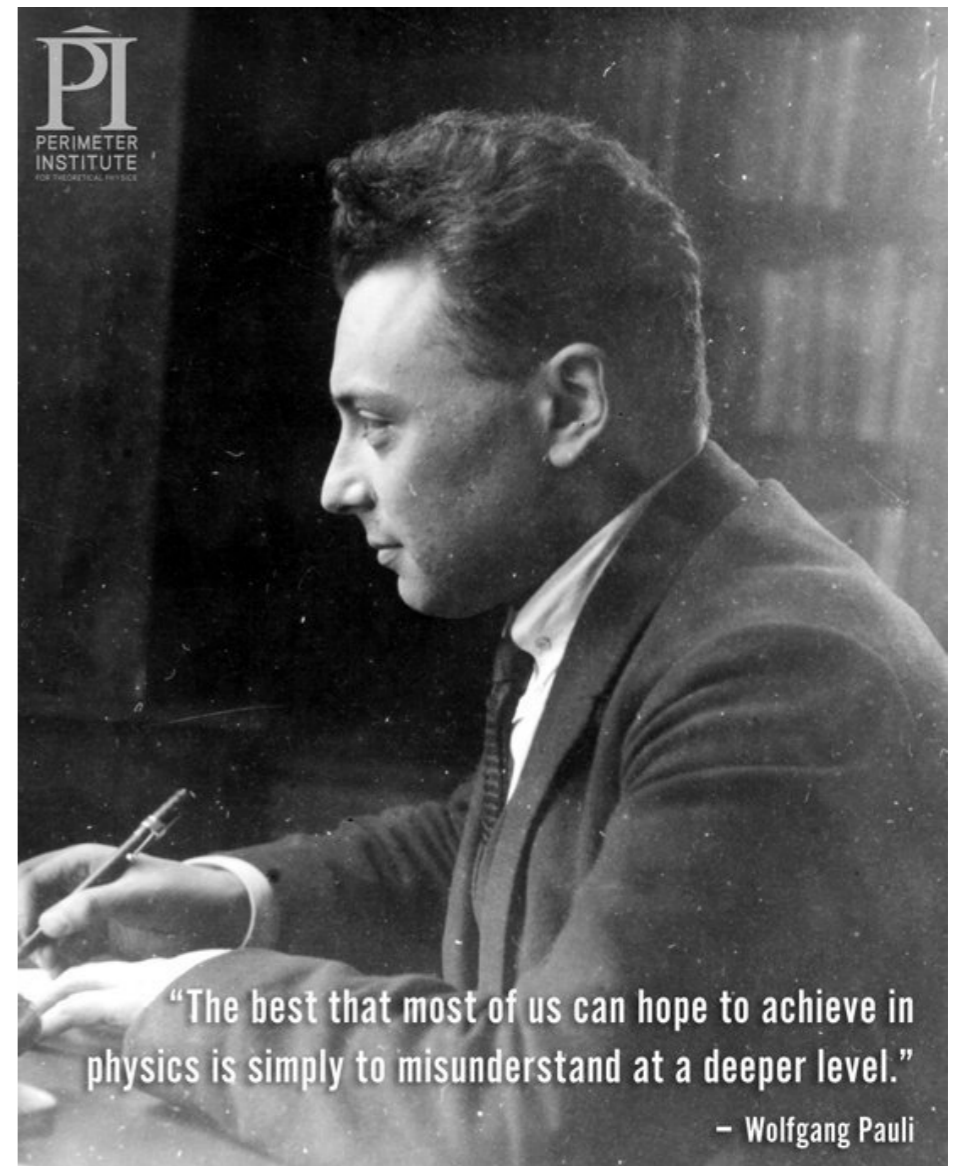
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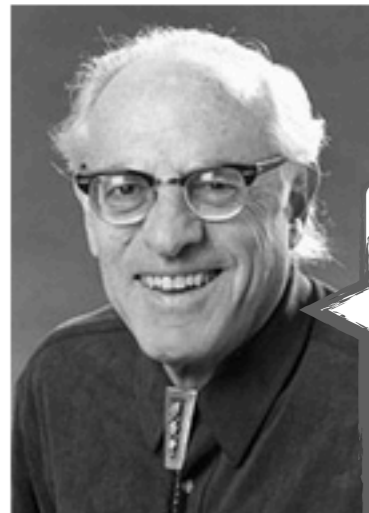
"I have done a terrible thing. I have postulated a particle that cannot be detected"

# DETECTION OF THE NEUTRINO

Everybody: "Nobody can do it!"

Cowan and Reines:

Challenge accepted!



Telegram Pauli!

Year 1956

*Detecting the Poltergeist*



Hanford Team 1953

# NEUTRINOS AND NOBEL PRIZES





# DETECTION OF THE ~~NEUTRINO~~



- Nobel Prize in Physics 1995



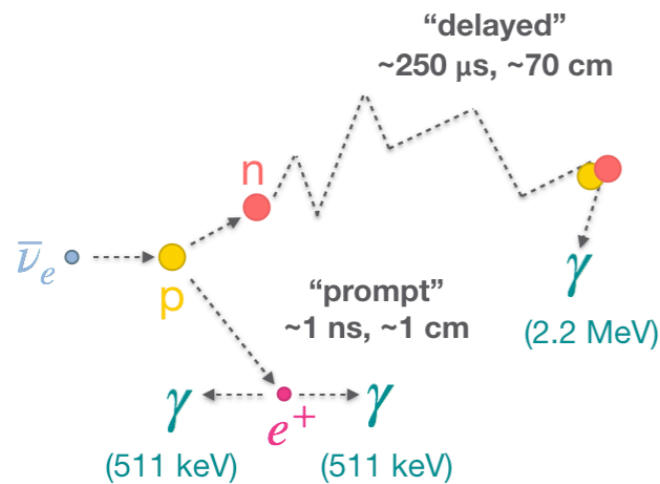
© University of California Regents  
**Frederick Reines**  
 Prize share: 1/2

*Detecting the Poltergeist*



Hanford Team 1953

# DETECTION OF THE ANTI-NEUTRINO





# DETECTION OF THE MUON NEUTRINO



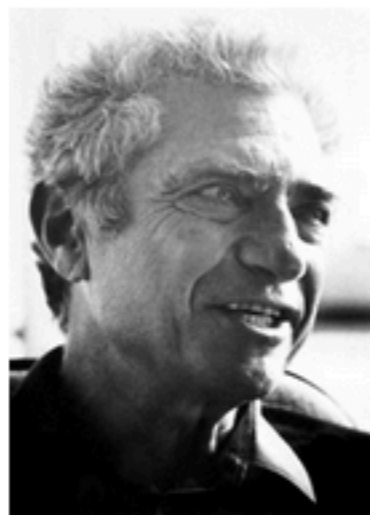
- Nobel Prize in Physics 1988



Leon M. Lederman  
Prize share: 1/3



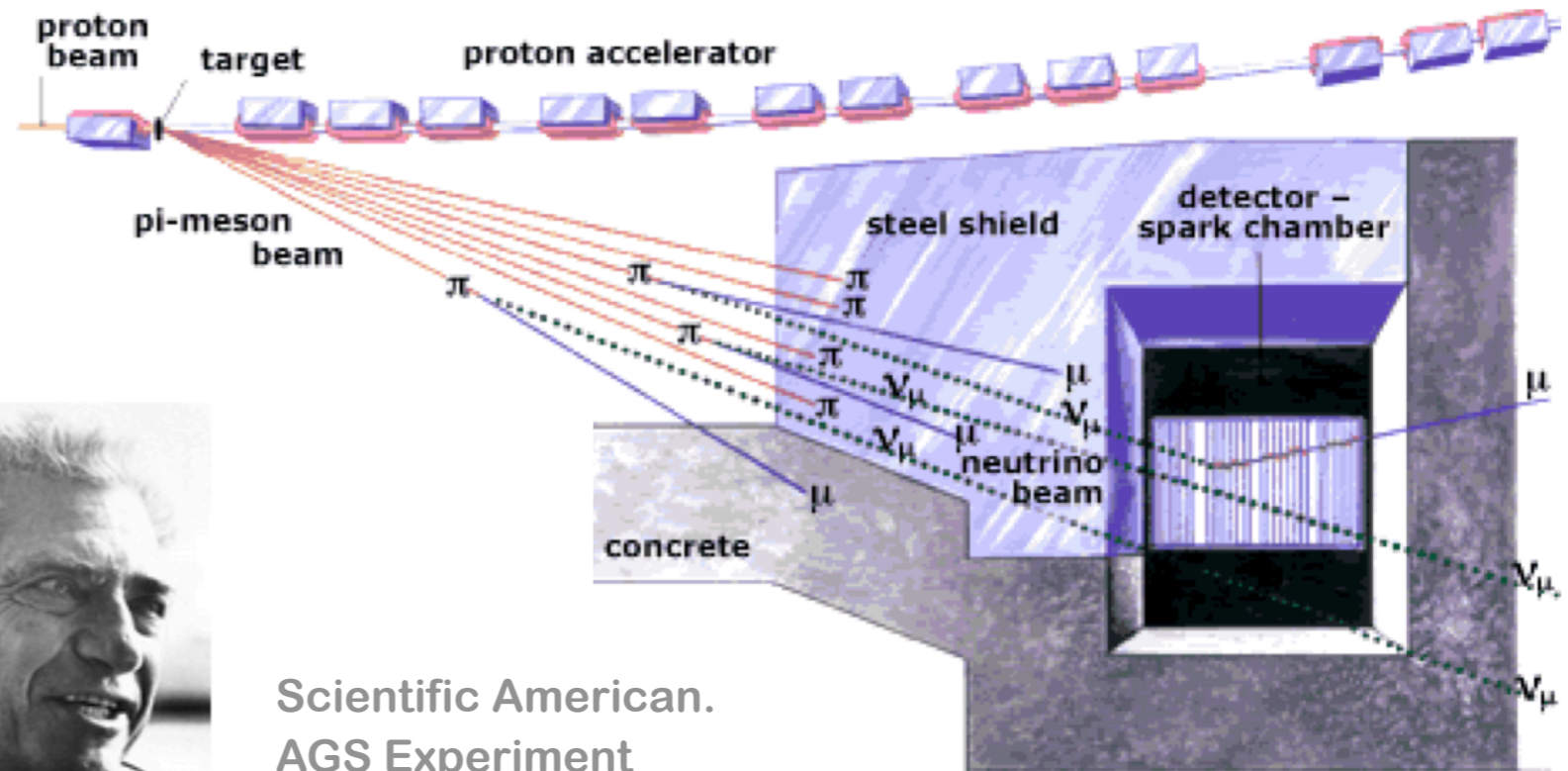
Melvin Schwartz  
Prize share: 1/3



Jack Steinberger  
Prize share: 1/3

The Nobel Prize in Physics 1988 was awarded jointly to Leon M. Lederman, Melvin Schwartz and Jack Steinberger *"for the neutrino beam method and the demonstration of the doublet structure of the leptons through the discovery of the muon neutrino"*.

Beam (tamed) neutrinos  
Muon neutrinos  $\neq$  electron neutrinos





# DETECTION OF THE COSMIC NEUTRINOS



Raymond Davis Jr.  
Prize share: 1/4

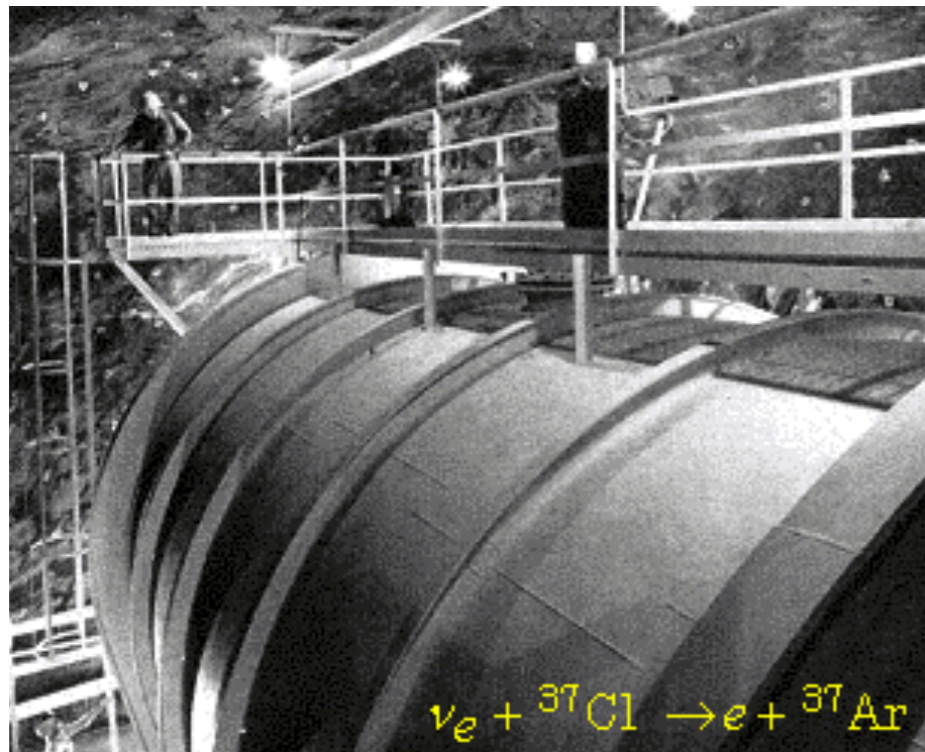


Masatoshi Koshiba  
Prize share: 1/4

- **Nobel Prize in Physics 2002:**

Pioneering contributions to astrophysics, in particular for the detection of cosmic neutrinos

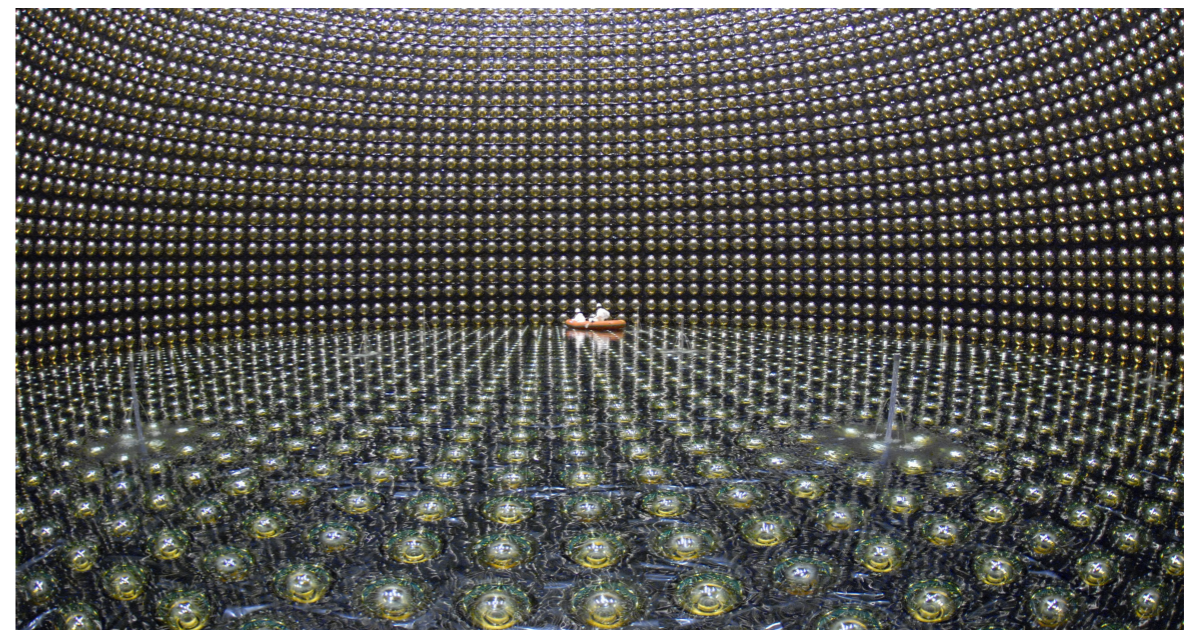
Radiochemical detection of solar- $\nu$



Mitglied der Helmholtz-Gemeinschaft

Zara Bagdasarian - Neutrino hunters around the world

Water Cherenkov detection of atmospheric  $\nu$

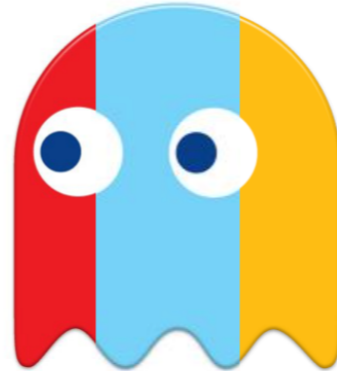


(Homestake) (Super-Kamiokande)

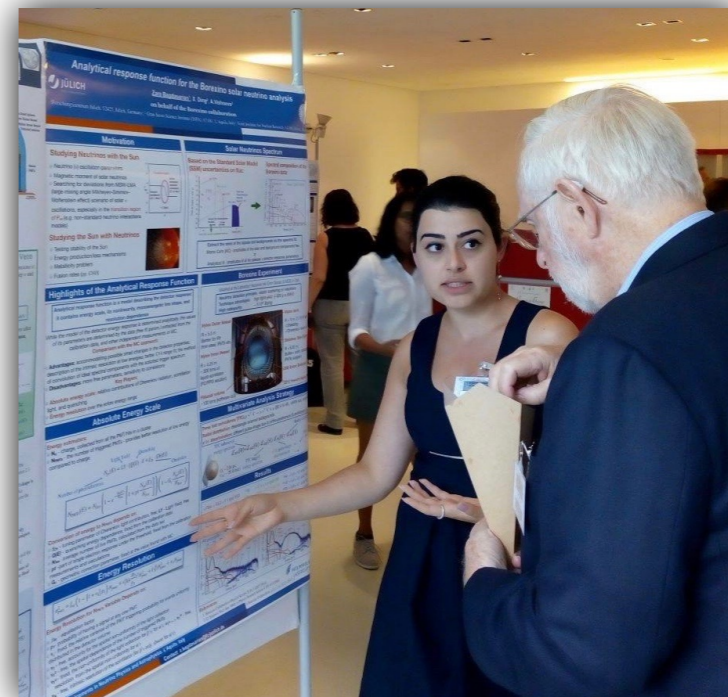
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# DISCOVERY OF NEUTRINO OSCILLATIONS



- **Nobel Prize in Physics 2015:**  
Discovery of neutrino oscillations



**Arthur McDonald  
(SNO)**

Credit: Lindau Nobel Laureate Meeting

**Takaaki Kajita (Super-K)**

Mitglied der Helmholtz-Gemeinschaft

**Zara Bagdasarian - Neutrino hunters around the world**

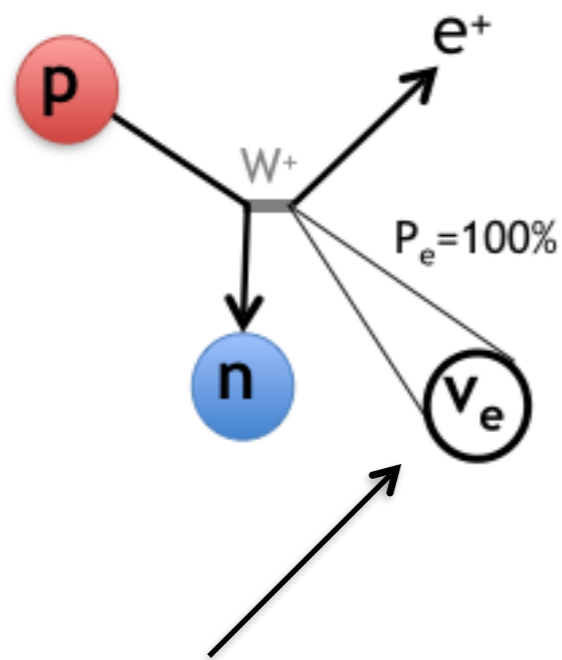
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# NEUTRINO OSCILLATIONS



## $\nu$ production

e.g.  $\beta^+$ -decay



Weak interaction  
creates neutrino in  
flavor eigenstate.

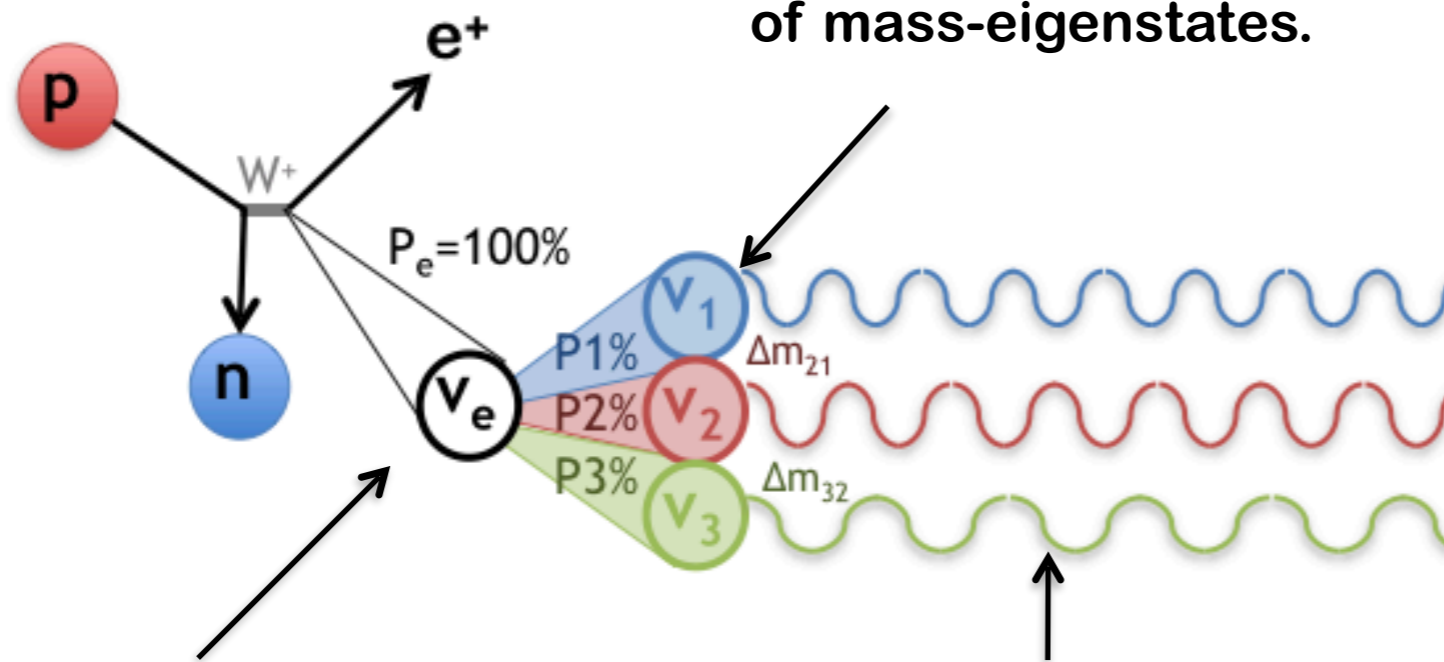


# NEUTRINO OSCILLATIONS



## $\nu$ production

e.g.  $\beta^+$ -decay



## $\nu$ propagation

as coherent superposition of mass-eigenstates.

Weak interaction creates neutrino in flavor eigenstate.

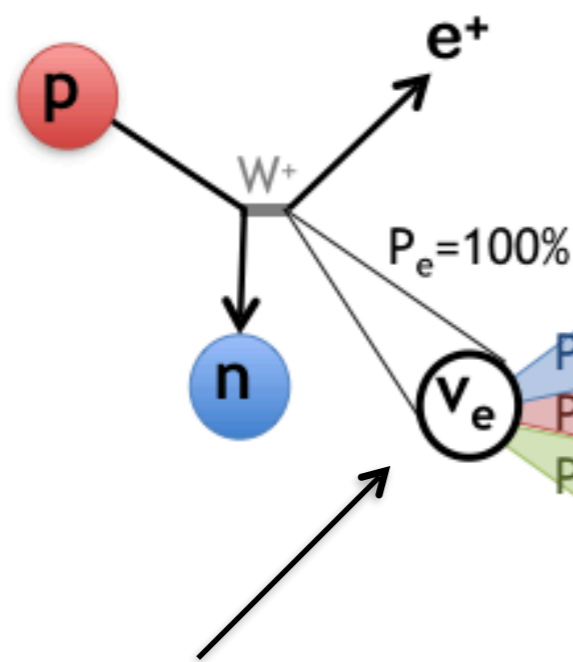
Different masses create a phase difference over time.

# NEUTRINO OSCILLATIONS



## $\nu$ production

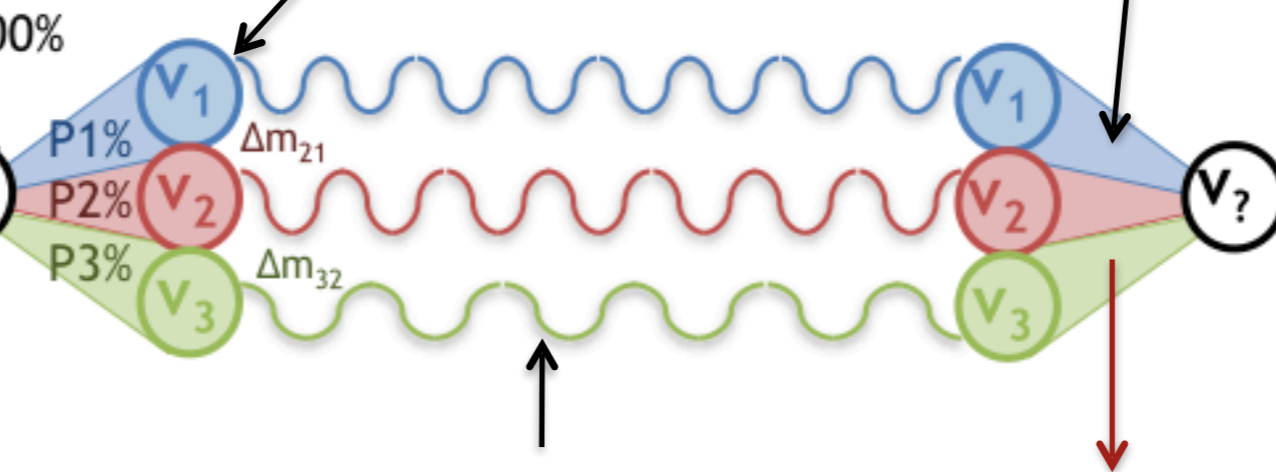
e.g.  $\beta^+$ -decay



Weak interaction creates neutrino in flavor eigenstate.

## $\nu$ propagation

as coherent superposition of mass-eigenstates.



Different masses create a phase difference over time.

## $\nu$ detection

as flavor eigenstate: superposition of mass eigenstates has changed because of the phase factors.

$$P = P_e\% : \nu_e$$

$$P_\mu\% : \nu_\mu$$

$$P_\tau\% : \nu_\tau$$

Finite probability to detect a different neutrino flavor!

$$P(\nu_e \rightarrow \nu_e) = 1 - \sin^2 2\theta_{12} \sin^2 \frac{1.27 \Delta m_{12}^2 L}{E}$$

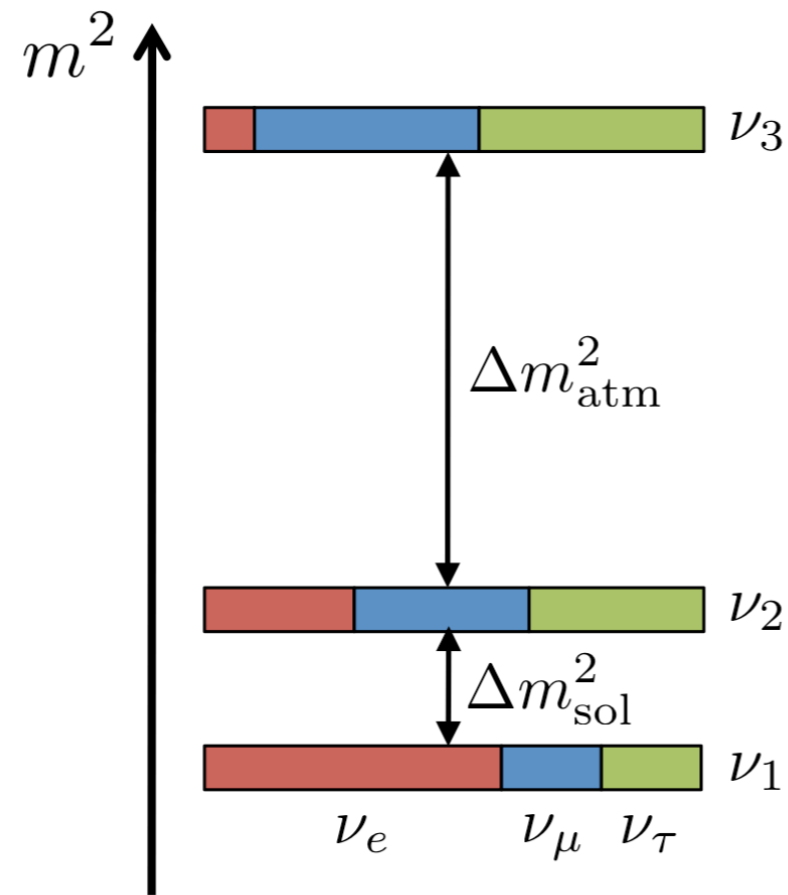


# IT'S ALL ABOUT THAT MASS

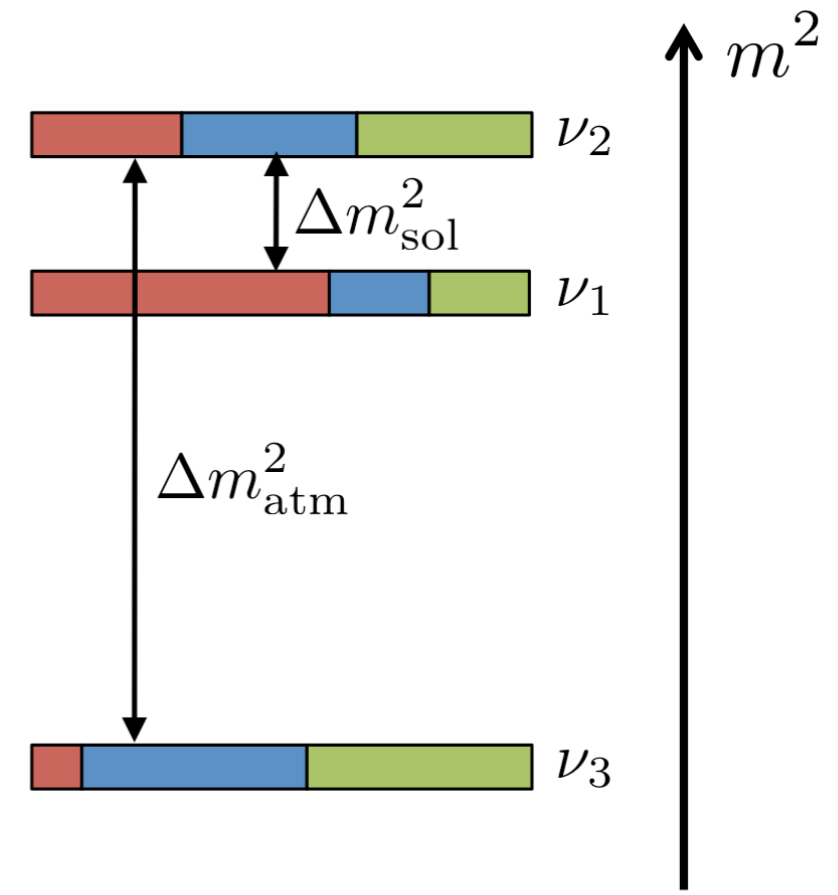
## Mass ordering



## Normal Ordering (NO)



## Inverse Ordering (IO)

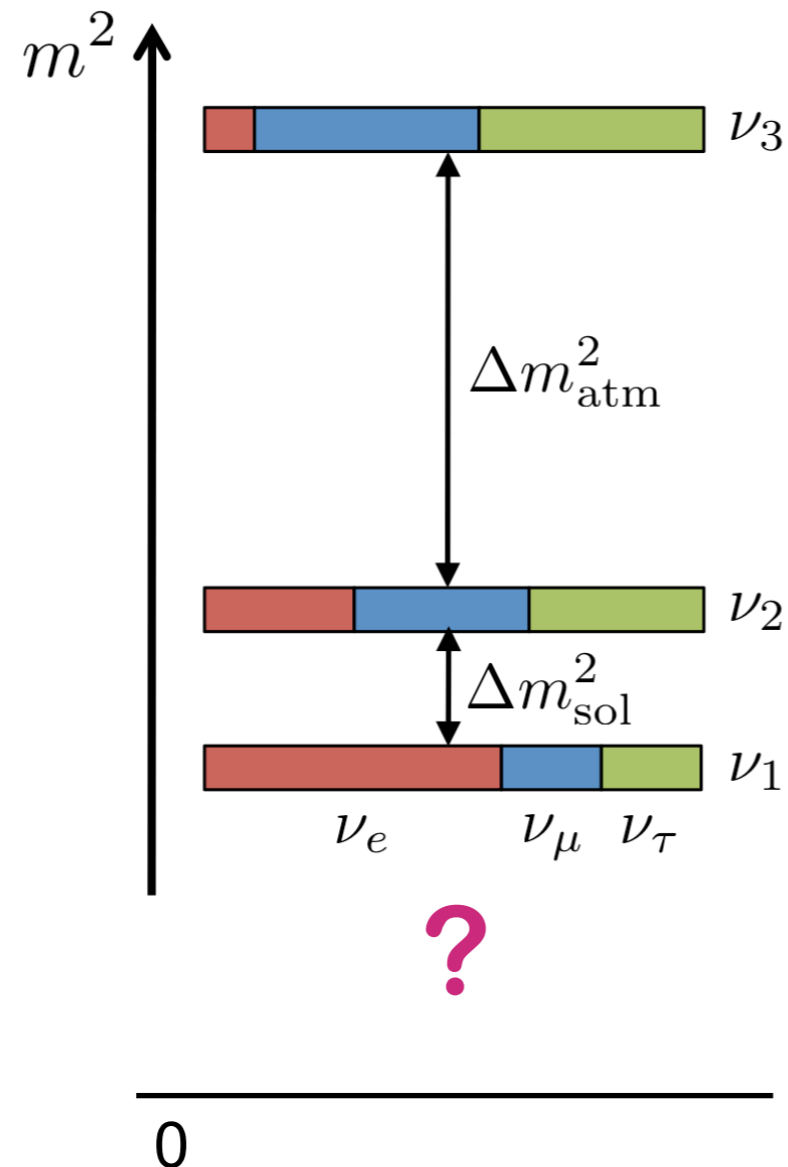


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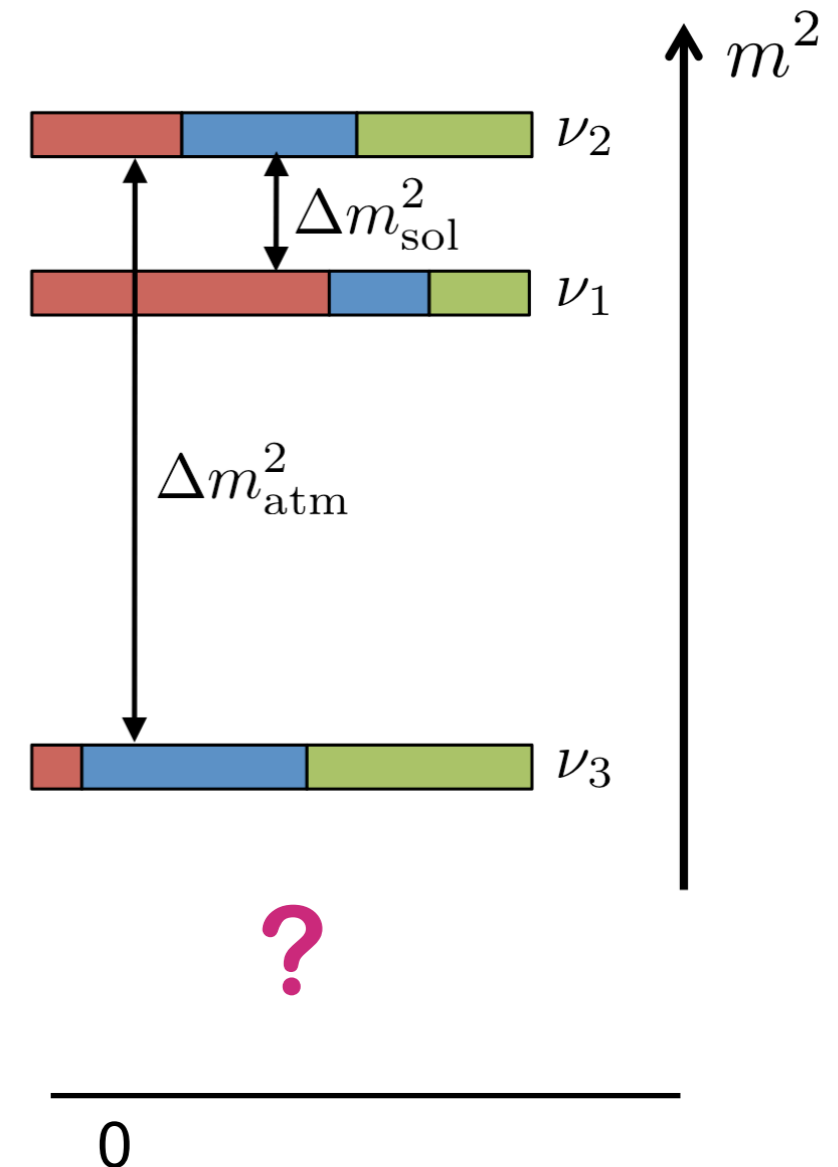
Mass ordering  
Absolute Mass



## Normal Ordering (NO)



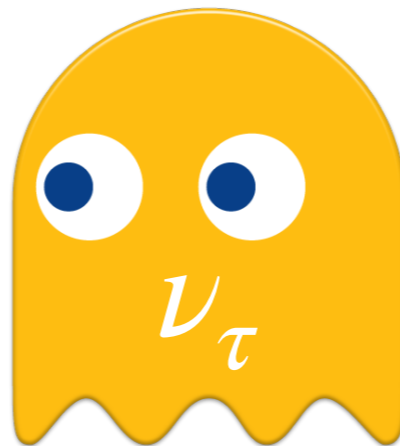
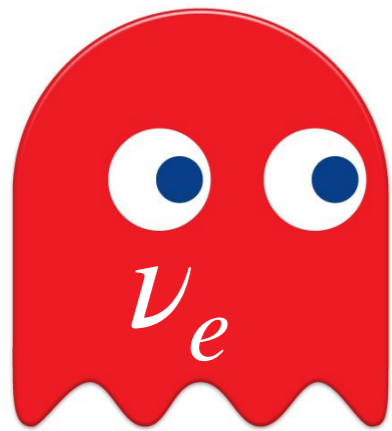
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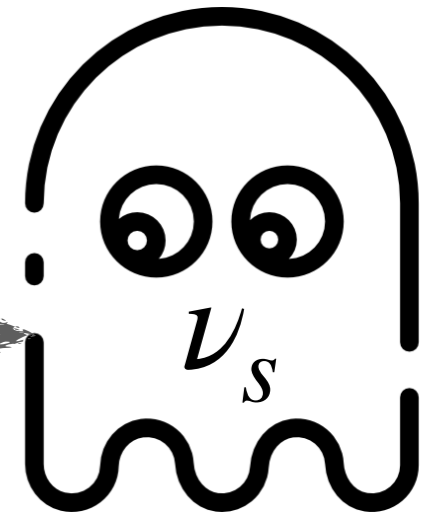


# MORE THAN THREE?

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{14} \sin^2 \frac{1.27 \Delta m_{14}^2 L}{E}$$



Can I come play  
with you?



# DIRAC OR MAJORANA

$$\nu \neq \bar{\nu}$$

**Dirac**

**vs.**

$$\nu = \bar{\nu}$$

**Majorana**



**Mechanism for matter-antimatter asymmetry in the universe?**



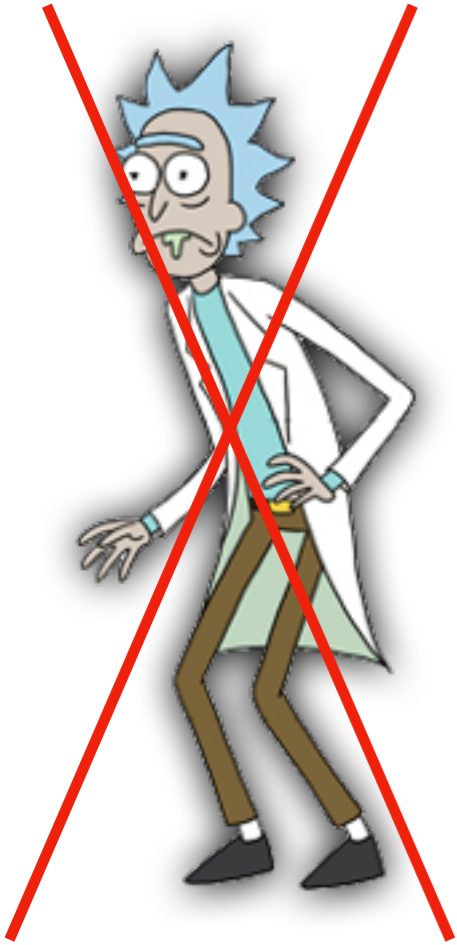
# WHO ARE THE NEUTRINO HUNTERS?



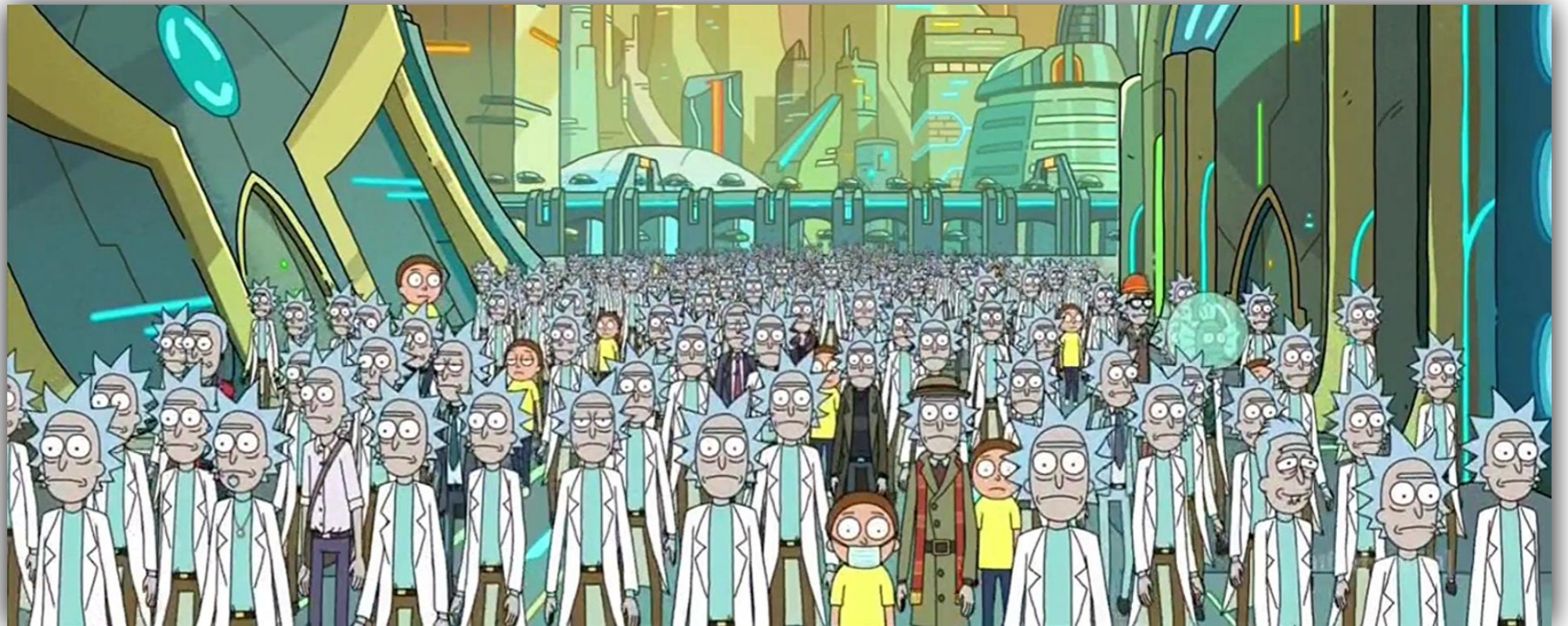
- A lonely grey-haired crazy scientist?

Source: Rick and Morty

# WHO ARE THE NEUTRINO HUNTERS?



- A lonely grey-haired crazy scientist?
- Maybe crazy, but not alone, and not only grey-haired



Source: Rick and Morty



# WHO ARE THE NEUTRINO HUNTERS?



Recent photo from the biggest neutrino conference (Heidelberg, Germany)

Source: <https://www.mpi-hd.mpg.de/nu2018/>



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Source: <https://www.mpi-hd.mpg.de/nu2018/>



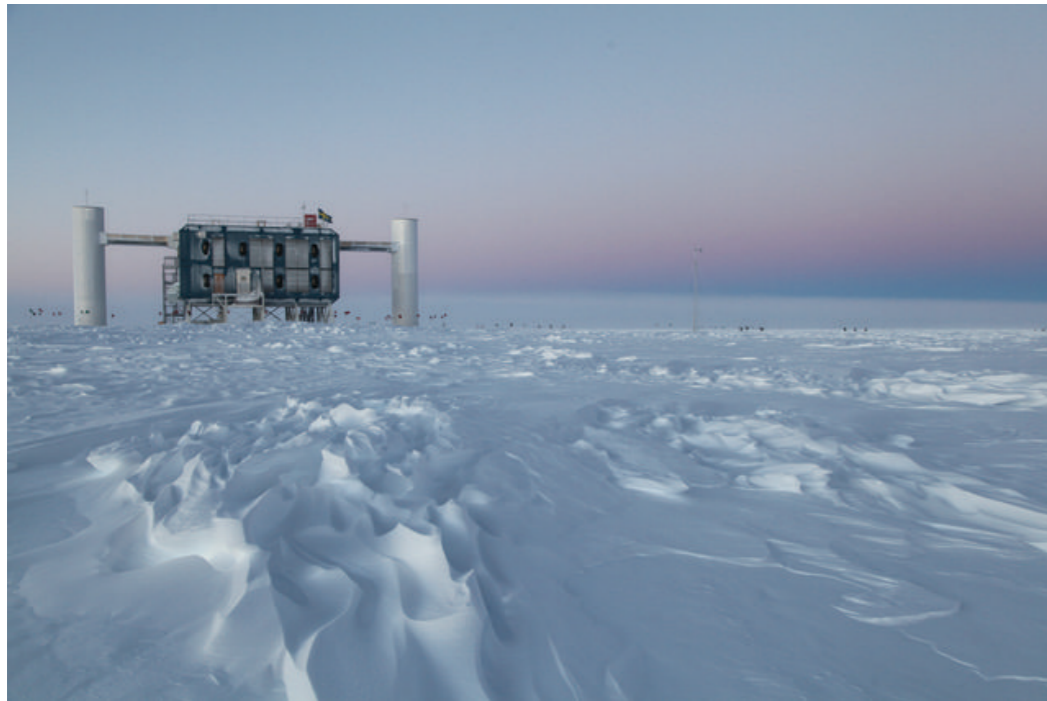
# WHAT ARE NEUTRINO HUNTERS DOING?

Build the laboratories

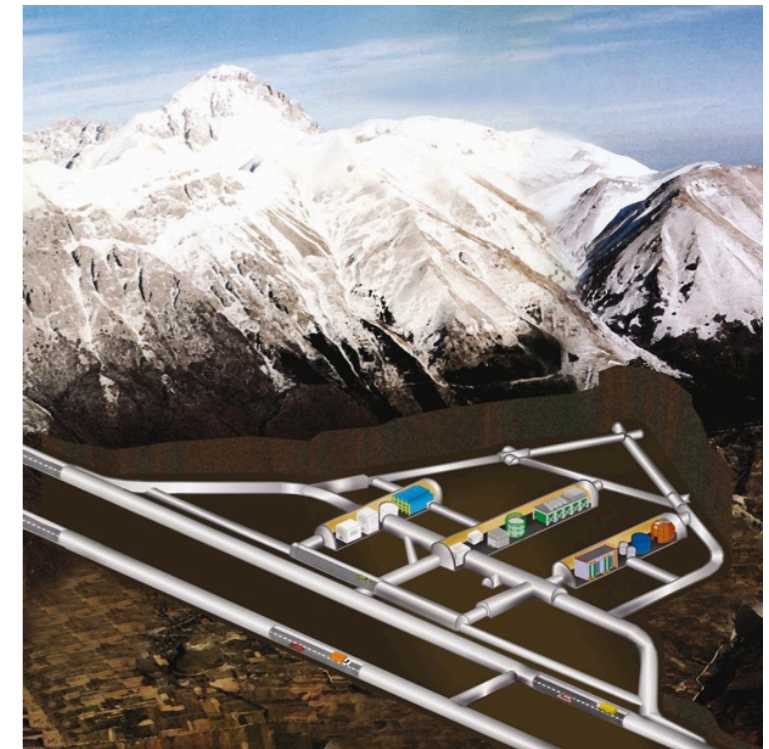
- In the mines
- In the mountains
- In the ice



Credit:SNOLAB



Credit: IceCube/NSF



Credit: Laboratori Nazionali del Gran Sasso



# WHAT ARE NEUTRINO HUNTERS DOING?

Build the laboratories

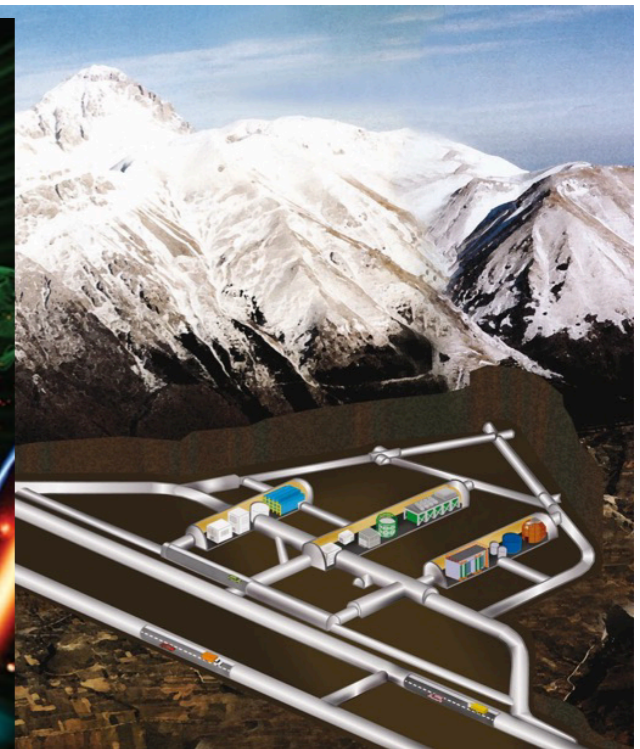
- In the mines
- In the mountains
- In the ice



Search for the ghosts!  
Solve mysteries



Credit: IceCube/NSF



Laboratori Nazionali del Gran Sasso



# LABORATORI NAZIONALI DEL GRAN SASSO (LNGS)



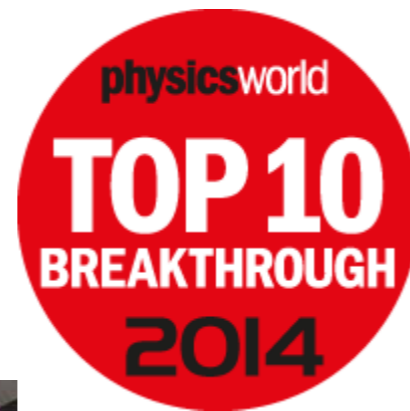
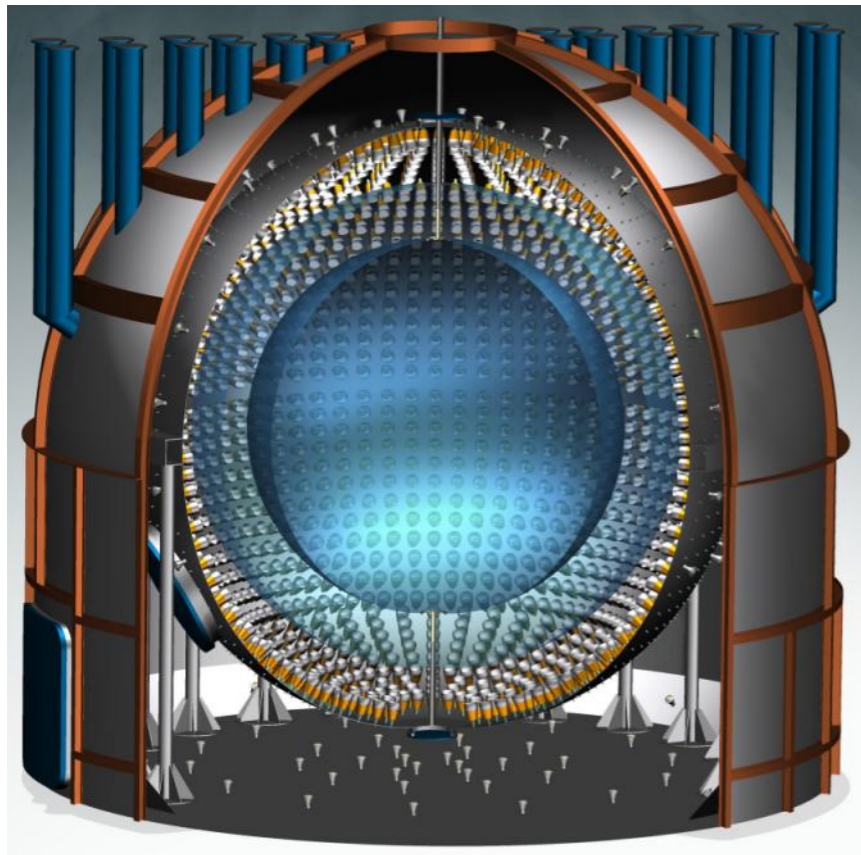
Credit: Laboratori Nazionali del Gran Sasso



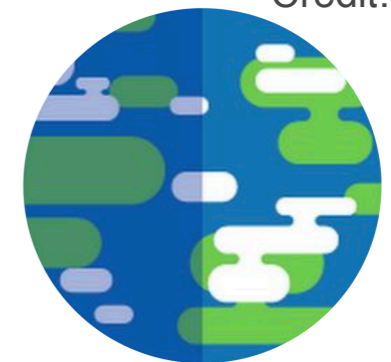
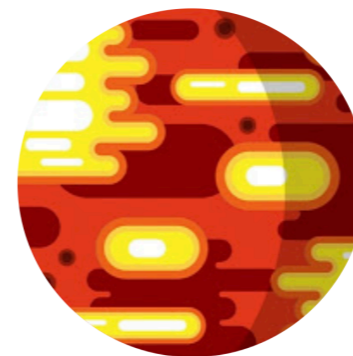
Credit: Zara Bagdasarian



# BOREXINO @ LNGS



Credit: Borexino



The most radiopure spot in the world

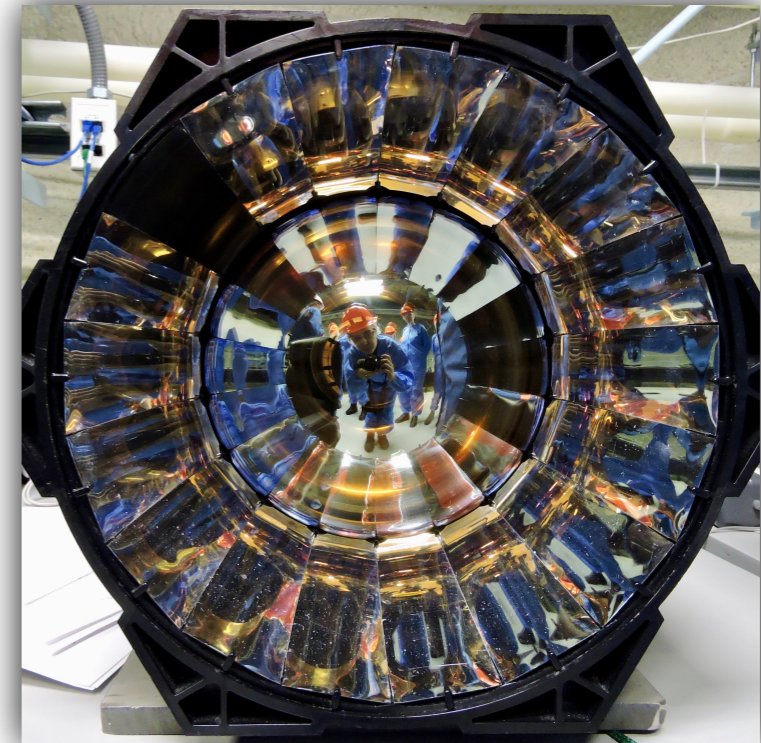
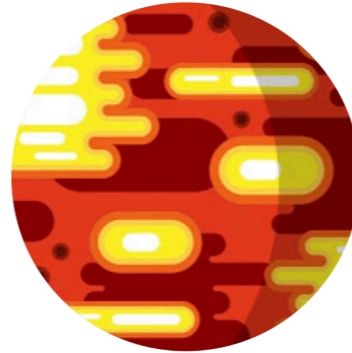




# SUDBURY NATIONAL OBSERVATORY



- 2 km down the mine
- Solved solar neutrino problem

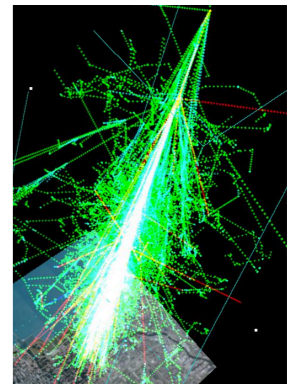
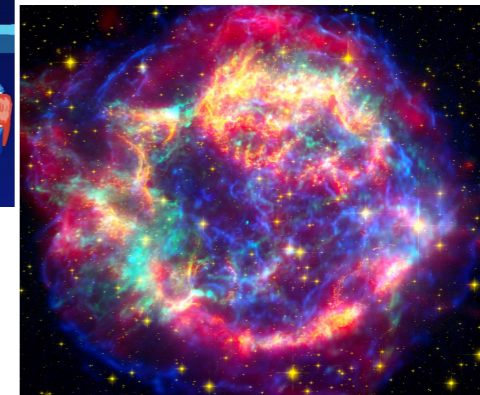
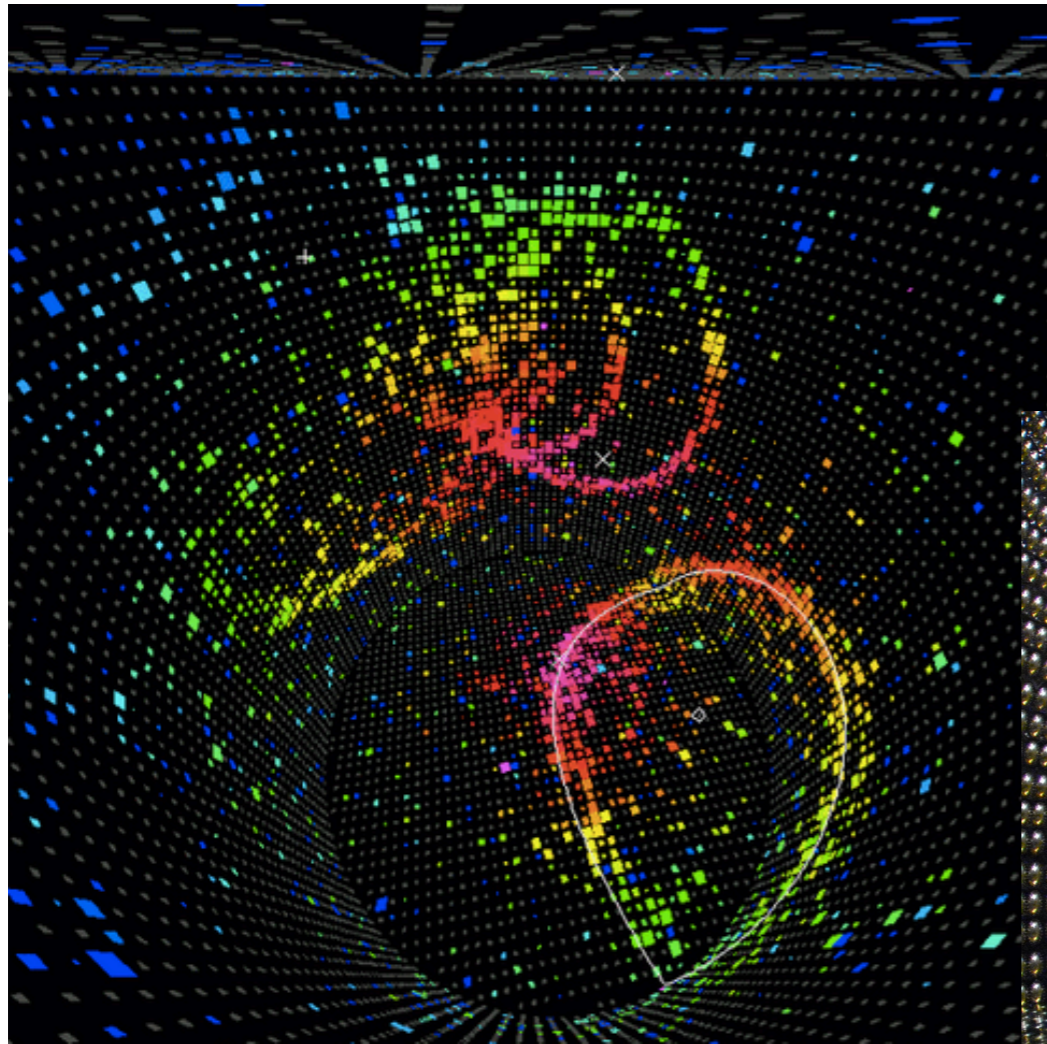


Credit: Zara Bagdasarian

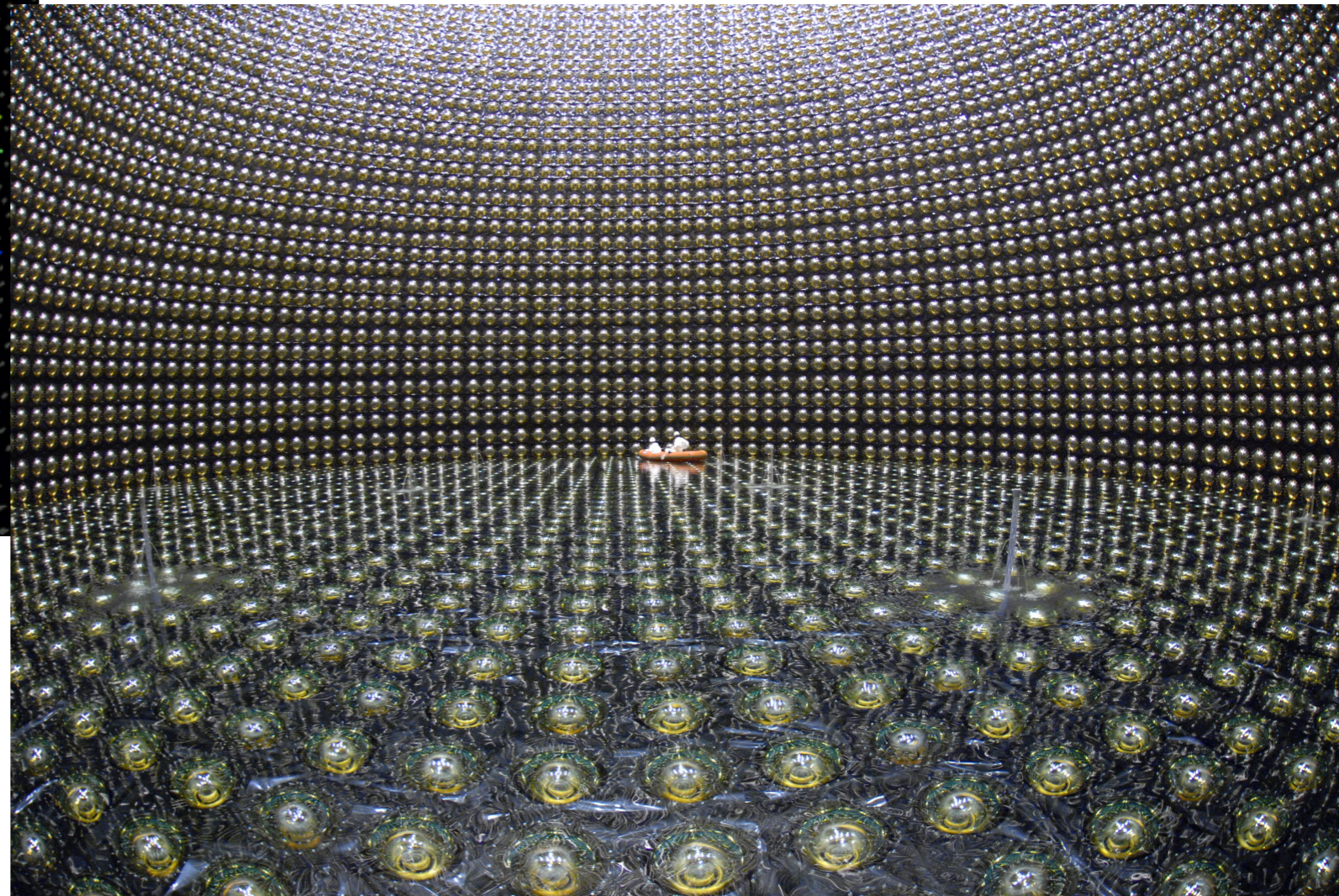




# SUPER-KAMIOKANDE



a massive tank containing  
50,000 tonnes of ultra pure  
water:  
over 33 Olympic sized  
swimming pools!

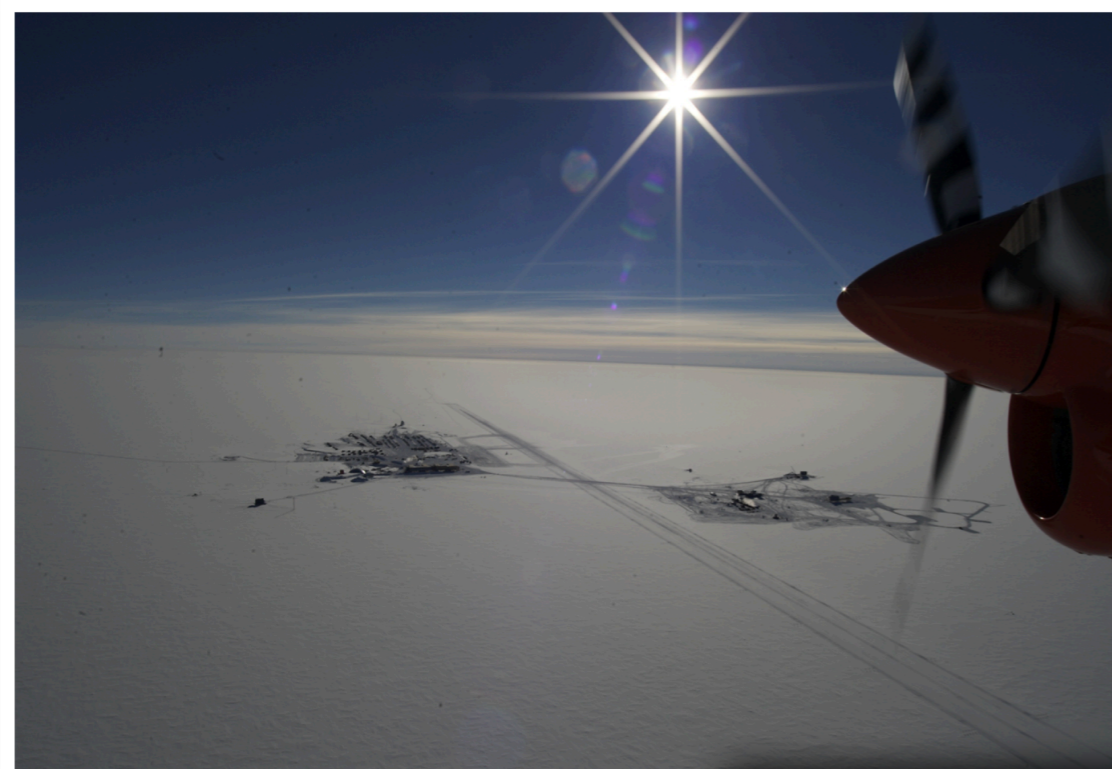


Future: from Super to Hyper

Credit: Super-Kamiokande

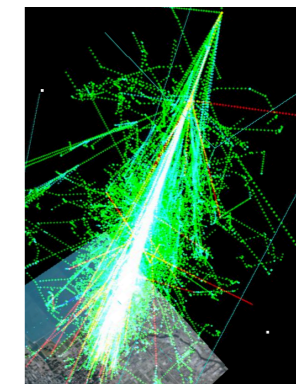
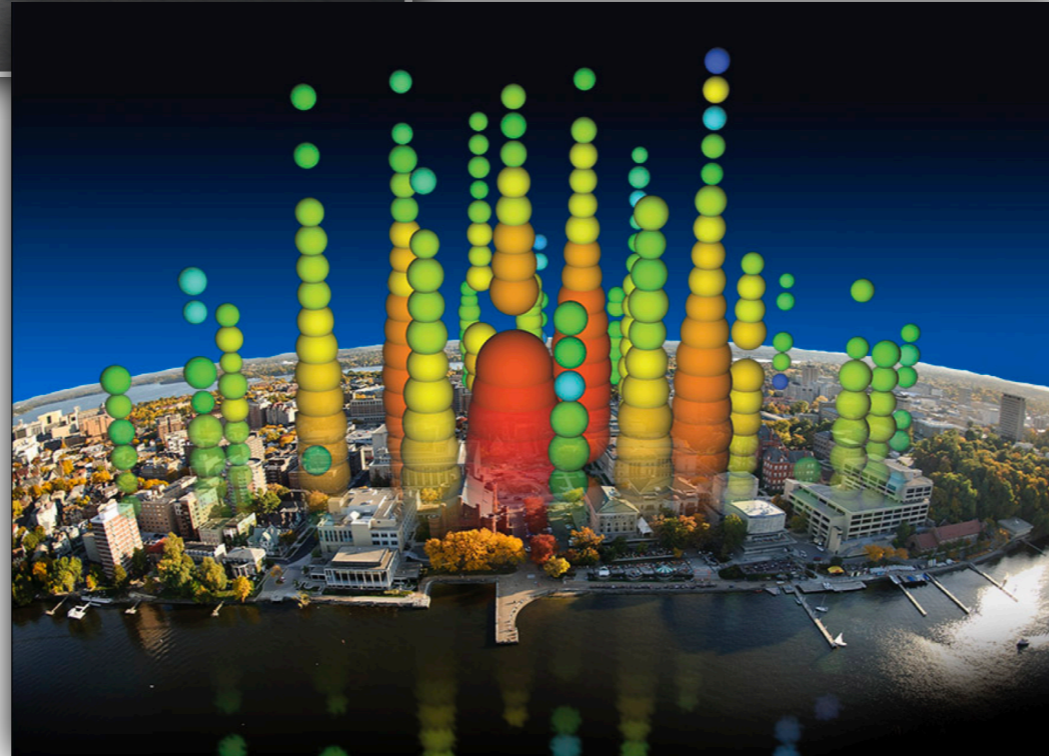
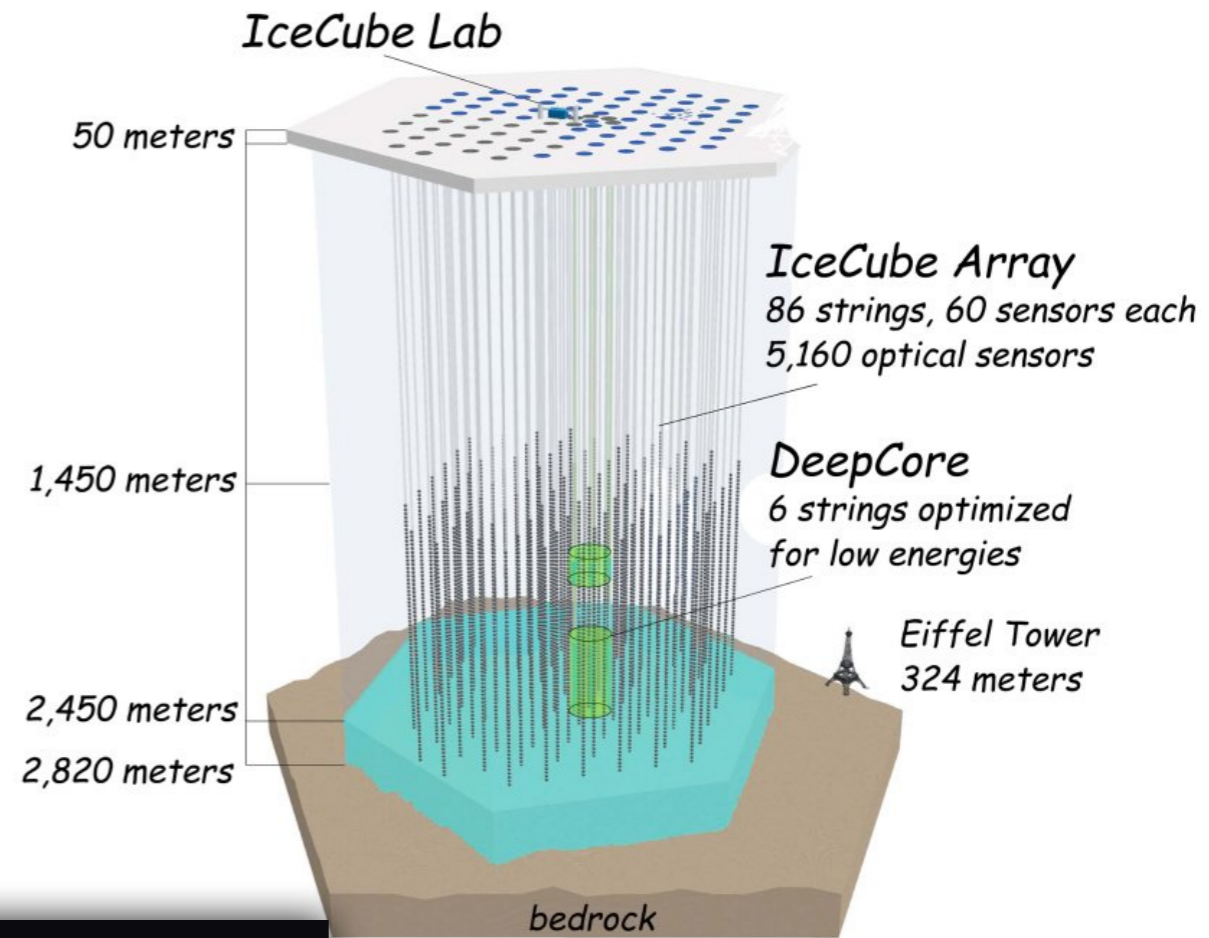


# ICE CUBE



ultra-transparent ice below 1.5 km

Credit: IceCube

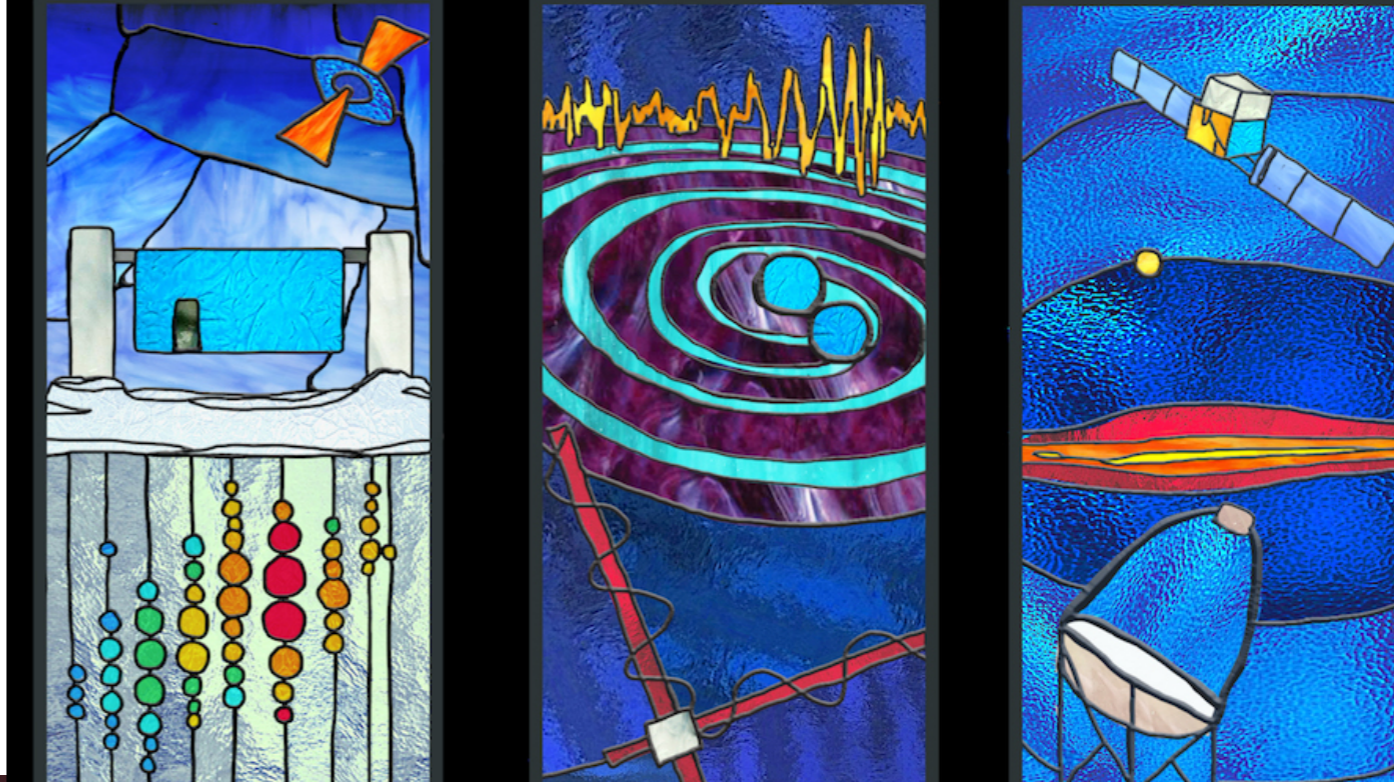




# ICE CUBE



Credit: IceCube/Science



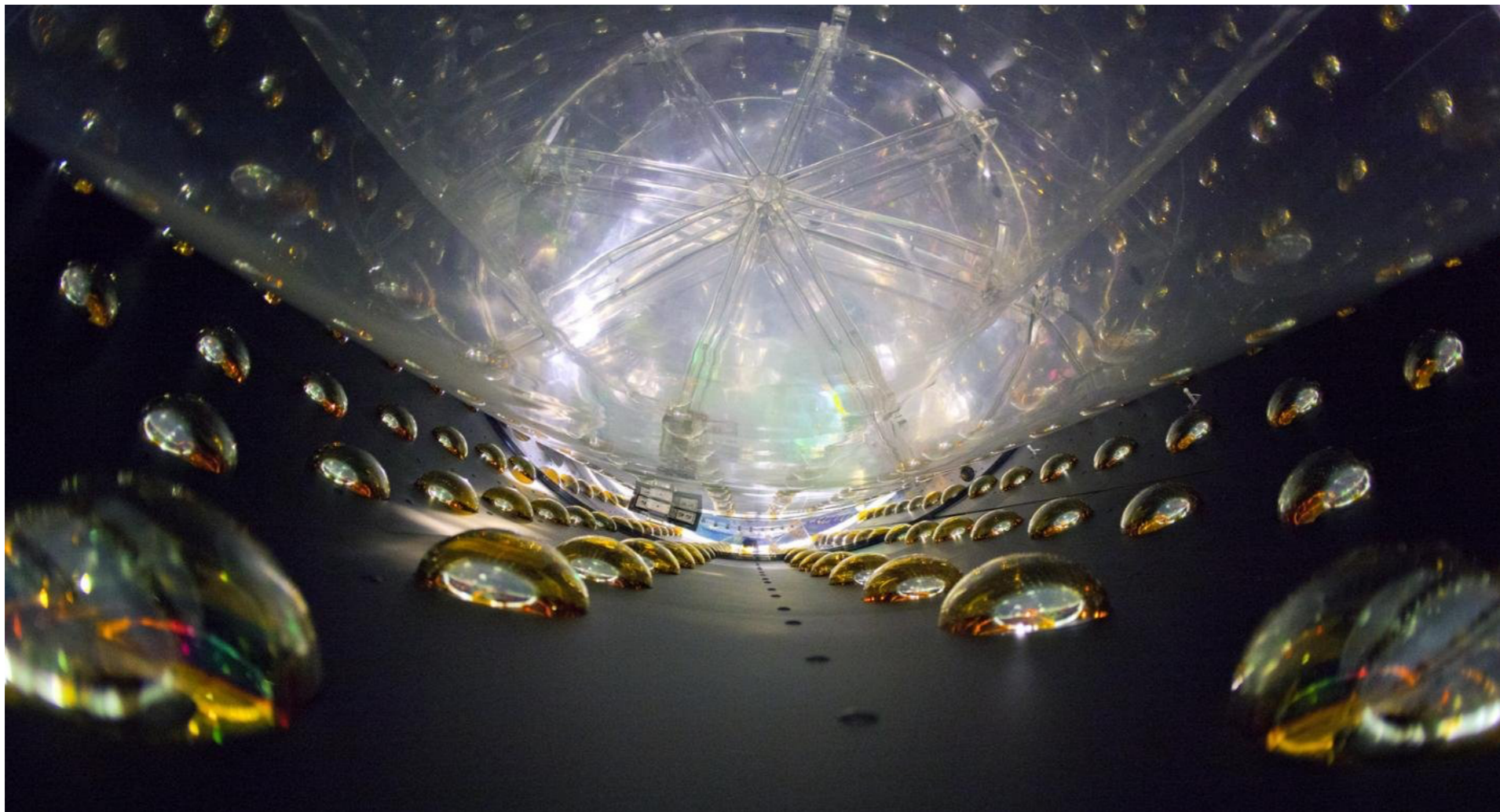
Credit: IceCube





# DAYA BAY

- Reactor antineutrinos: record 4 million antineutrinos dataset



$$\sin^2 2\theta_{13} = 0.0856$$
$$\Delta m^2_{ee} = 2.52 \times 10^{-3} \text{ eV}^2$$

precision of  
3.4% and 2.8%

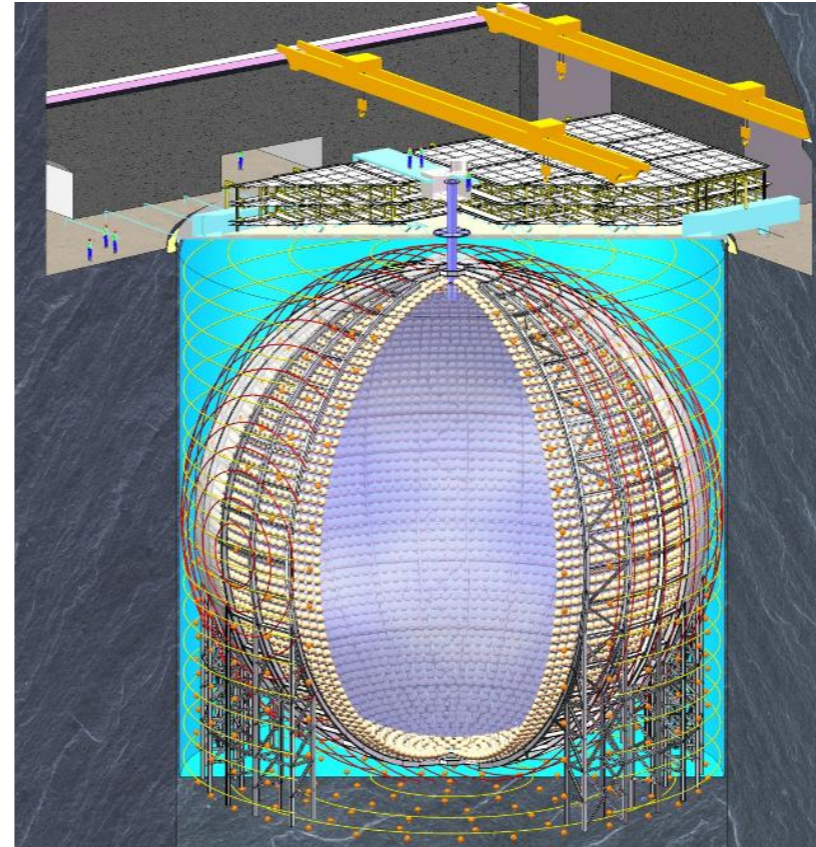
Credit: Daya Bay



# ANTINEUTRINOS FROM REACTORS



Credit: RENO



Credit: JUNO Collaboration



- Other reactor experiments:
  - DOUBLE CHOOZ (France)
  - RENO (Korea)
  - JUNO (China) (upcoming)

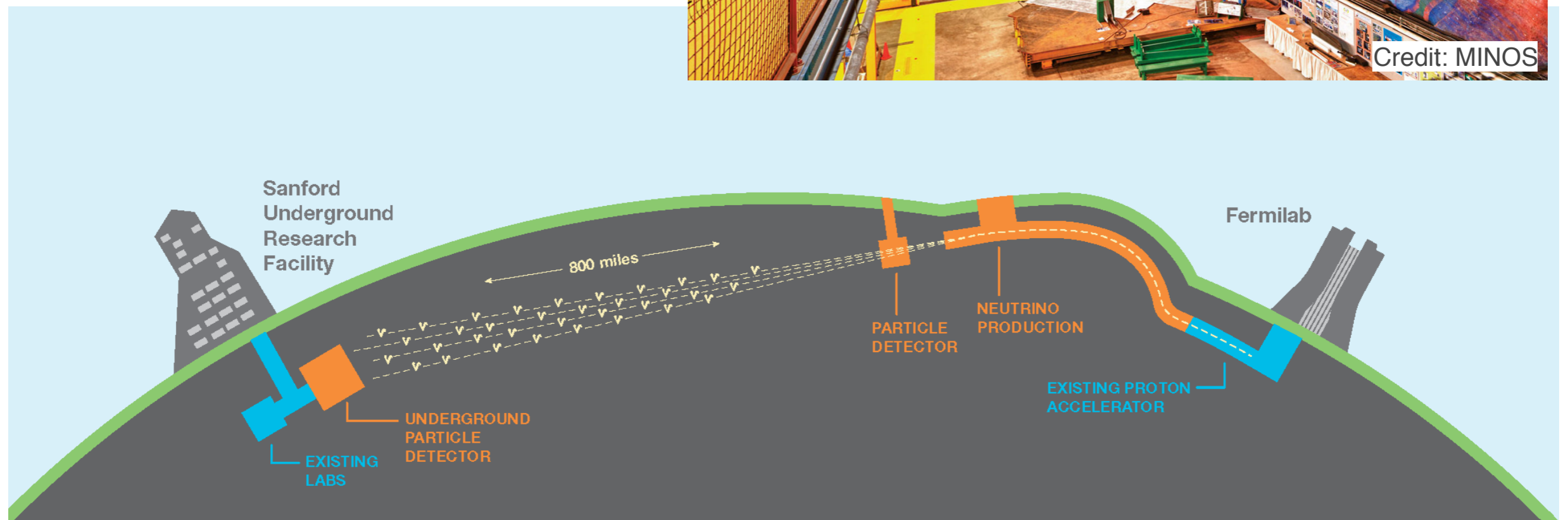


Credit: Double Chooz



# NEUTRINOS FROM MAN-MADE BEAMS

- MINOS (USA)
- NOvA(USA)
- T2K (Japan)
- DUNE (USA) (future)



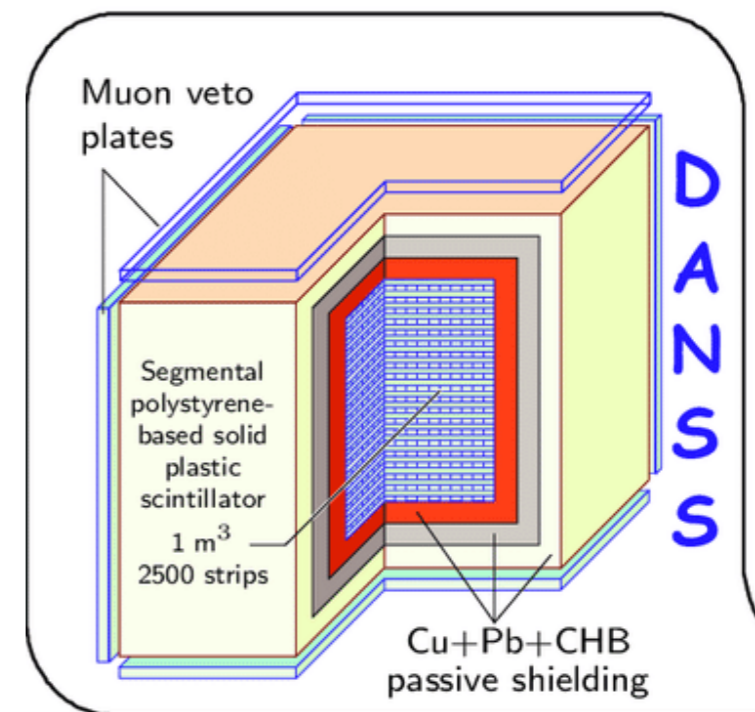
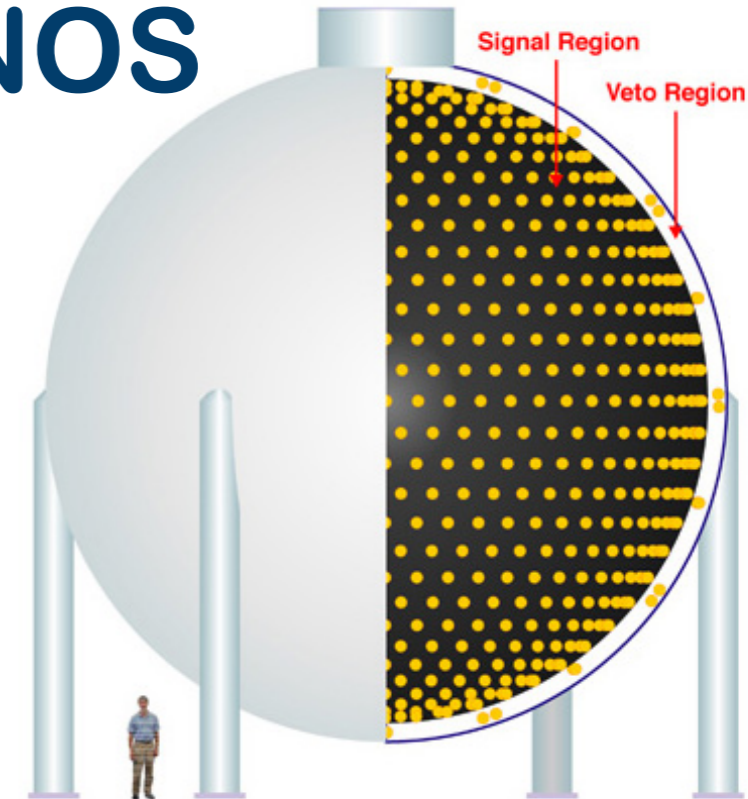
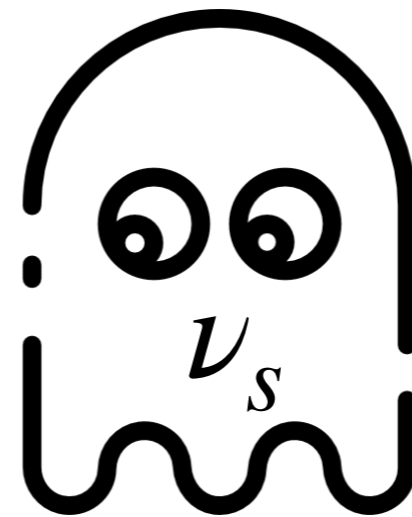
CP-violation: differences between neutrinos and antineutrinos

Credit: DUNE Collaboration



# SEARCH FOR STERILE NEUTRINOS

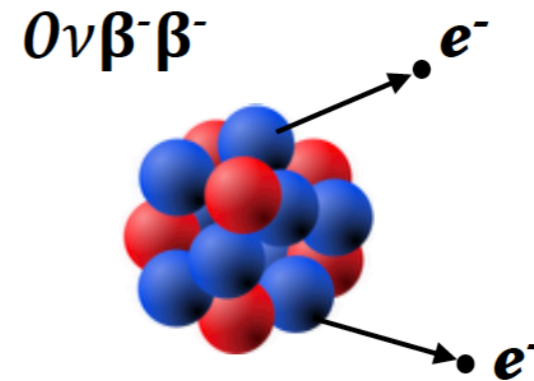
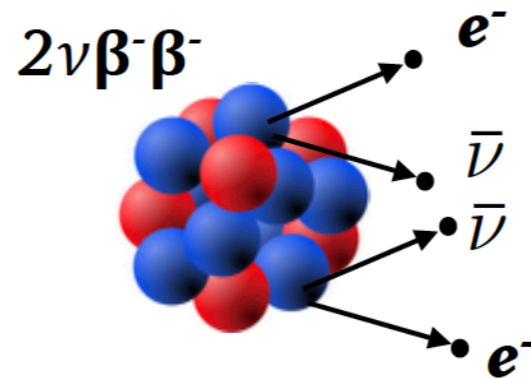
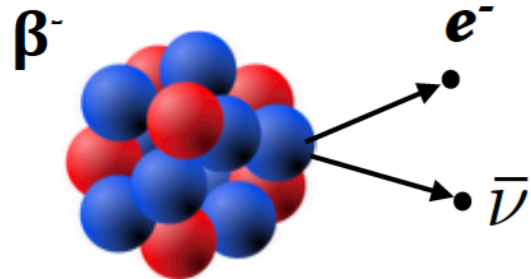
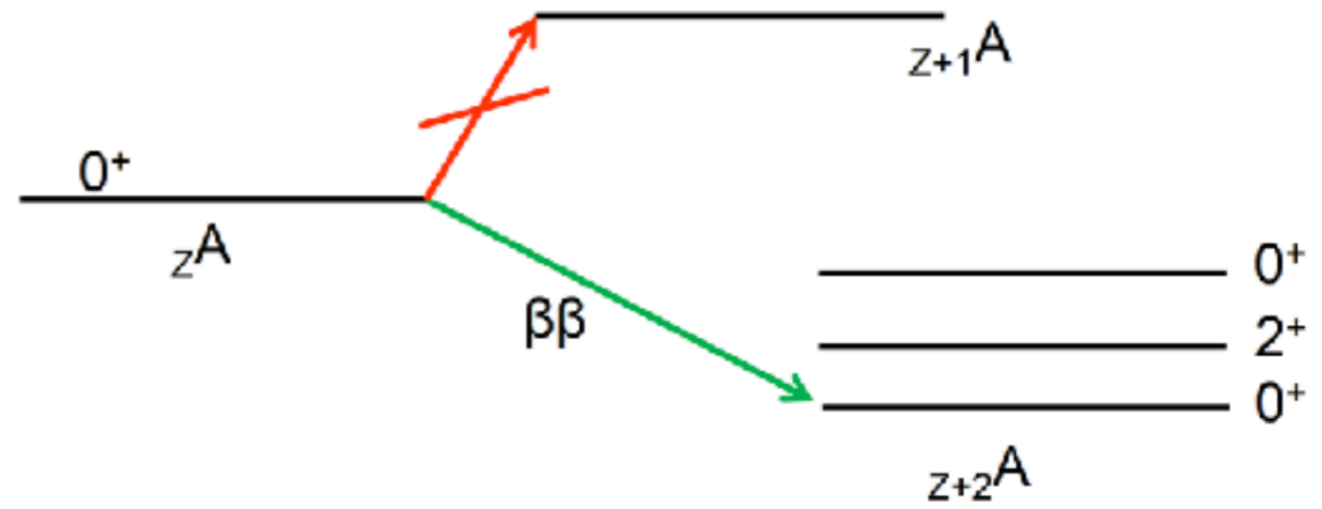
- MiniBooNE (USA) (recently in the news)
- DANSS (Russia)
- NEOS (Korea)
- PROSPECT (USA)





# SEARCHES FOR $0\nu\beta\beta$ DECAY

Elements for which normal beta decay is suppressed:  
Germanium, Xenon, Tellurium



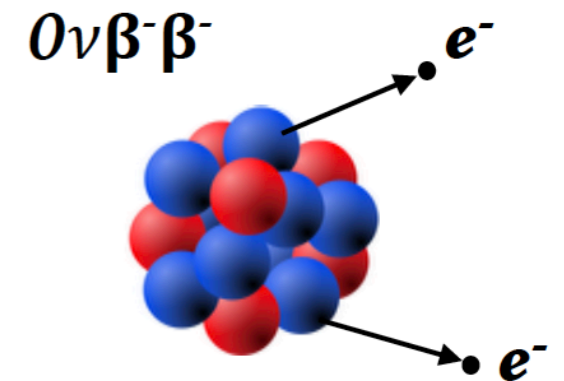
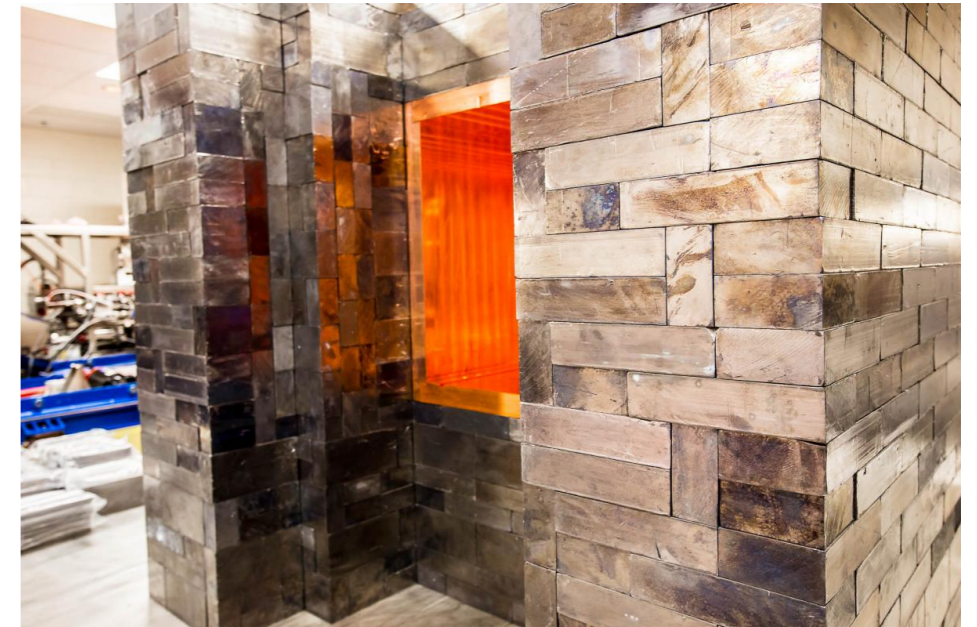
Processes within the Standard Model

Neutrinos are their own antiparticles  
Lepton number is not conserved



# SEARCHES FOR $0\nu\beta\beta$ DECAY

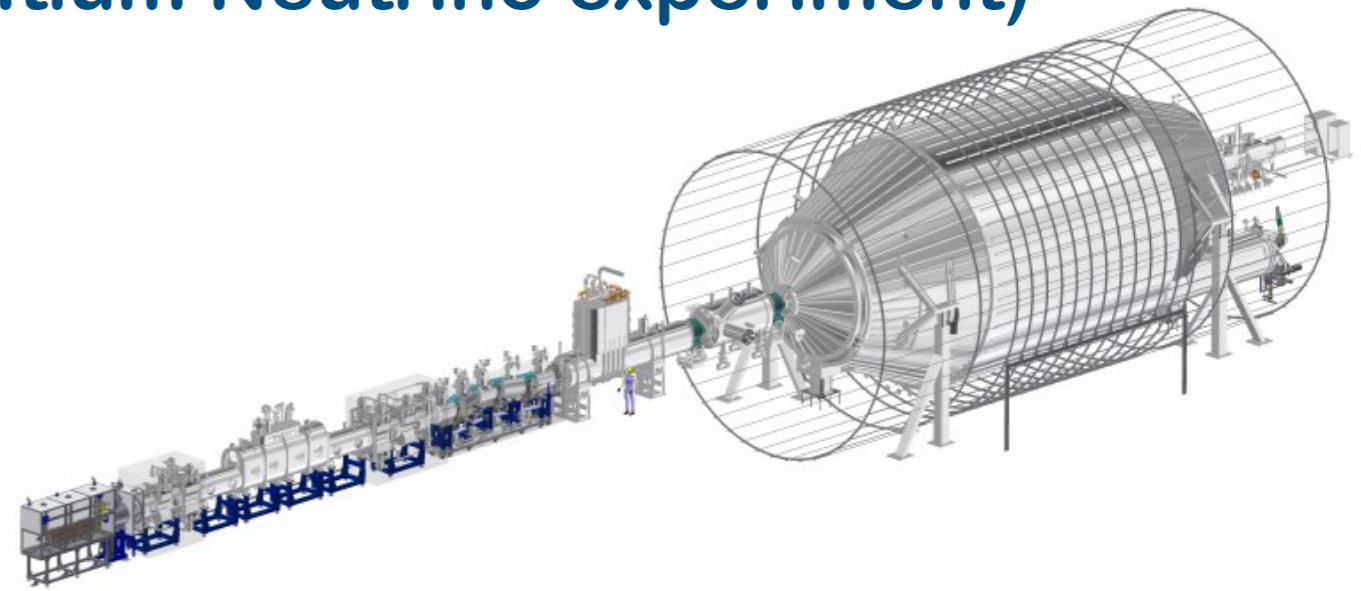
- GERDA (Italy)  $m_{\beta\beta} < (110, 260) \text{ meV}$
- MAJORANA DEMONSTRATOR (USA)
- CUORE (Italy)



- EXO (USA)
- KAMLand- Zen (Japan)



# KATRIN (the Karlsruhe TRitium Neutrino experiment)

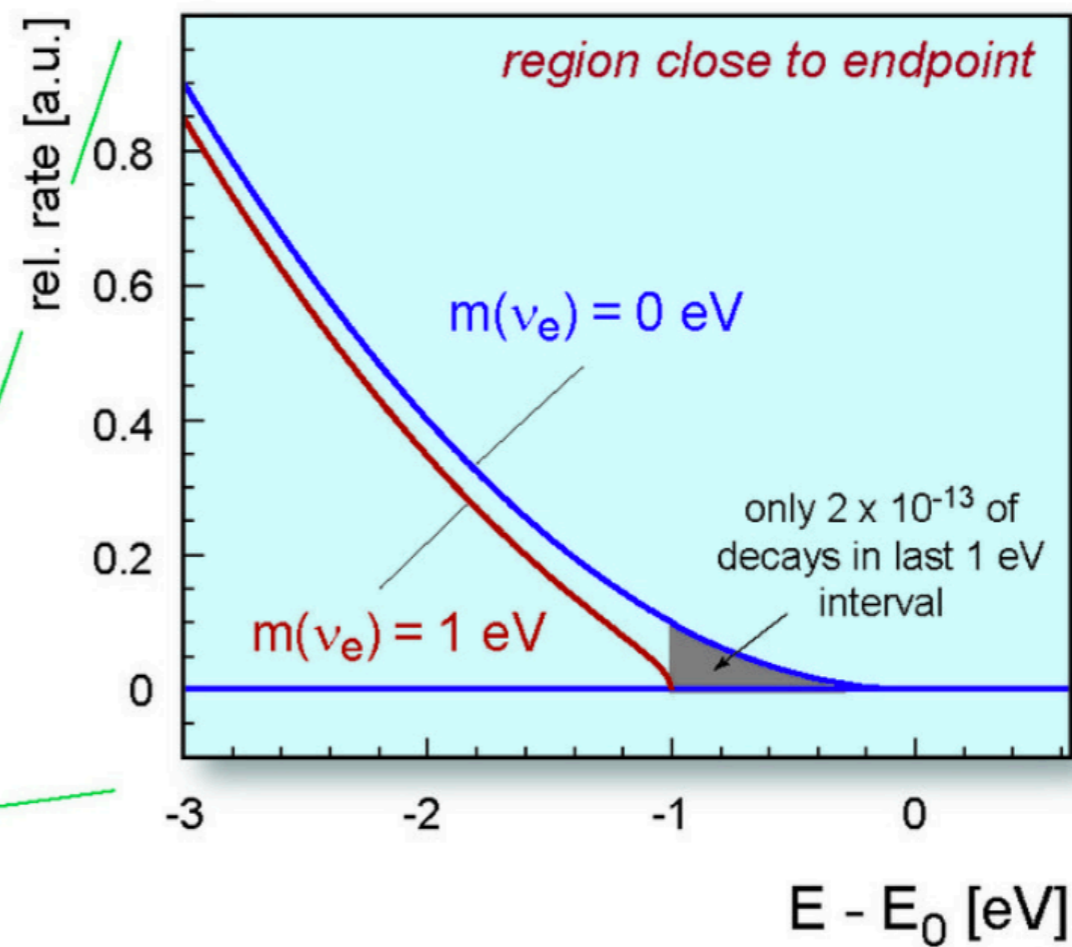
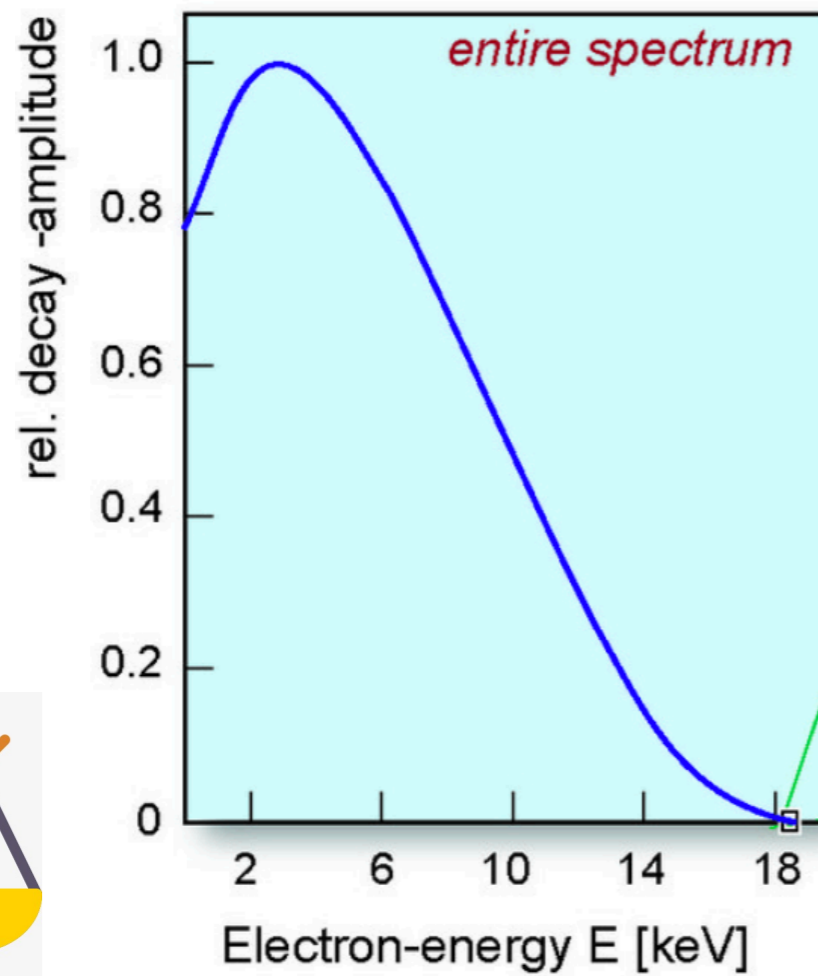
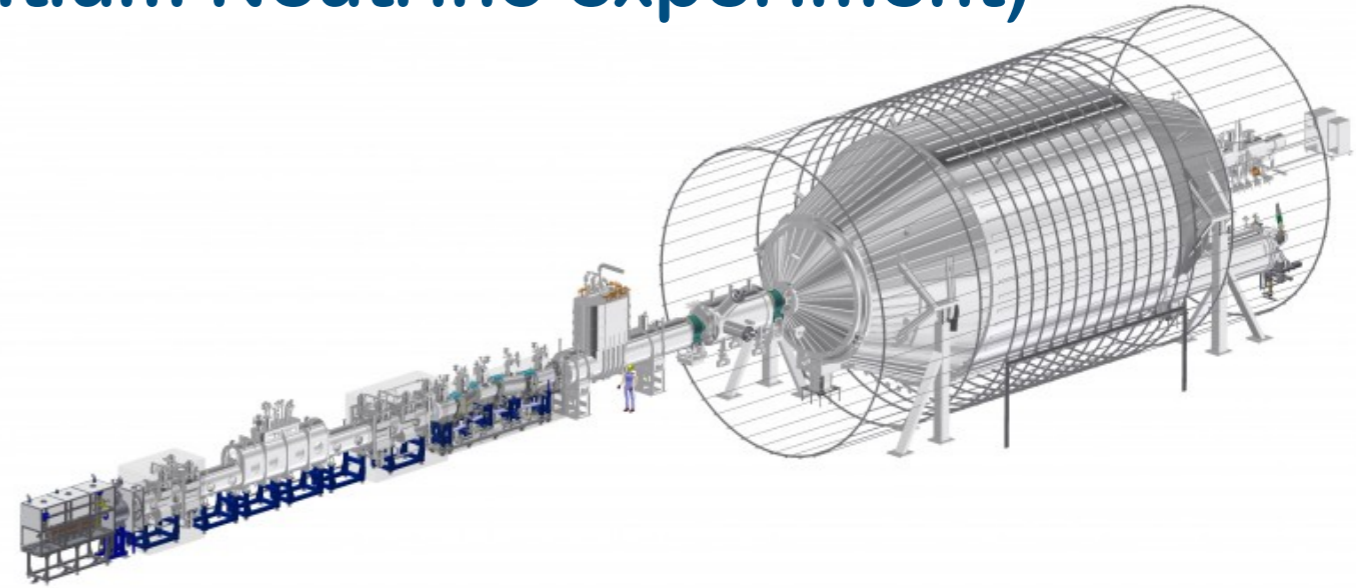
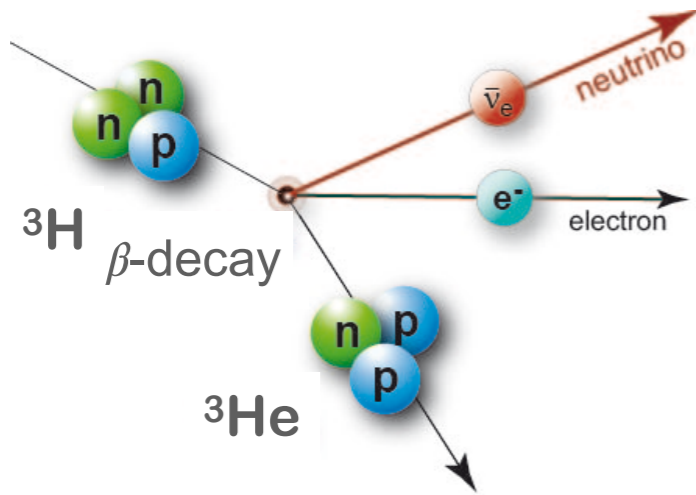


**Goal: Absolute mass scale of neutrinos**



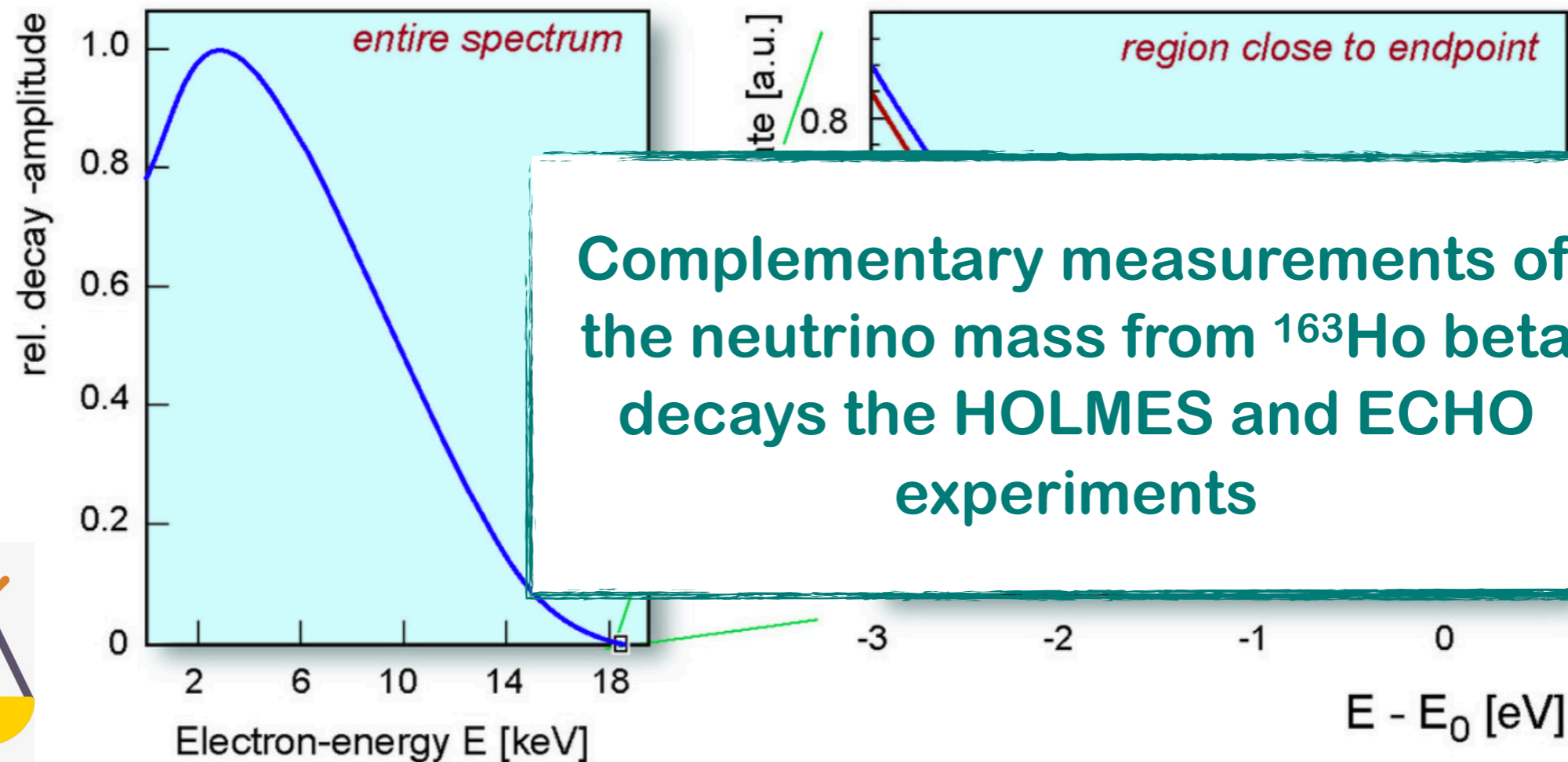
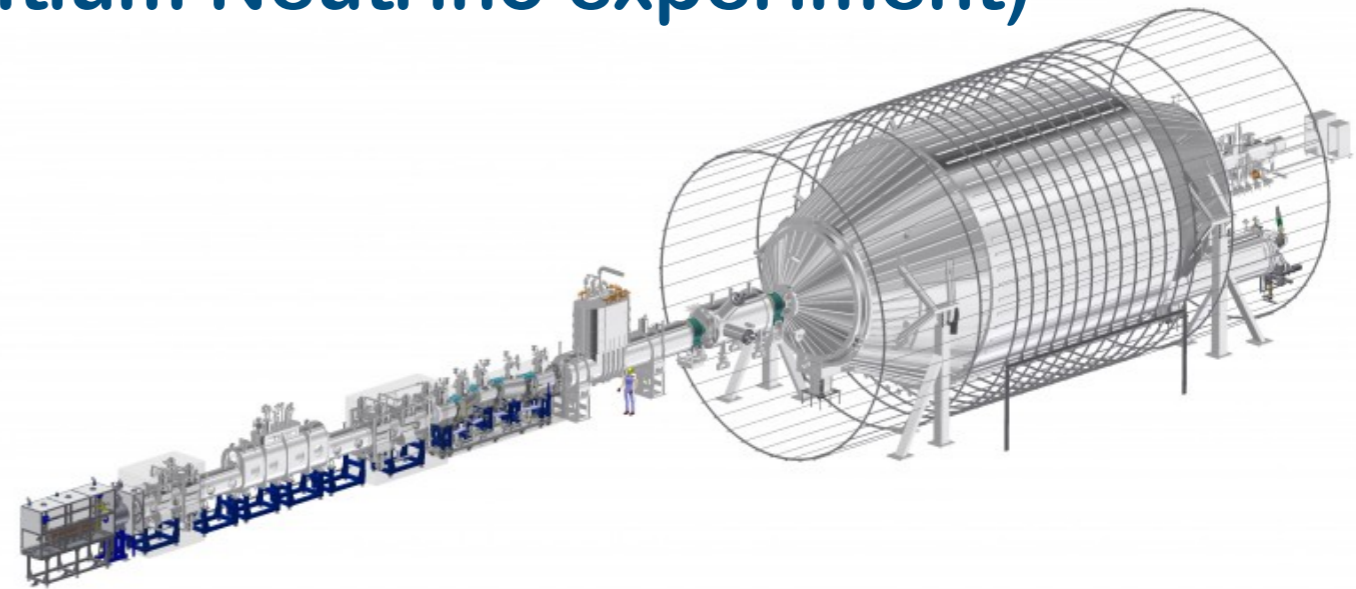
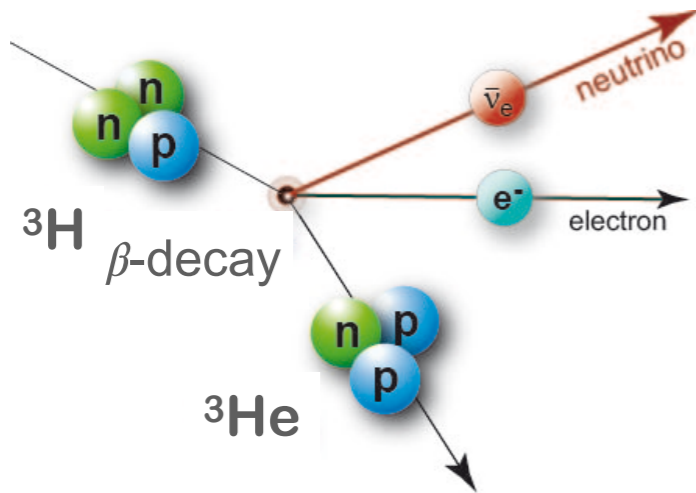


# KATRIN (the KARlsruhe TRitium Neutrino experiment)





# KATRIN (the KARlsruhe TRitium Neutrino experiment)

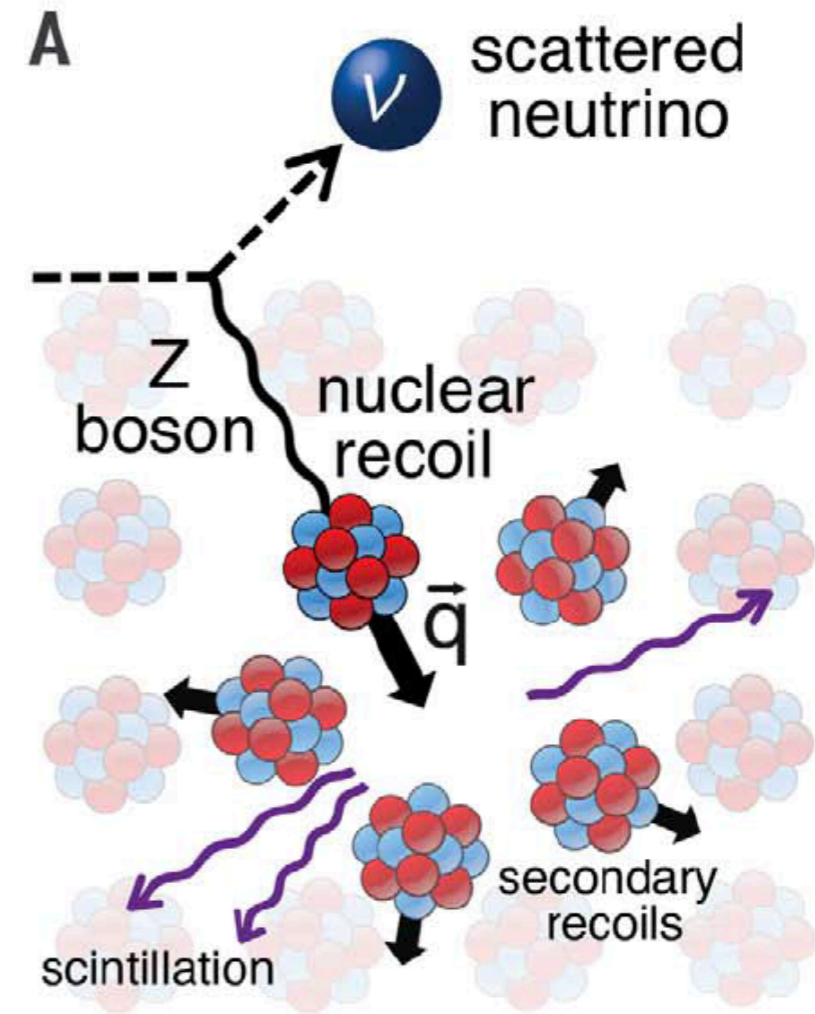




# NOT ALL ABOUT THE SIZE



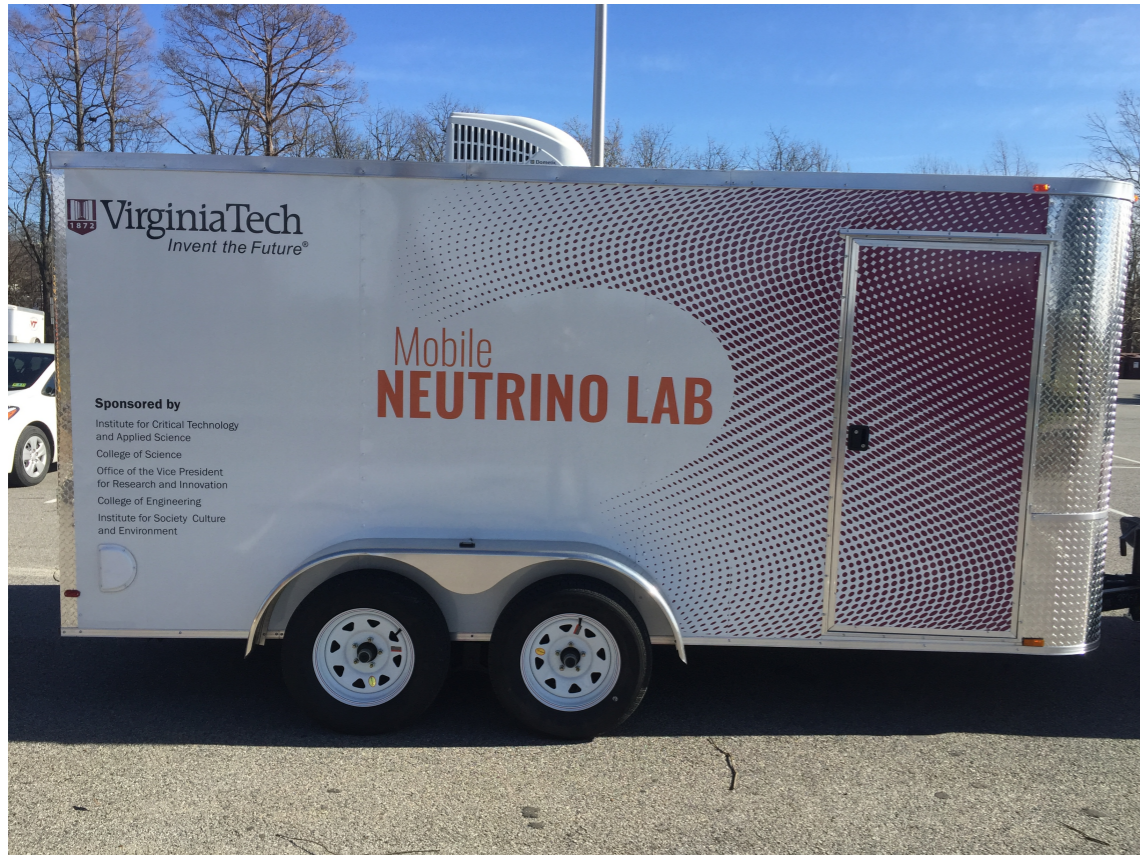
# COHERENT



First observation of coherent elastic neutrino-nucleus scattering (2017):  
Large cross-section  $\rightarrow$  small detector size



# PORTABLE NEUTRINO DETECTORS?



Why not drive to the nuclear reactors?



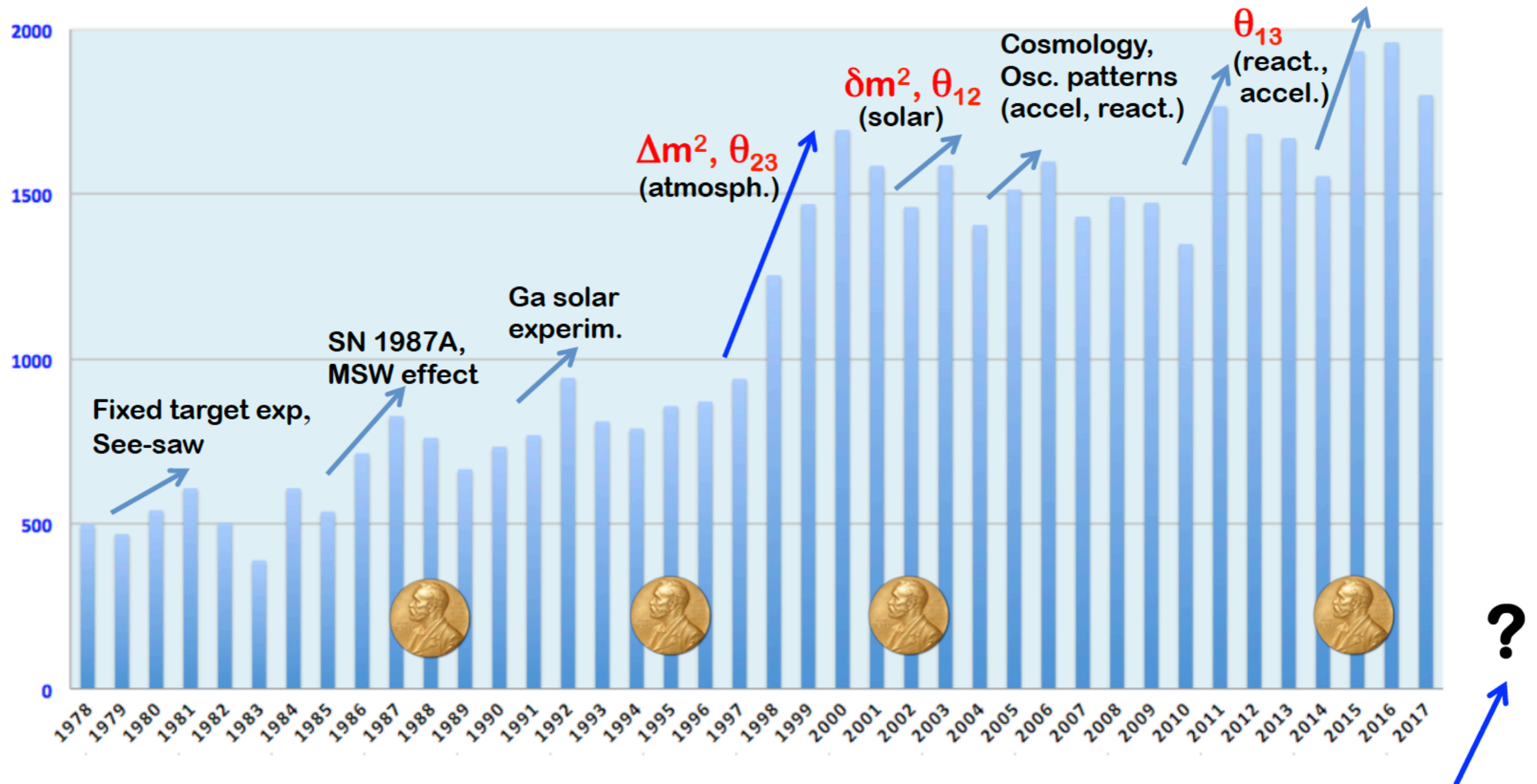
MiniCHANDLER - portable neutrino lab



# STATE OF NEUTRINO PHYSICS FIELD

N. of #neutrino# preprints per year (1978-2018) from 

**Pattern reflects breakthroughs and peaks of interest...** UHE  $\nu$ , CPV tests



**... with more to come!**

Credit: Eligio Lisi doi:10.5281/zenodo.1286744



# SUMMARY

- **Huge progress in the last decades**
- **Still wide range of open questions**
- **Wide range of experiments:**
  - **Long-baseline beam experiments -> oscillations parameters**
  - **Spectrometers and cosmology -> the absolute mass measurement**
  - **Clean experiments -> solar neutrinos and neutrino less double beta decay**
  - **Exciting future full of experiments**

# THANKS FOR YOUR ATTENTION



# QUESTIONS ARE WELCOME